

# An update on charming penguins in charmless B decays



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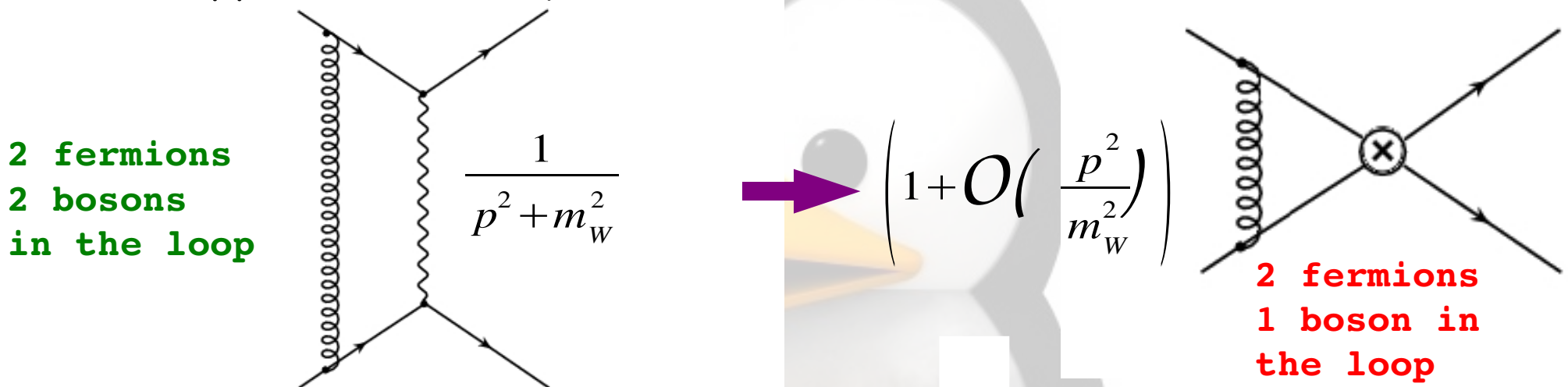
paper in  
preparation

# Current Experimental Situation

- The establishment of  $CP$  violation in  $K^+\pi^-$  and the observed discrepancy with respect to  $K^+\pi^0$  called for a “puzzle” in  $K\pi$  decays. For us, this is just the indication that charming and  $GIM$  penguins are not the end of the story
- the enhancement with time of  $BR(K^0\pi^0)$  introduced some tension in the commonly accepted models, interpreted as a possible hint of  $SU(2)$  breaking
- The large set of measurements (including  $S$  and  $C$ ) from BaBar and Belle allows to study not only  $K\pi$  but also  $PV$  modes

# The OPE and decay amplitudes

Since  $m_b \sim 4\text{GeV}$  and  $m_W \sim 80\text{GeV}$ , weak interaction can be replaced by an effective local theory, contracting the  $W$  propagator to a point (similar approach with  $t$  quark)



This operation breaks the ultraviolet behavior of the theory.

$$\int \frac{d^4 p}{p^6} \approx \int \frac{dp}{p^3} \rightarrow 0 \quad \longrightarrow \quad \int \frac{d^4 p}{p^4} \approx \int \frac{dp}{p} \rightarrow \infty$$

To remove the  $\infty$  after integrating out the heavy degrees of freedom we need to renormalize the theory. New operators couplings are generated

# The effective Hamiltonian

After the renormalization of the effective theory we get

$$\mathcal{H}_{\text{eff}} = \frac{G_F}{\sqrt{2}} \sum_{p=u,c} \lambda_p \left( C_1 Q_1^p + C_2 Q_2^p + \sum_{i=3,\dots,10} C_i Q_i + C_{11} Q_{12} + C_{12} Q_{12} \right) + \text{h.c.}$$

$$Q_1^p = (\bar{p}b)_{V-A} (\bar{s}p)_{V-A},$$

$$Q_2^p = (\bar{p}_i b_j)_{V-A} (\bar{s}_j p_i)_{V-A},$$

**Tree level  
operators**

$$Q_3 = (\bar{s}b)_{V-A} \sum_q (\bar{q}q)_{V-A},$$

$$Q_4 = (\bar{s}_i b_j)_{V-A} \sum_q (\bar{q}_j q_i)_{V-A},$$

**Penguin  
operators**

$$Q_5 = (\bar{s}b)_{V-A} \sum_q (\bar{q}q)_{V+A},$$

$$Q_6 = (\bar{s}_i b_j)_{V-A} \sum_q (\bar{q}_j q_i)_{V+A},$$

$$Q_7 = (\bar{s}b)_{V-A} \sum_q \frac{3}{2} e_q (\bar{q}q)_{V+A},$$

$$Q_8 = (\bar{s}_i b_j)_{V-A} \sum_q \frac{3}{2} e_q (\bar{q}_j q_i)_{V+A},$$

**EW Penguin  
operators**

$$Q_9 = (\bar{s}b)_{V-A} \sum_q \frac{3}{2} e_q (\bar{q}q)_{V-A},$$

$$Q_{10} = (\bar{s}_i b_j)_{V-A} \sum_q \frac{3}{2} e_q (\bar{q}_j q_i)_{V-A},$$

$$Q_{7\gamma} = \frac{-e}{8\pi^2} m_b \bar{s} \sigma_{\mu\nu} (1 + \gamma_5) F^{\mu\nu} b,$$

$$Q_{8g} = \frac{-g_s}{8\pi^2} m_b \bar{s} \sigma_{\mu\nu} (1 + \gamma_5) G^{\mu\nu} b,$$

**(cromo)magnetic operators**

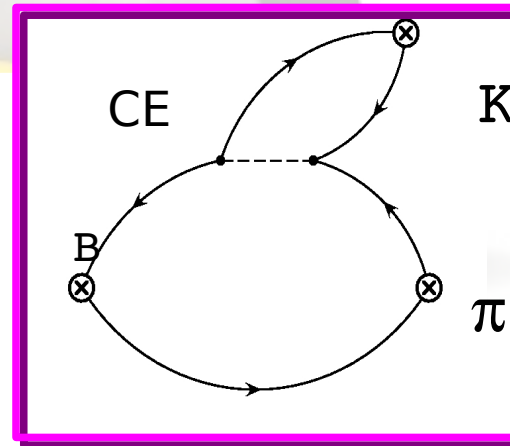
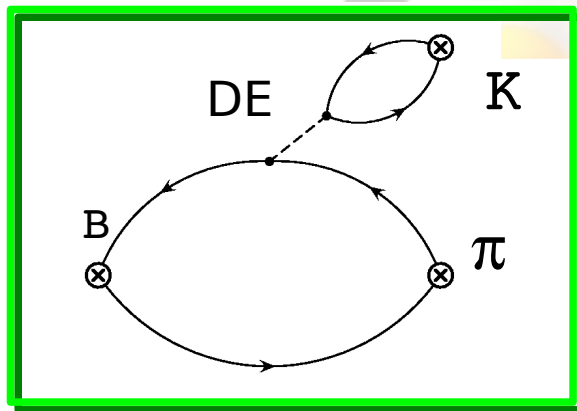
# Contractions of the $H_{eff}$

Contracting (Wick theorem)  $H_{eff}$  on initial and final states

$$A(B^0 \rightarrow K^+ \pi^-) = \langle K^+ \pi^- | H_{eff} | B^0 \rangle = \sum_{i=1,10} C_i(\mu) \langle K^+ \pi^- | Q_i(\mu) | B^0 \rangle$$

All the **perturbative physics** (scale  $> \mu$ ) in the **Wilson coeff.  $C_i(\mu)$** . All the **non-perturbative physics** (scale  $< \mu$ ) in the **matrix elements**. The unphysical dependence on  $\mu$  has to cancel out. One operator can produce several diagram topologies. Example: tree level operators generate

$\langle Q \rangle_{DE}(\mu)$  and  $\langle Q \rangle_{CE}(\mu)$



# The RGI combinations

One can rearrange contractions into **Renormalization Group Invariant** combinations, corresponding to physical quantities (**Buras & Silvestrini, hep-ph/9812392**). Example: T and  $T_c$  (trees) correspond to the RGI's  $E_1$  and  $E_2$

$$E_1 = C_1 \langle Q_1 \rangle_{DE} + C_2 \langle Q_2 \rangle_{CE}$$

$$E_2 = C_1 \langle Q_1 \rangle_{CE} + C_2 \langle Q_2 \rangle_{DE}$$

Penguins are more complicated

$$P_1 = C_1 \langle Q_1 \rangle_{CP}^c + C_2 \langle Q_2 \rangle_{DP}^c + \sum_{i=2}^5 \left( C_{2i-1} \langle Q_{2i-1} \rangle_{CE} + C_{2i} \langle Q_{2i} \rangle_{DE} \right) \\ + \sum_{i=3}^{10} \left( C_i \langle Q_i \rangle_{CP} + C_i \langle Q_i \rangle_{DP} \right) + \sum_{i=2}^5 \left( C_{2i-1} \langle Q_{2i-1} \rangle_{CA} + C_{2i} \langle Q_{2i} \rangle_{DA} \right)$$

$$P_1^{\text{GIM}} = C_1 \left( \langle Q_1 \rangle_{CP}^c - \langle Q_1 \rangle_{CP}^u \right) + C_2 \left( \langle Q_2 \rangle_{DP}^c - \langle Q_2 \rangle_{DP}^u \right)$$

Every RGI corresponds to a contraction of the  $J_\mu J^\mu$  interaction term of the Standard Model (i.e. RGIs are the physical quantities)

