

Localising the H.E.S.S. Galactic Centre point source

Isabel Braun

O. Bolz, C. van Eldik, G. Hermann, W. Hofmann

Max-Planck-Institut für Kernphysik, Heidelberg, Germany

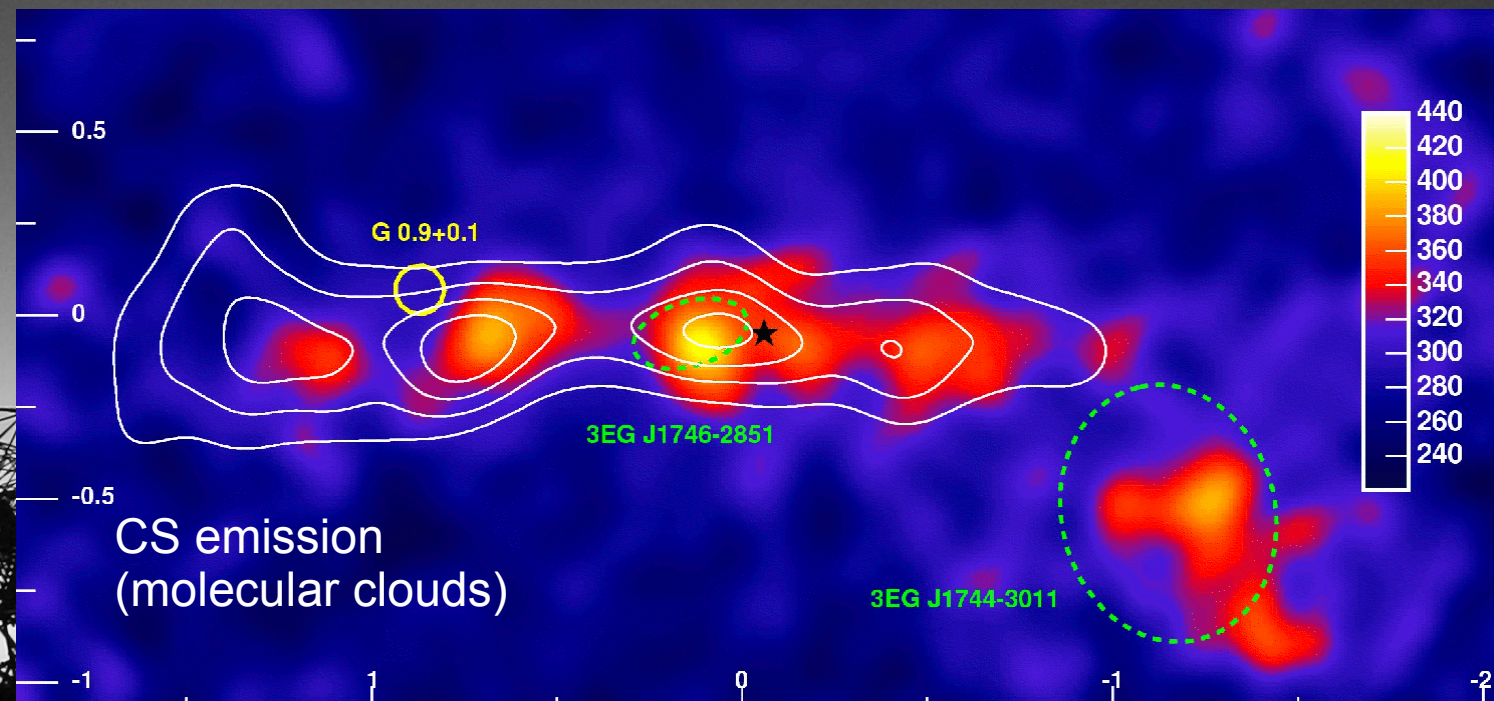
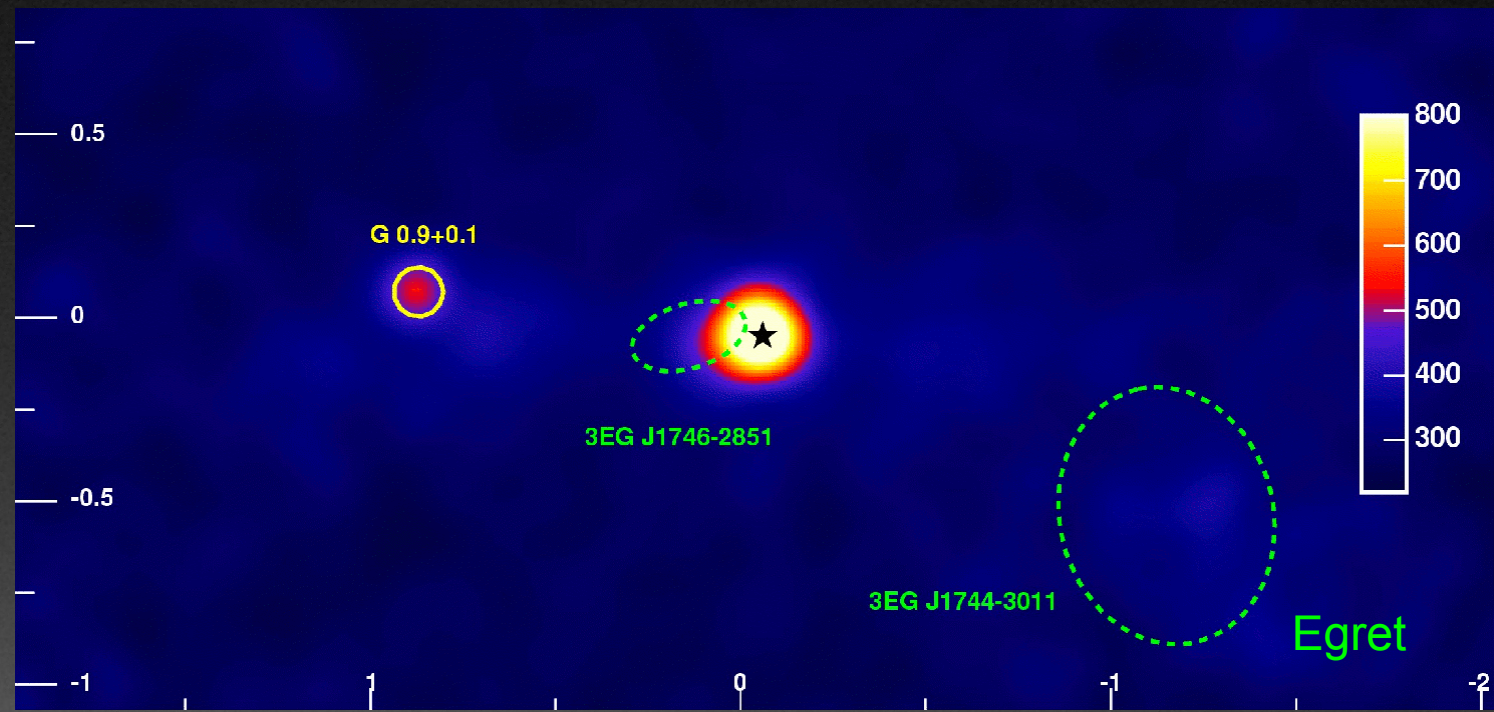
J. Hinton