# The Decay Amplitude

The final formula is simplified and the dependence on  $\mu$  is formally canceled out

$$\begin{aligned}
 A(B^0 \rightarrow K^+ \pi^-) &= V_{ts} V_{tb}^* \times P_1 - V_{us} V_{ub}^* \times \{E_1 - P_1^{GIM}\} \\
 A(B^+ \rightarrow K^0 \pi^+) &= -V_{ts} V_{tb}^* \times P_1 + V_{us} V_{ub}^* \times \{A_1 - P_1^{GIM}\} \\
 \sqrt{2} \cdot A(B^+ \rightarrow K^+ \pi^0) &= V_{ts} V_{tb}^* \times P_1 - V_{us} V_{ub}^* \times \{E_1 + E_2 + A_1 - P_1^{GIM}\} \\
 \sqrt{2} \cdot A(B^0 \rightarrow K^0 \pi^0) &= -V_{ts} V_{tb}^* \times P_1 - V_{us} V_{ub}^* \times \{E_2 + P_1^{GIM}\}
 \end{aligned}$$

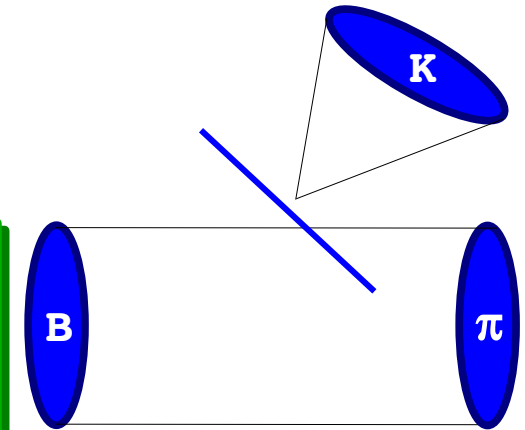
- ✚ We know  $C(\mu)$  from perturbative calculations
- ✚ We still miss a technique to calculate matrix elements to obtain the values of the RGI

# Perturbative Approaches

Perturbative calculations are possible in  $m_b \rightarrow \infty$  limit

- + pQCD by Keum, Li & Sanda [hep-ph/0004004](#)
- + QCD Factorization by BBNS [hep-ph/0006124](#)
- + SCET by BPRS [hep-ph/0401188](#)

$$\langle B^0 | J_\mu J^\mu | K \pi \rangle = \langle B^0 | J_\mu | \pi \rangle \langle 0 | J^\mu | K \rangle (1 + O(\alpha_s)) + O\left(\frac{\Lambda_{\text{QCD}}}{m_b}\right)$$

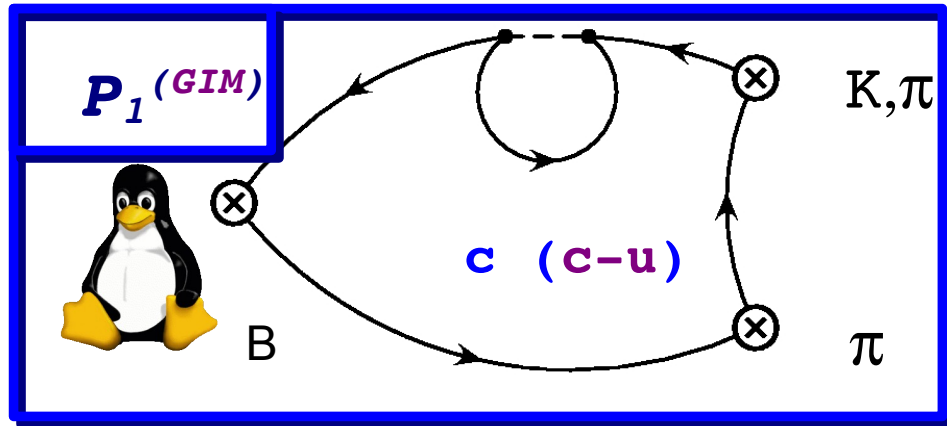


- + A **clear demonstration** that **penguins** do factorize is still **missing** (Bauer et al. [hep-ph/0401188](#) vs Beneke et al. [hep-ph/0411171](#))
- + As previously pointed out (Ciuchini et al. [hep-ph/9703353](#))  $\Lambda_{\text{QCD}}/m_b$  contributions may play a **relevant role in phenomenology** ( $m_b \not\rightarrow \infty$ )

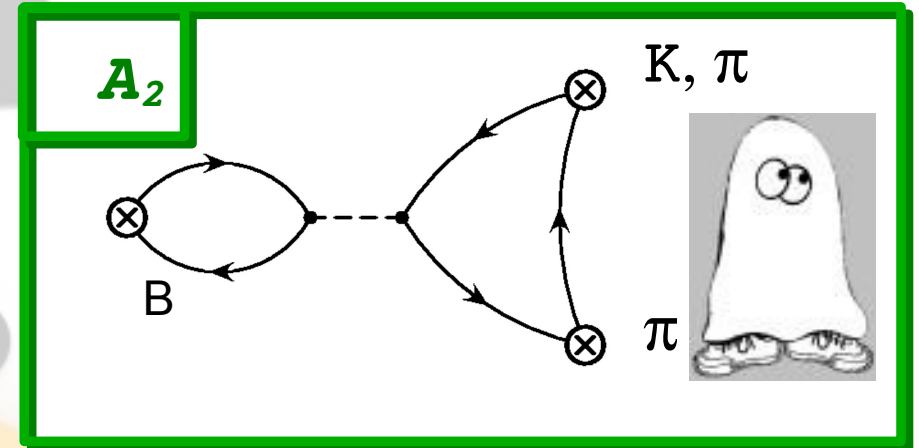


# $RGI \sim \Lambda_{\text{QCD}}/m_b$ in $b \rightarrow s$ decays

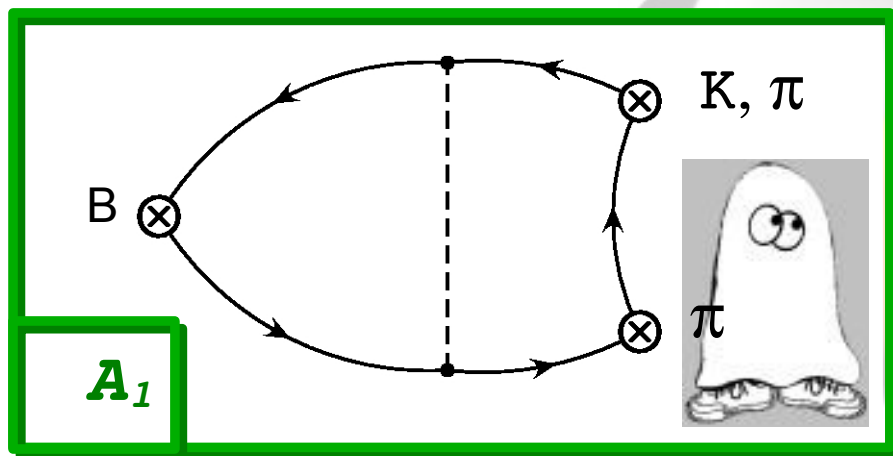
Charming and GIM penguins (c-u)



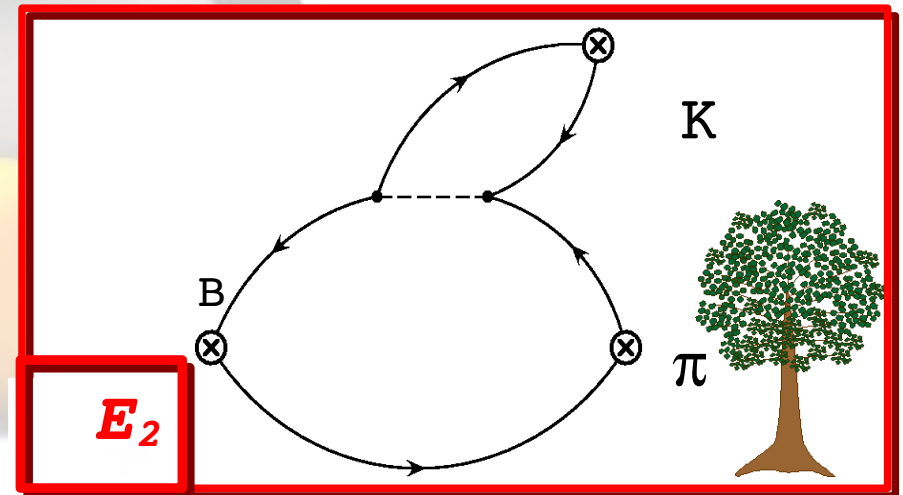
Disconnected Annihilation



Connected Annihilation



Connected emission



# A Few Comments

One can move from the exact calculation to some model dependent approach, but all  $\Lambda_{\text{QCD}}/m_b$  terms have to be considered

- +  $P_1$  is doubly Cabibbo enhanced, so it plays the major role
- + Nevertheless, the others are important

In principle: we have enough observables to determine all the parameters (4 complex RGI) and keep some predictive power

In practice: we are not precisely sensitive to the doubly Cabibbo suppressed  $\Lambda_{\text{QCD}}/m_b$  terms (which are  $\sim\%$  corrections to BR's)

What we can do:

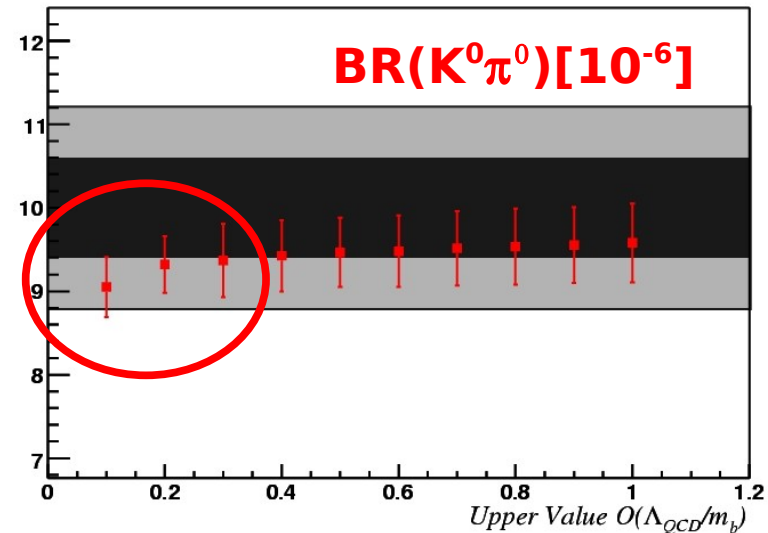
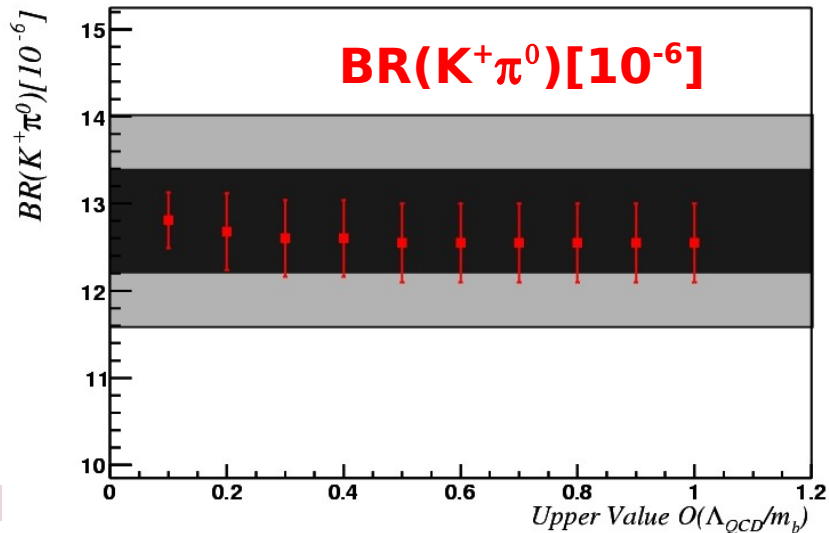
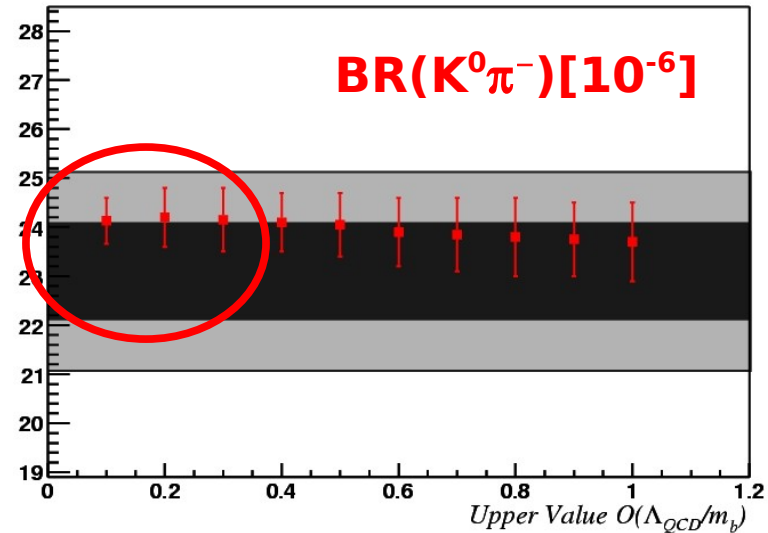
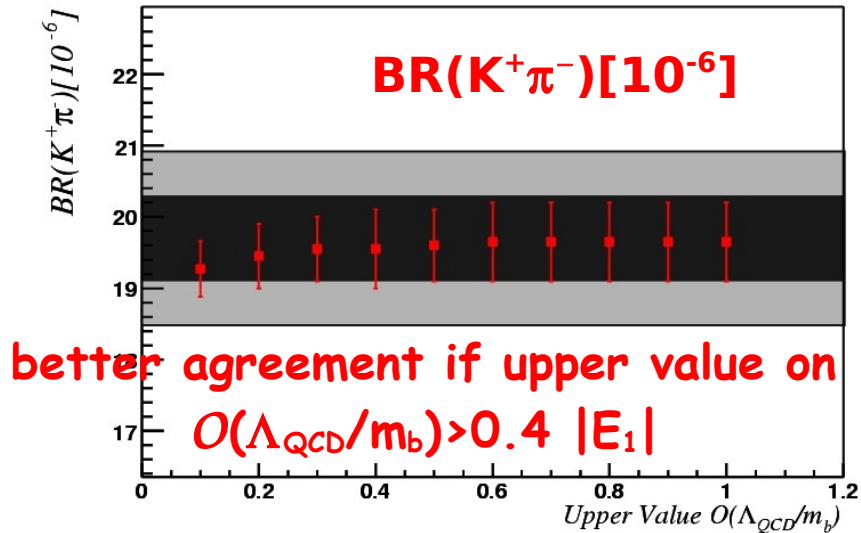
- + Describe  $E_1$  in factorization
- + Fit for the leading term  $P_1$
- + Vary the other in some *a-priori fixed range*
- + Use BR and direct CP asymmetries to obtain information on  $S(K^0\pi^0)$

We can still obtain a prediction on  $S$  within the Standard Model, but the error on it will depend on the chosen range. **So, we scan the upper bound on the range**

# Result on $B \rightarrow K\pi$ (I)

- $1\sigma$  exp range
- $2\sigma$  exp range

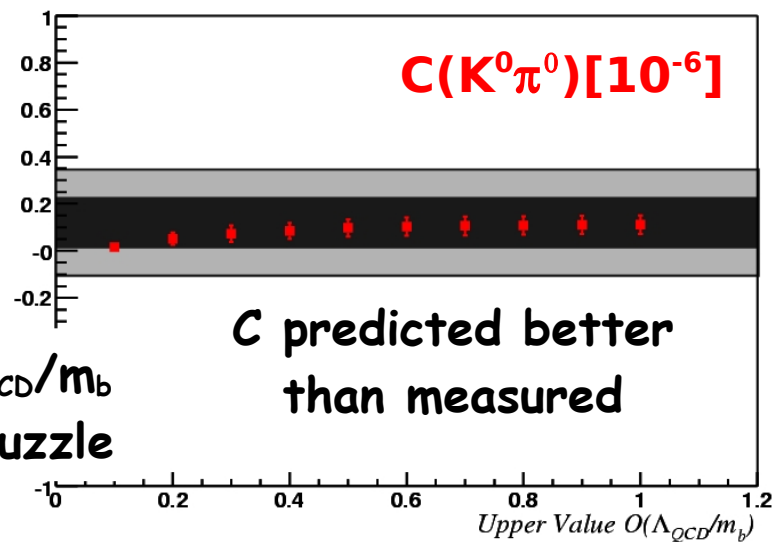
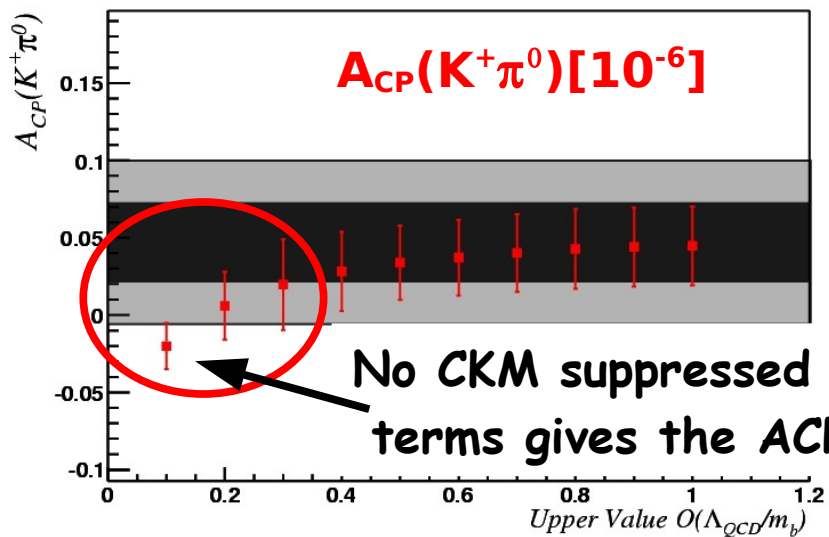
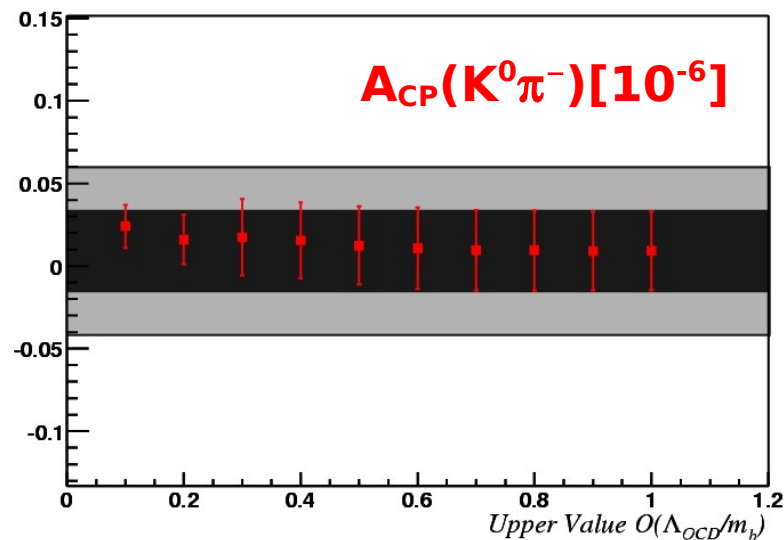
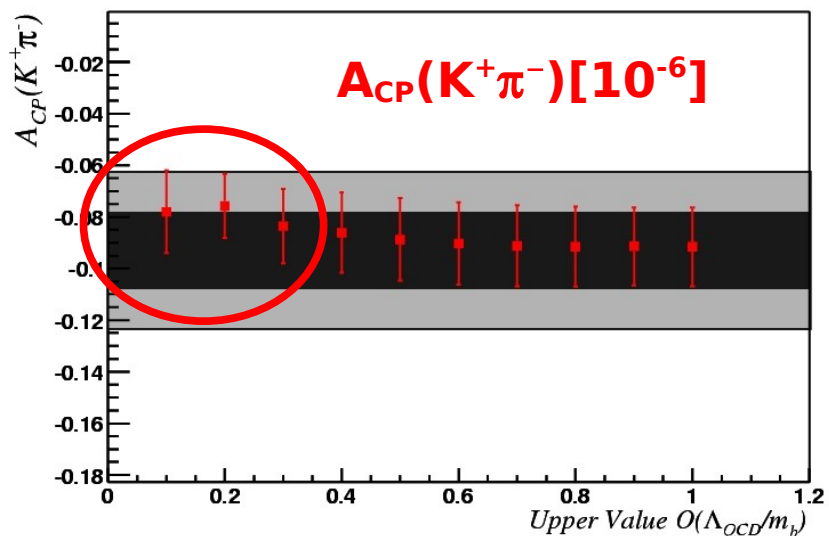
With all  $\Lambda_{\text{QCD}}/m_b$  corrections  
no BR  $K\pi$  puzzle!!!