School of Physics and Astronomy, University of Leeds, UK

for the H.E.S.S. collaboration



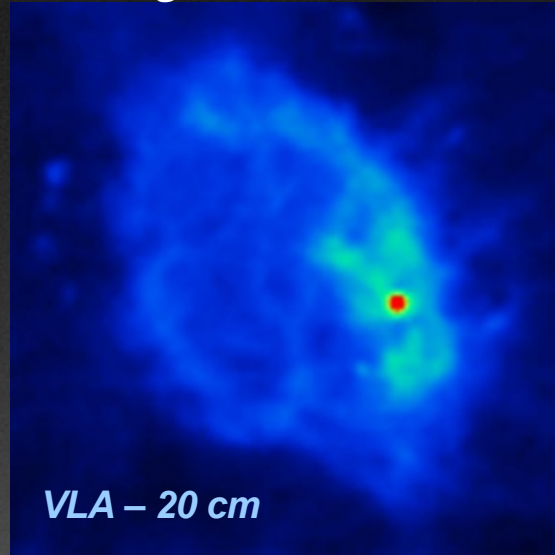
HESS J1745-290



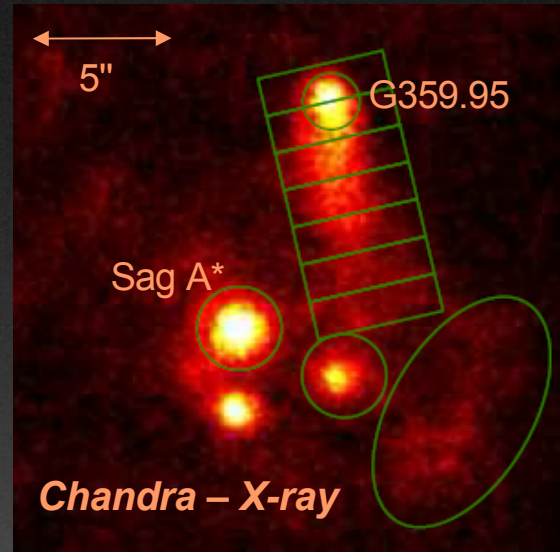
Discovery of very-high-energy gamma-rays from the Galactic Centre ridge,
H.E.S.S. collaboration, F. Aharonian et al., Nature 439 (2006) 695-698