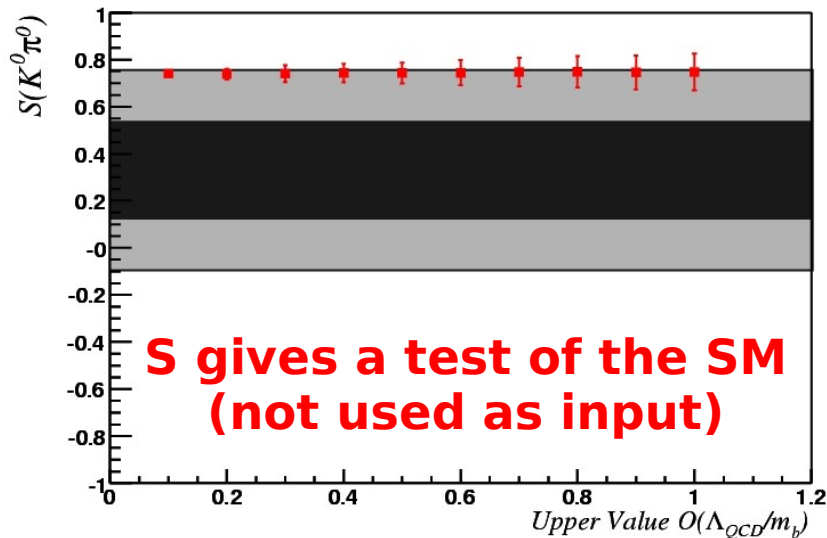
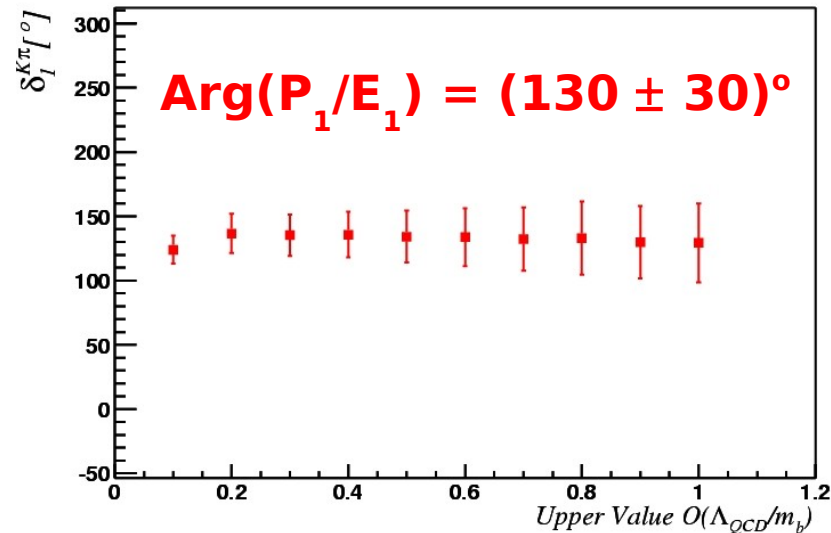
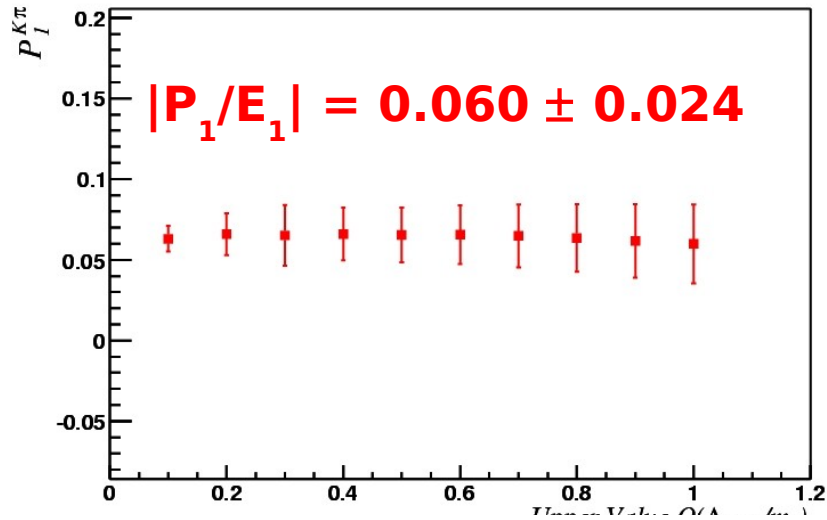
# Result on $B \rightarrow K\pi$ (II)

- $1\sigma$  exp range
- $2\sigma$  exp range

With all  $\Lambda_{\text{QCD}}/m_b$  corrections  
no  $A_{\text{CP}}$   $K\pi$  puzzle!!!



# Result on $B \rightarrow K\pi$ (III)



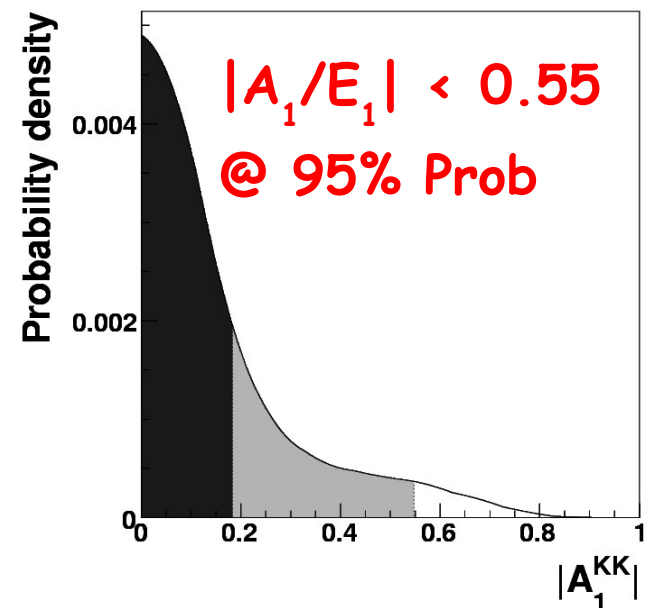
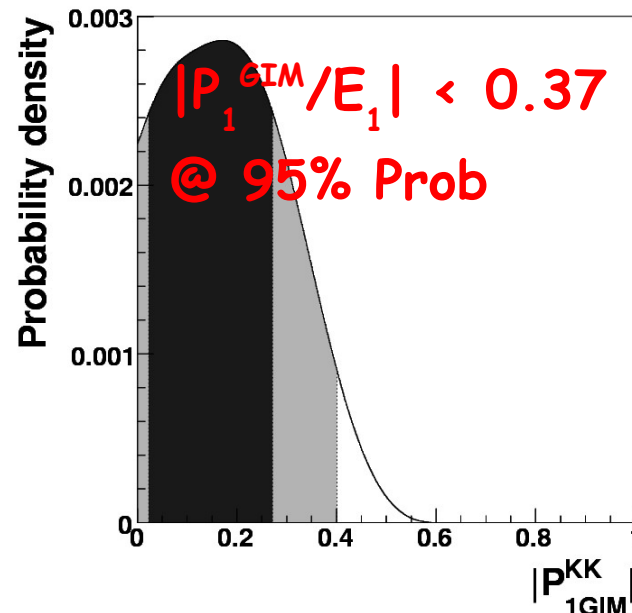
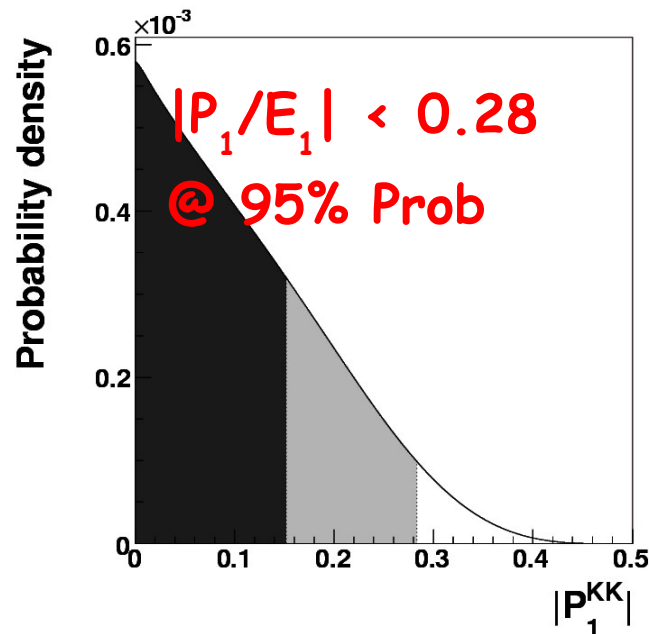
- ✚ The prediction on  $S(K^0\pi^0)$  is stable
- ✚ The error depends on the upper value of the range
- ✚ In a very conservative situation ( $O(\Lambda_{\text{QCD}}/m_b)/E_1 \in [0,1]$ ) we can still test the SM
- ✚ Limiting factor is still the exp. precision

# B → KK and the magnitude of $O(\Lambda_{\text{QCD}}/m_b)$

3 RGI (i.e. 5 real parameters to fit)  
 2BR, 2 direct CP asymmetry and S(K+K-)

$$\begin{aligned}
 A(B^+ \rightarrow K^+ K^0) &= -V_{td} V_{tb}^* \times P_1 + V_{ud} V_{ub}^* \times \{A_1 - P_1^{\text{GIM}}\} \\
 A(B^+ \rightarrow K^0 K^0) &= -V_{td} V_{tb}^* \times P_1 - V_{ud} V_{ub}^* \times \{P_1^{\text{GIM}}\}
 \end{aligned}$$

Large values of the parameters are suppressed.  
 Even with SU(3) broken @100% we do not expect large enhancements

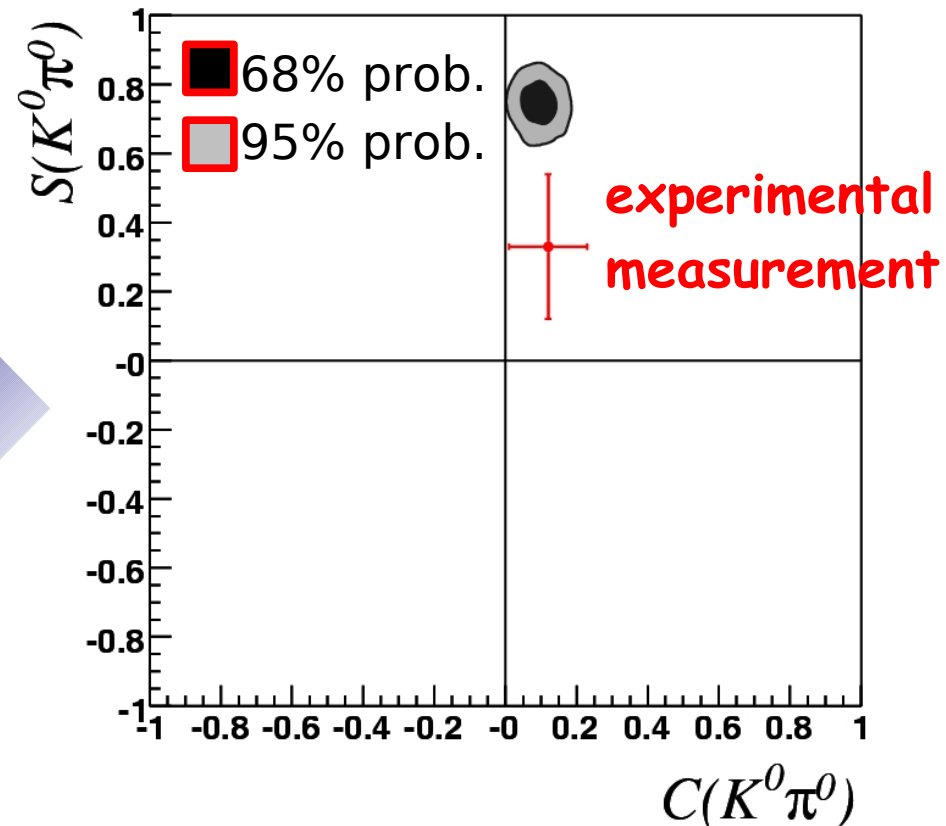
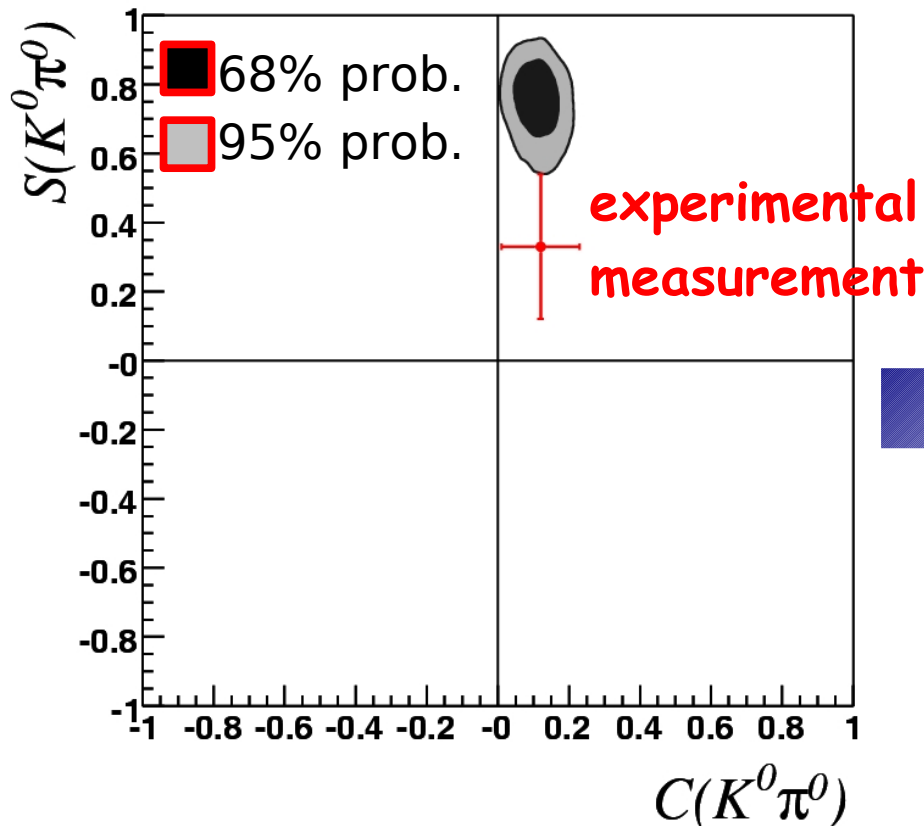


# Test of SM: $S_{K\pi}$ vs $C_{K\pi}$

Since  $C$  is better determined by the fit than by the experiment, we have information on it from the other variables + SU(2) relations (all possible sum rules you can imagine are implemented). We can remove also  $C$  from the set of inputs and look at the agreement in the  $S$  vs  $C$  plane