possible counterparts

SNR Sgr A East



PWN G359.95-0.04

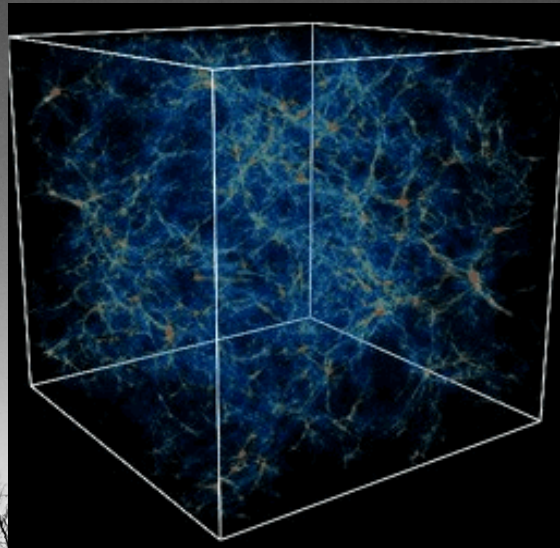


?

SMBH Sgr A*

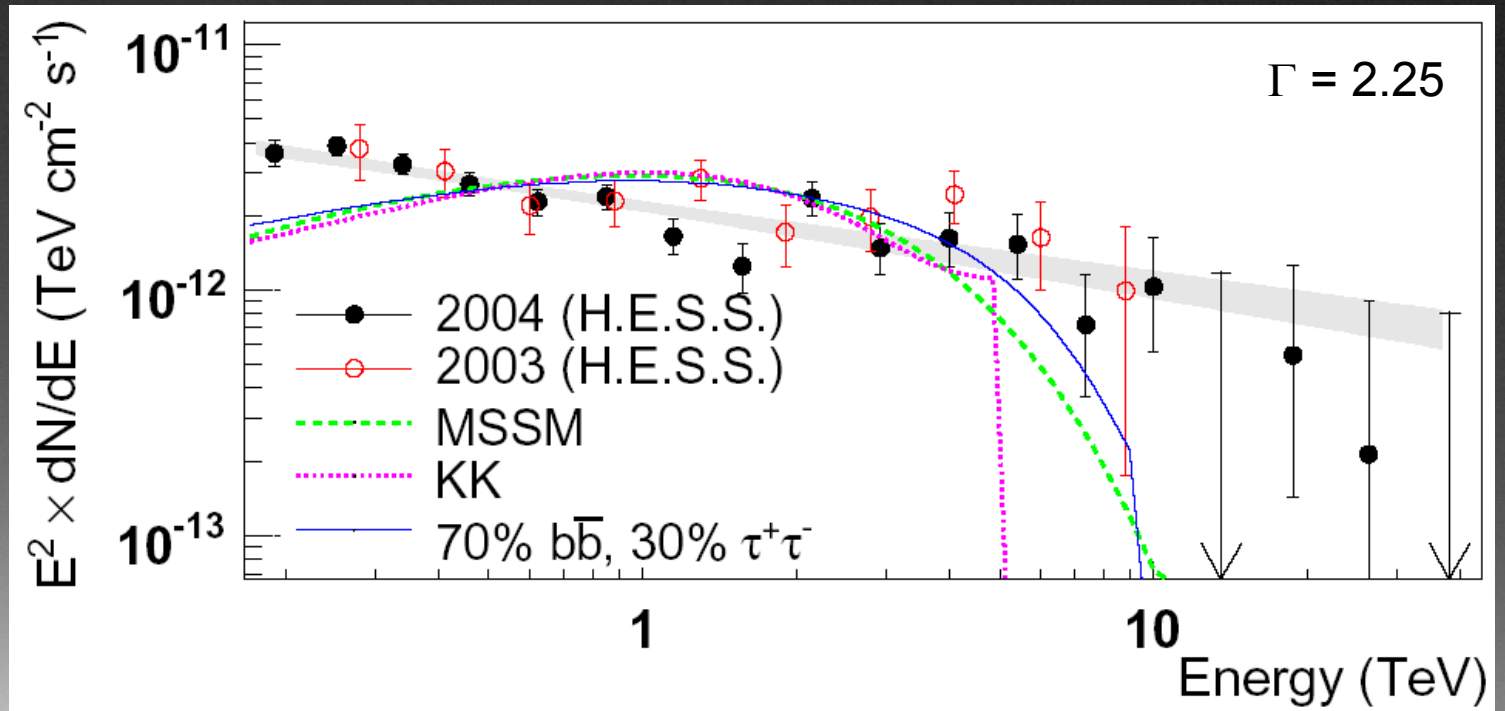


Dark Matter

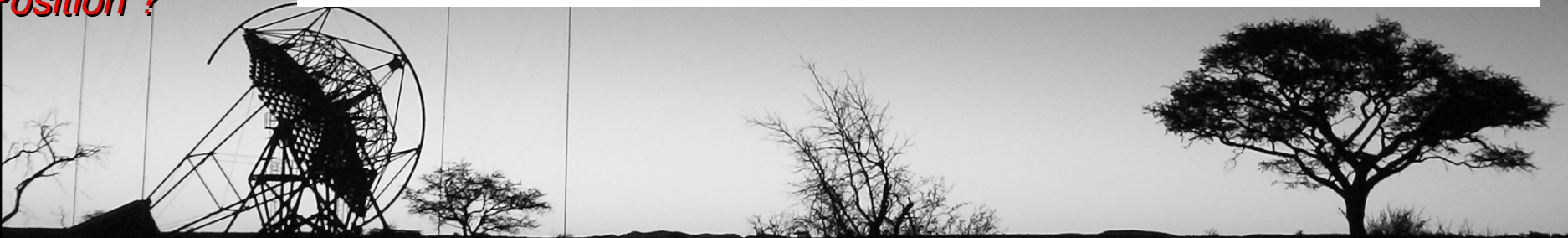


possible identification tools