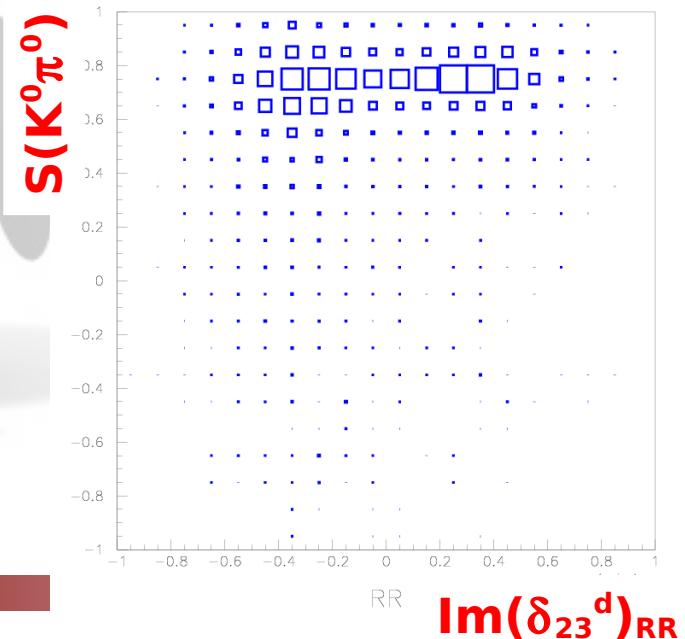
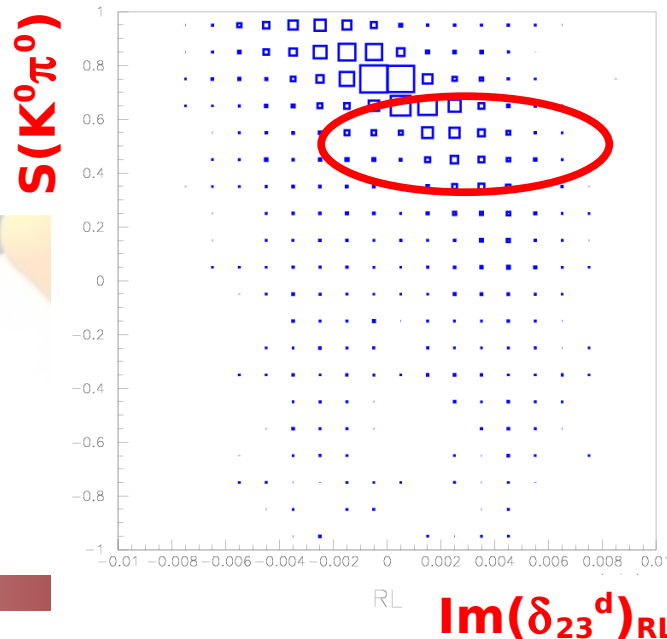
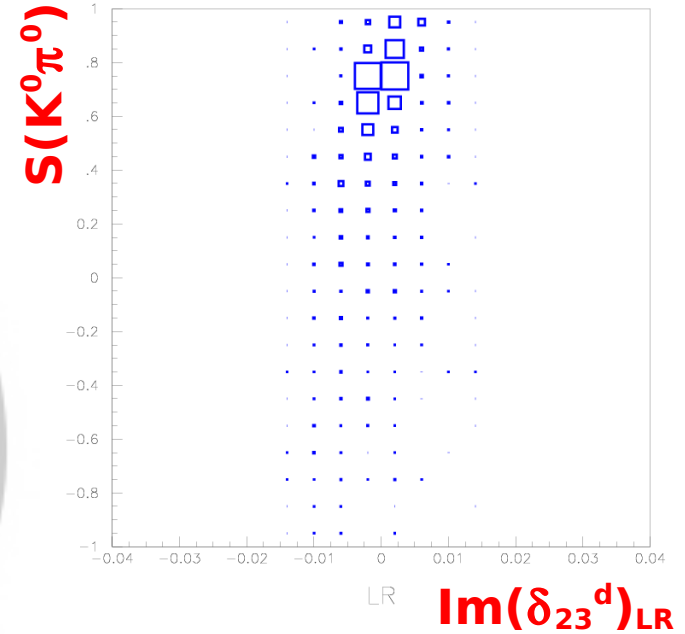
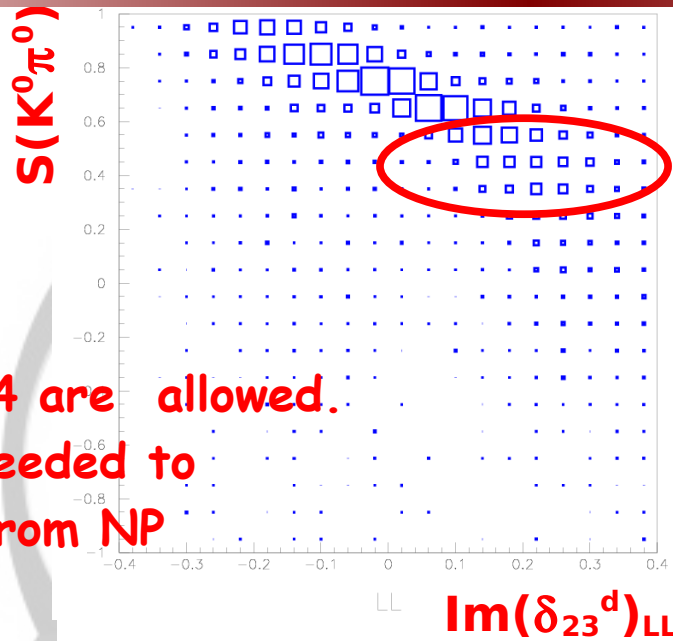
$O(\Lambda_{\text{QCD}}/m_b)/E_1 \in [0,1]$

$O(\Lambda_{\text{QCD}}/m_b)/E_1 \in [0,0.5]$



# What SUSY can do

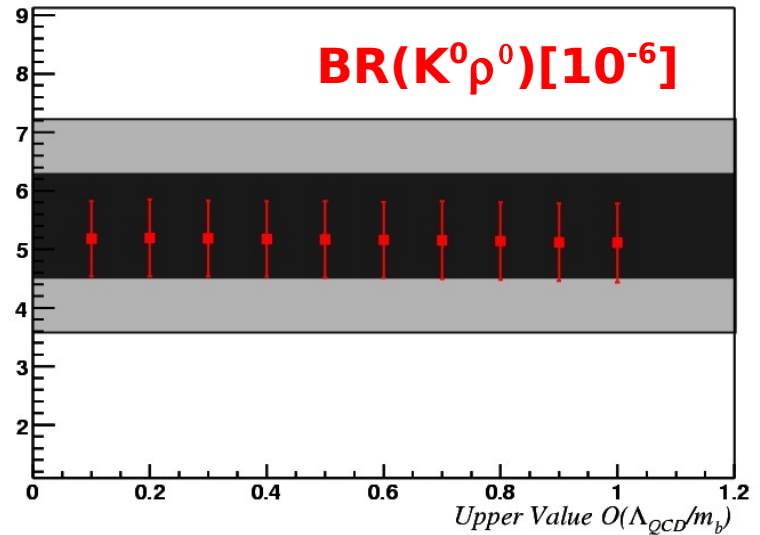
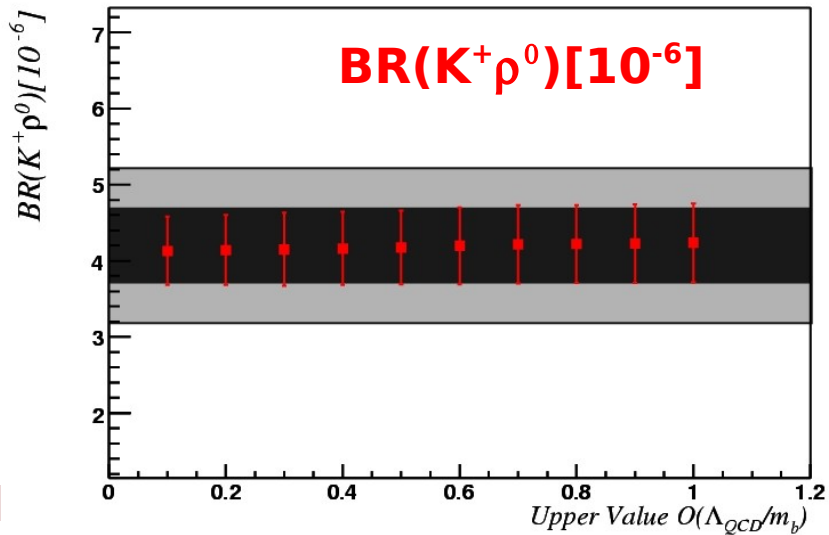
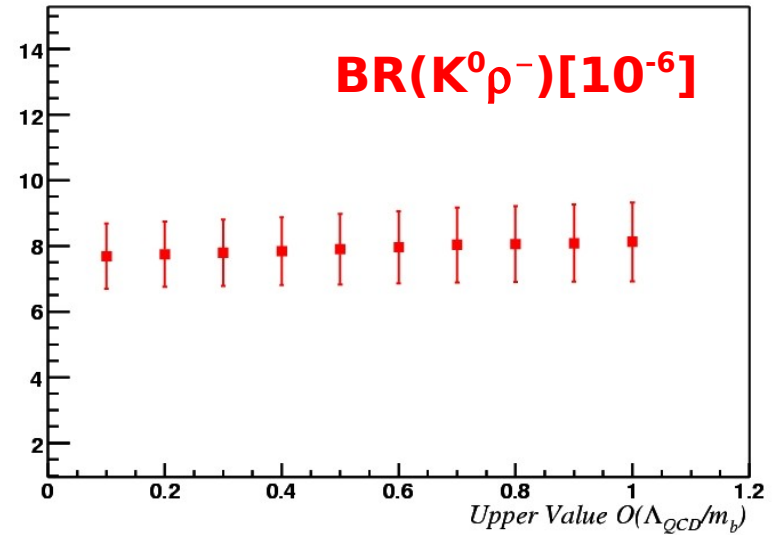
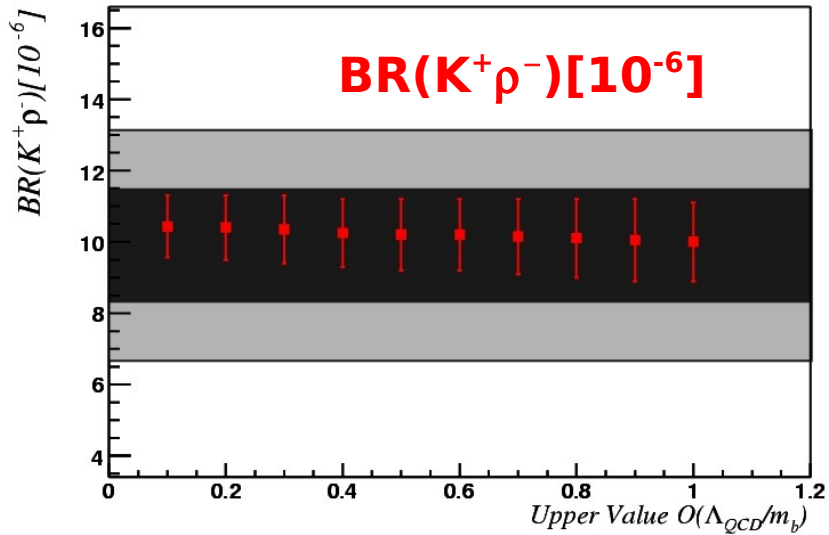
Typical values of 0.3-0.4 are allowed.  
SuperB statistics needed to disentangle SM from NP