- Extension *Point-like for H.E.S.S.*
- Variability *No variability on timescales of years to minutes detected*
- Spectrum



- *Position ?*

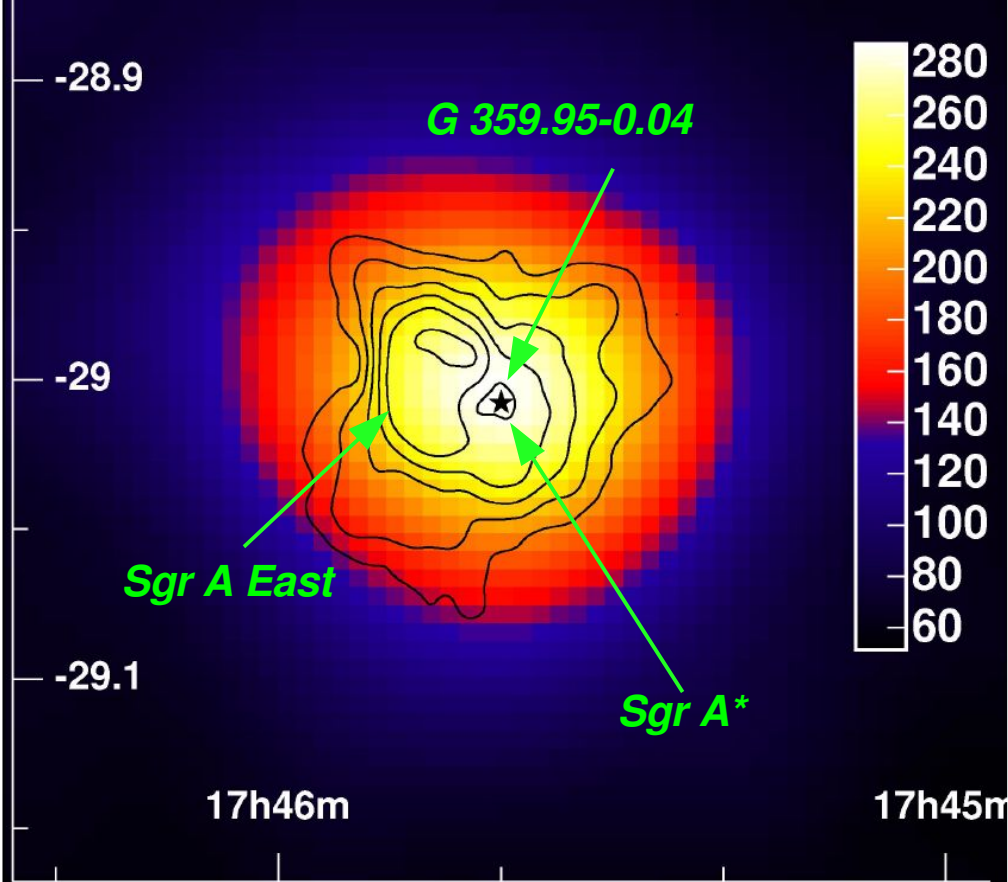


H.E.S.S. observations of the Galactic Center region and their possible dark matter interpretation,