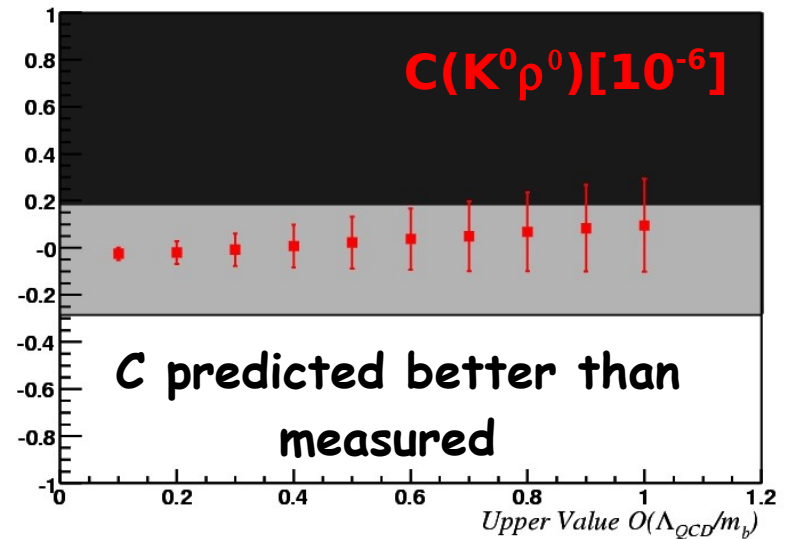
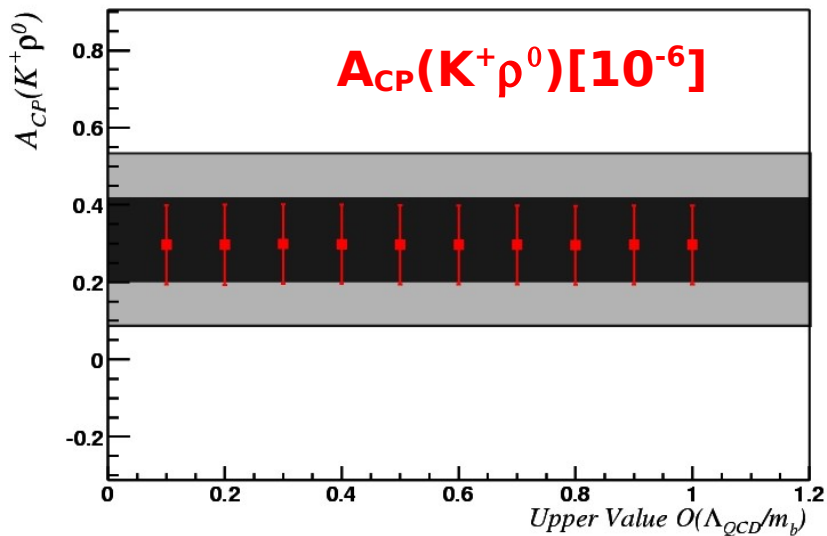
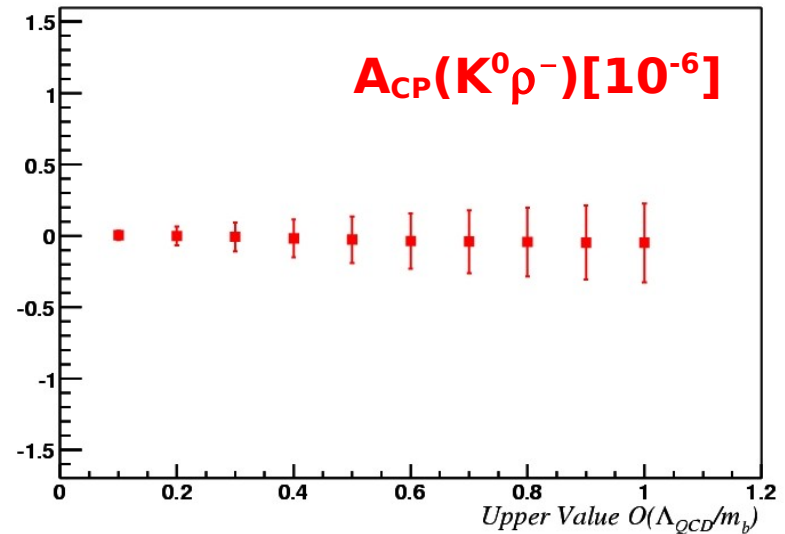
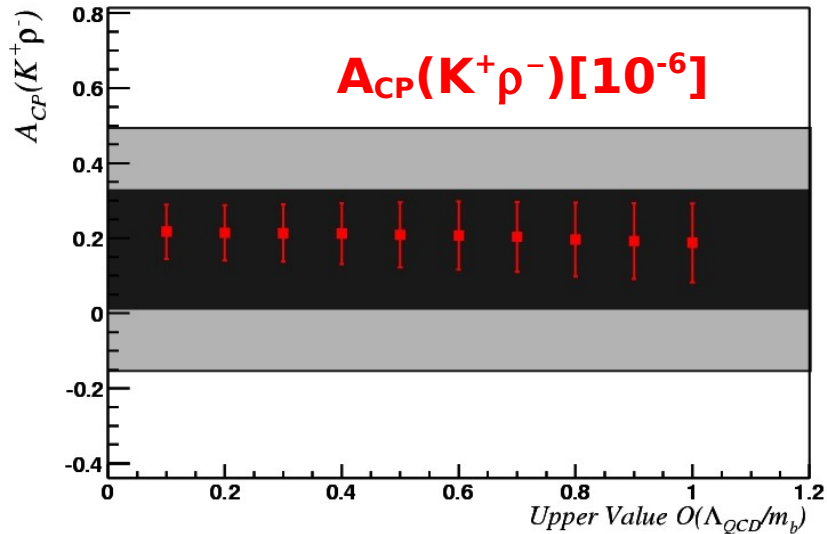
# Result on $B \rightarrow K\rho$ (I)

- $1\sigma$  exp range
- $2\sigma$  exp range

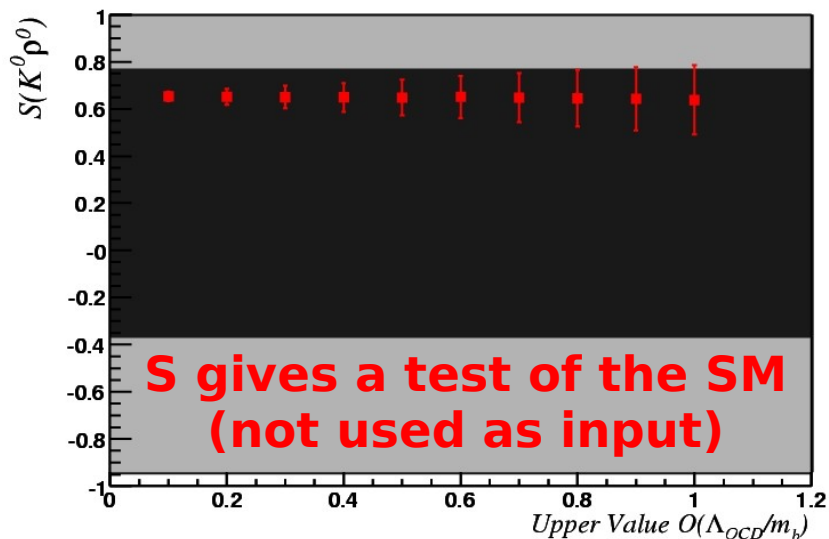
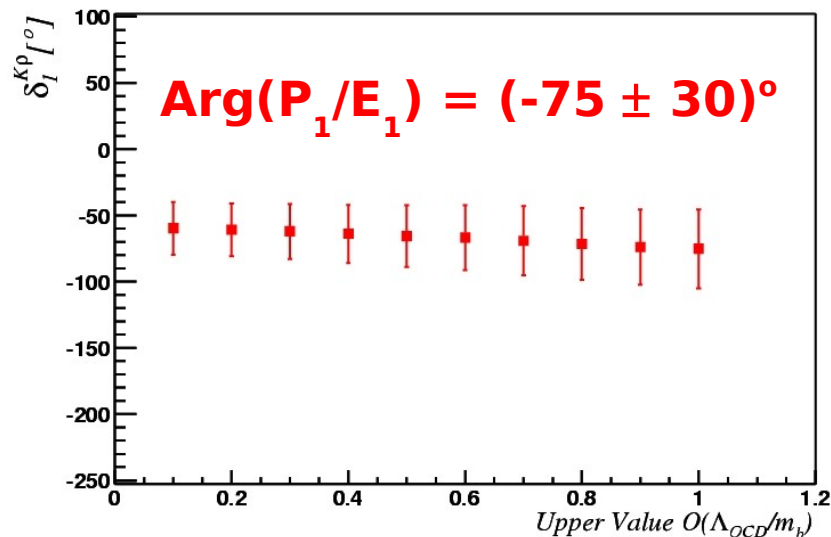
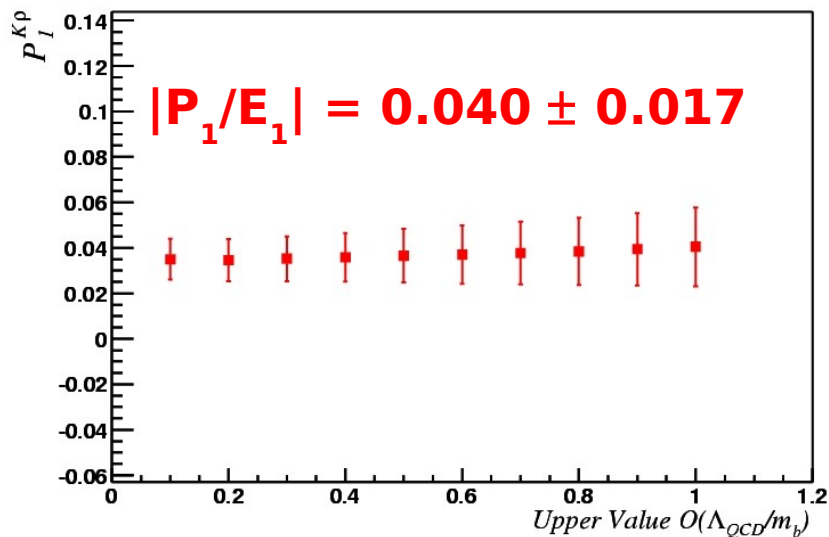