H.E.S.S. collaboration, F. Aharonian et al., Phys. Rev. Lett. 97 (2006) 221102

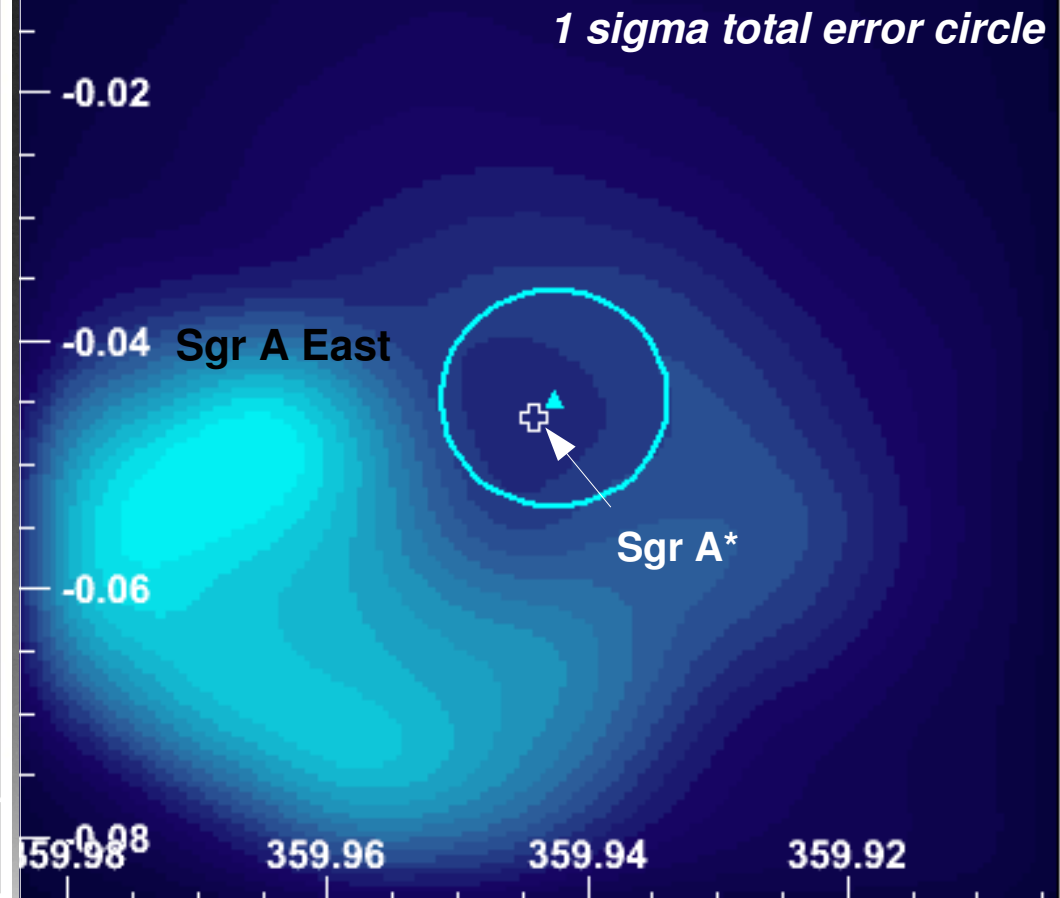
Position of HESS J1745-290

H.E.S.S. 2004



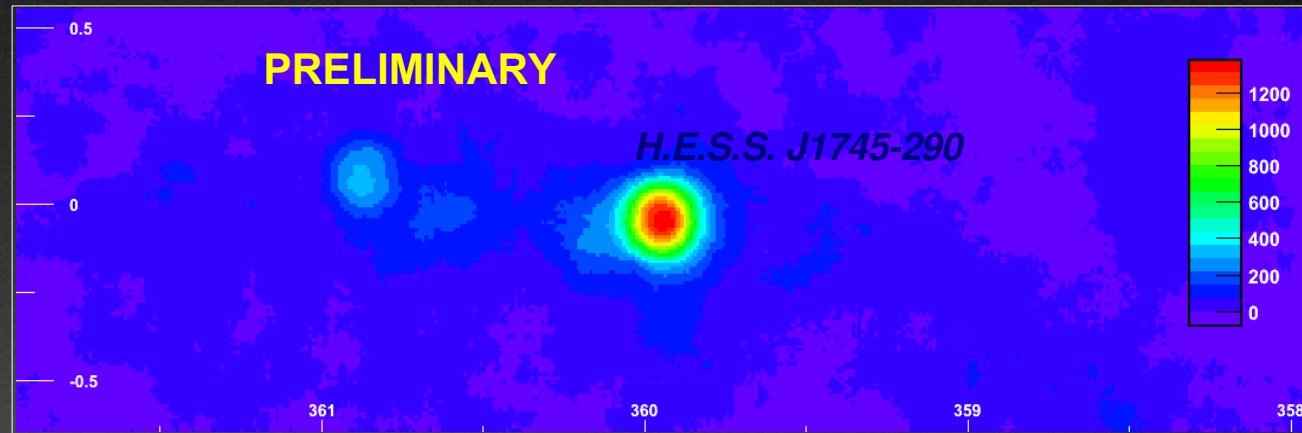
VLA 90cm

HESS J1745-290 (2004)
1 sigma total error circle



Pointing Improvements

- *New data 2005/2006:*
 - 73.2 h live time*
 - 44 σ significance*
 - 0.07° ang.res./event*



- *Improved Pointing Calibration :*
 - new method using guiding telescopes*
 - systematic error reduced from 28" to 9"!*



Position using 2005/2006 data

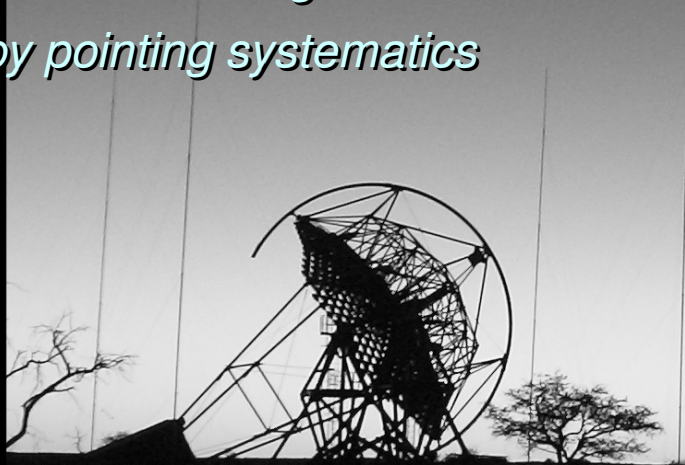
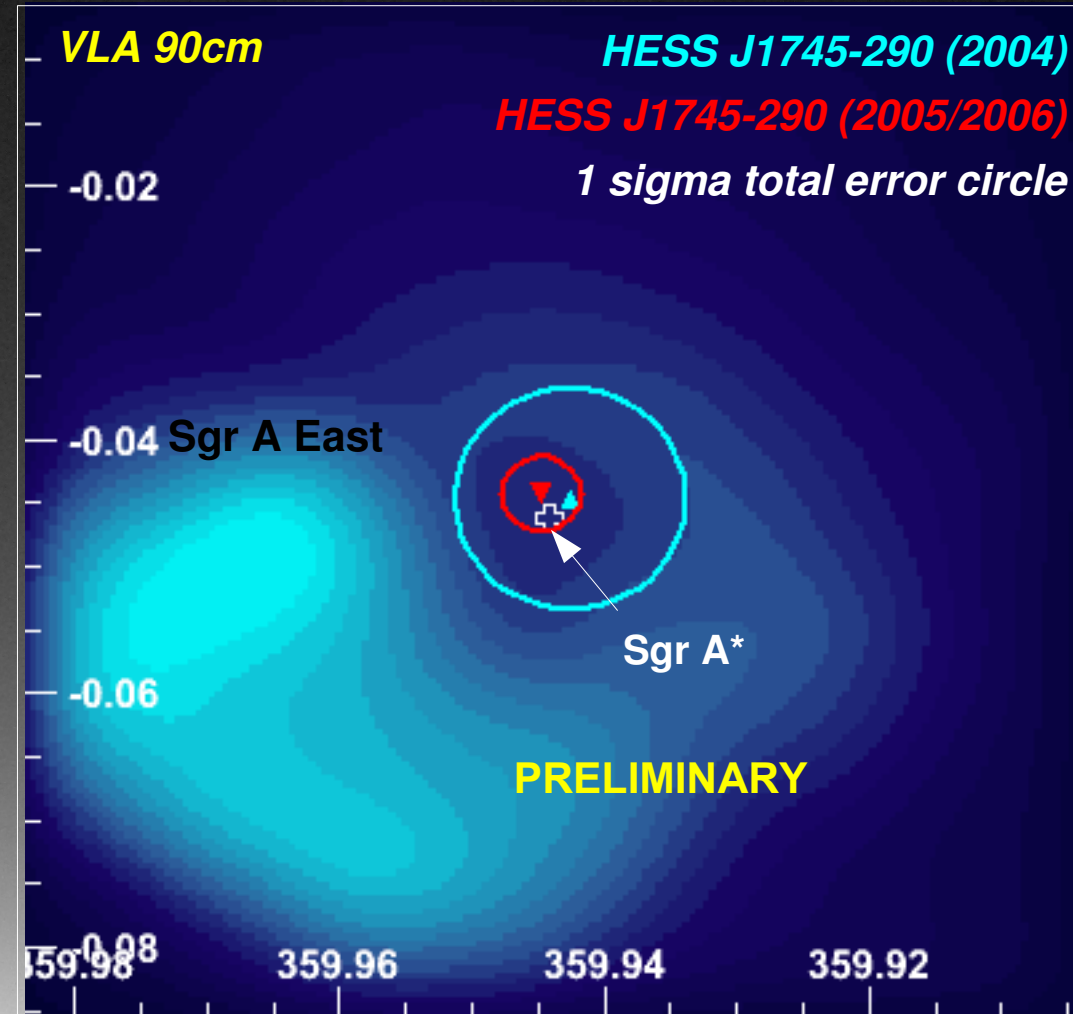
- best fit position:

$$l = 359^{\text{d}}56'41.1'' \pm 6.4'' \pm 6''$$

$$b = -0^{\text{d}}2'39.2'' \pm 5.9'' \pm 6''$$

preliminary statistical + systematic errors

- point-like after subtraction of diffuse emission
- position no longer dominated by pointing systematics



Position using 2005/2006 data

- best fit position:

$$l = 359^{\text{d}}56'41.1'' \pm 6.4'' \pm 6''$$

$$b = -0^{\text{d}}2'39.2'' \pm 5.9'' \pm 6''$$

preliminary statistical + systematic errors

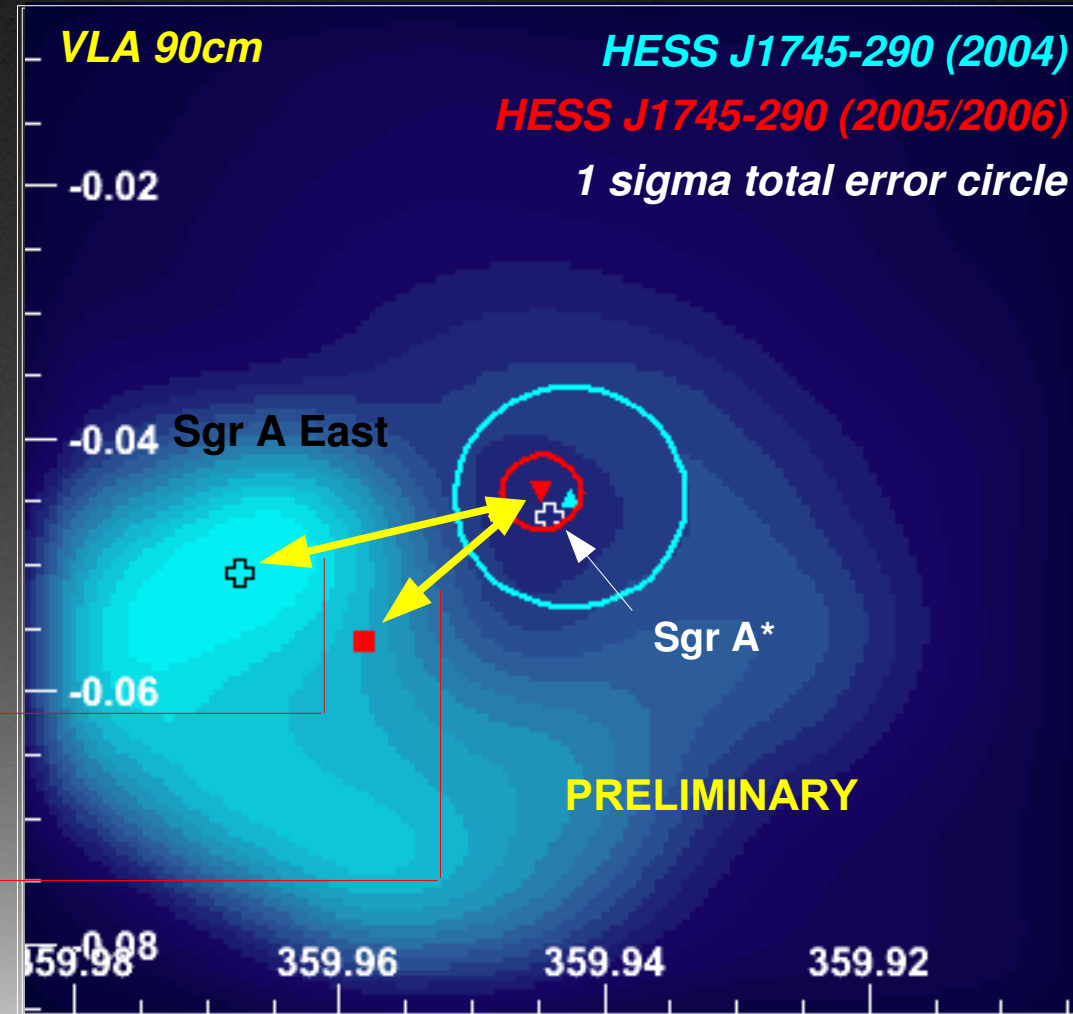
Assumptions

- VHE from Sgr A East radio maximum

-> excluded at 7σ level

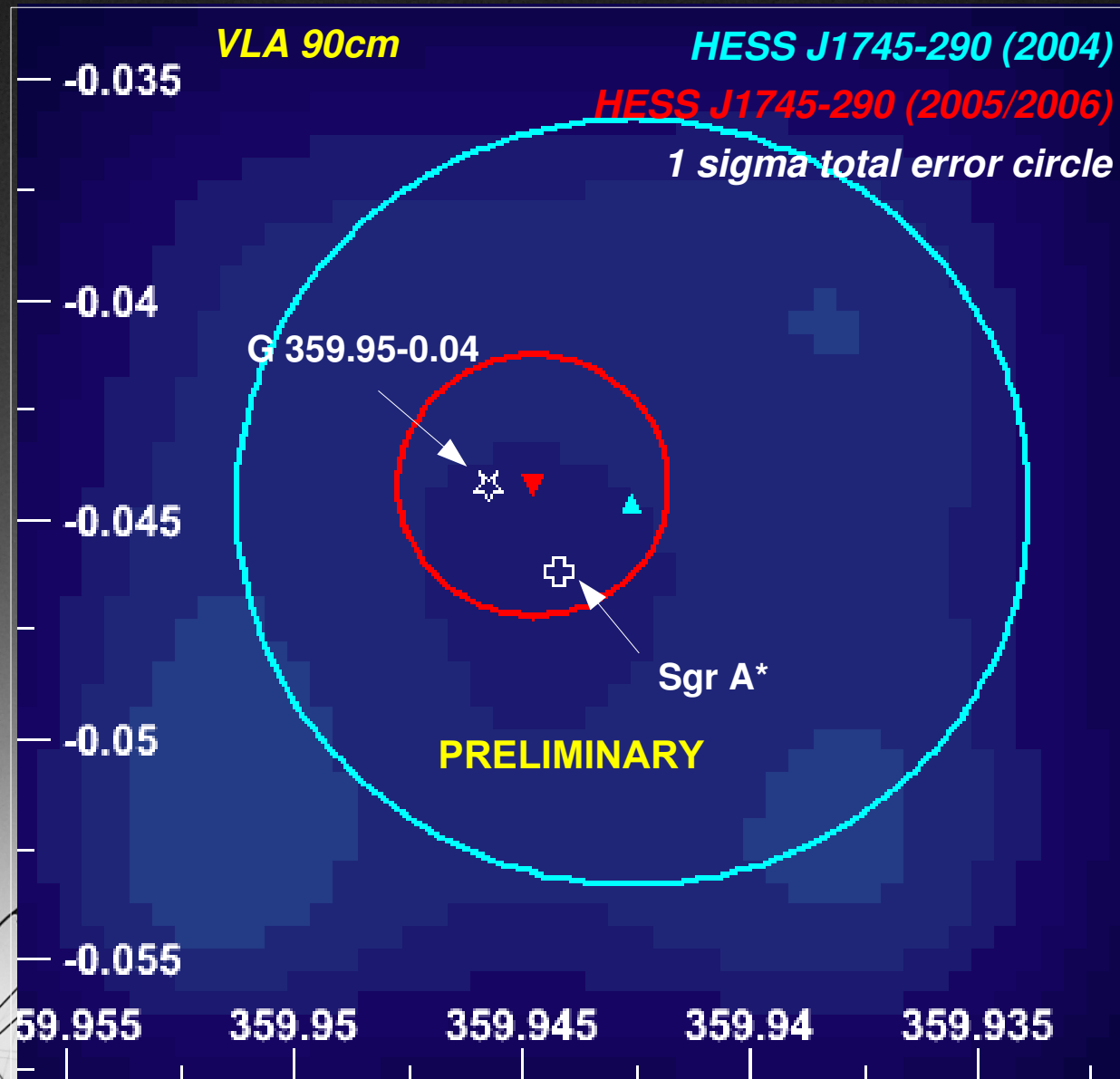
- VHE follows radio emission

-> excluded at 5.3σ level



-> bulk of TeV Gamma emission not from SNR Sgr A East!

Position using 2005/2006 data



-> good agreement for black hole and PWN

Summary

- *Competing candidates for TeV γ emission from the galactic centre*
- *Sgr A East can be excluded due to significant improvement of pointing accuracy*
- *SMBH Sgr A* and PWN G395.95-0.04 not resolvable*