# Result on $B \rightarrow K\rho$ (II)

- $1\sigma$  exp range
- $2\sigma$  exp range



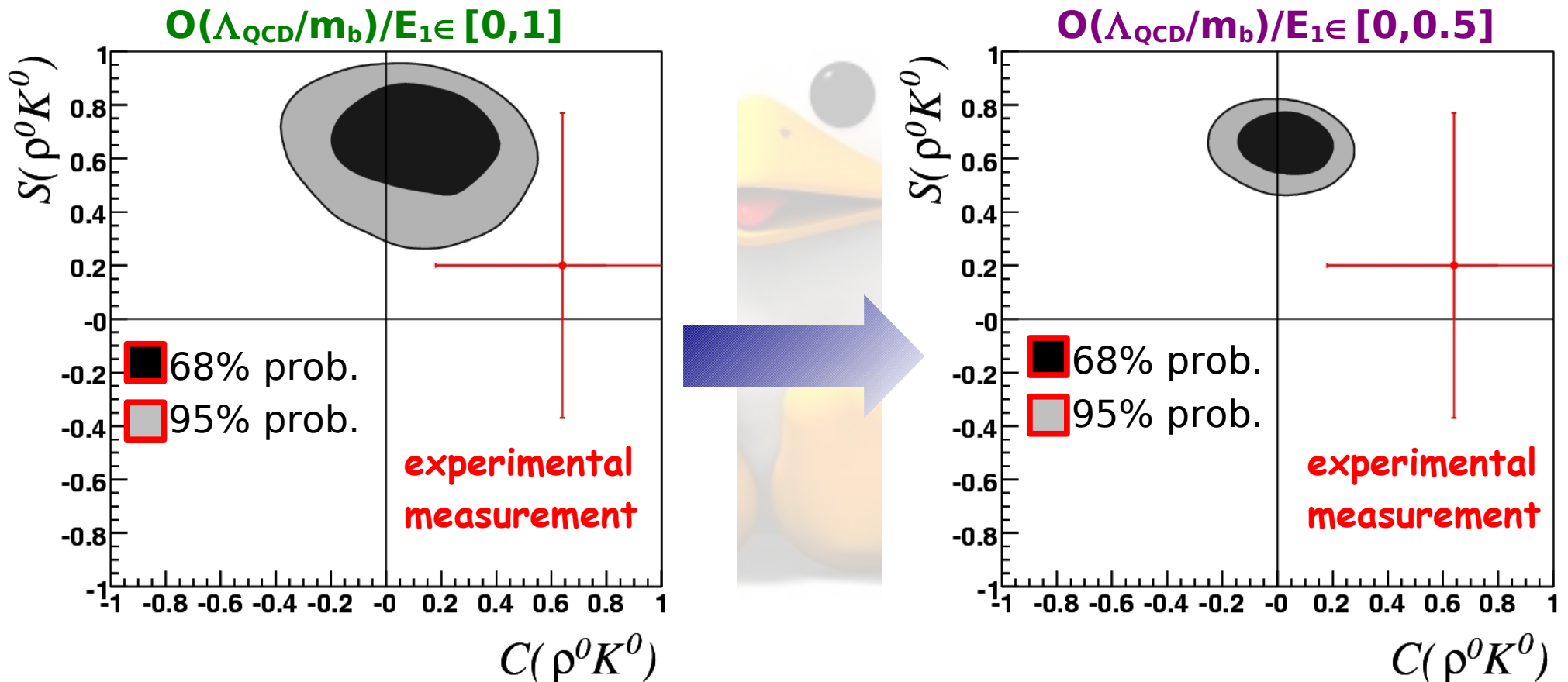
# Result on $B \rightarrow K\rho$ (III)



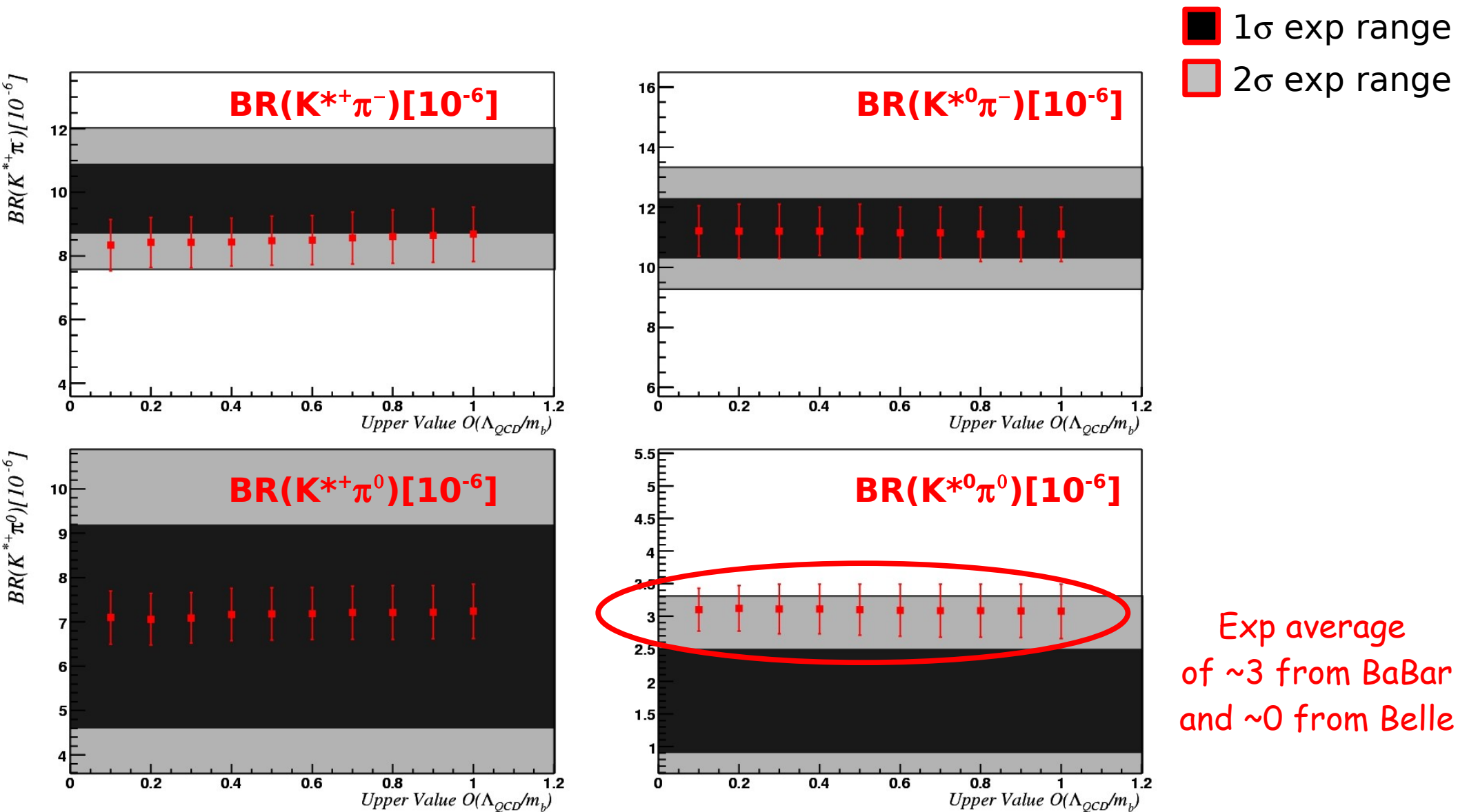
- The prediction on  $S(K^0\rho^0)$  is stable
- The error depends on the upper value of the range
- In a very conservative situation ( $O(\Lambda_{QCD}/m_b)/E_1 \in [0,1]$ ) we can still test the SM
- Limiting factor is still the exp. precision

# Test of SM: $S_{K^0}$ vs $C_{K^0}$

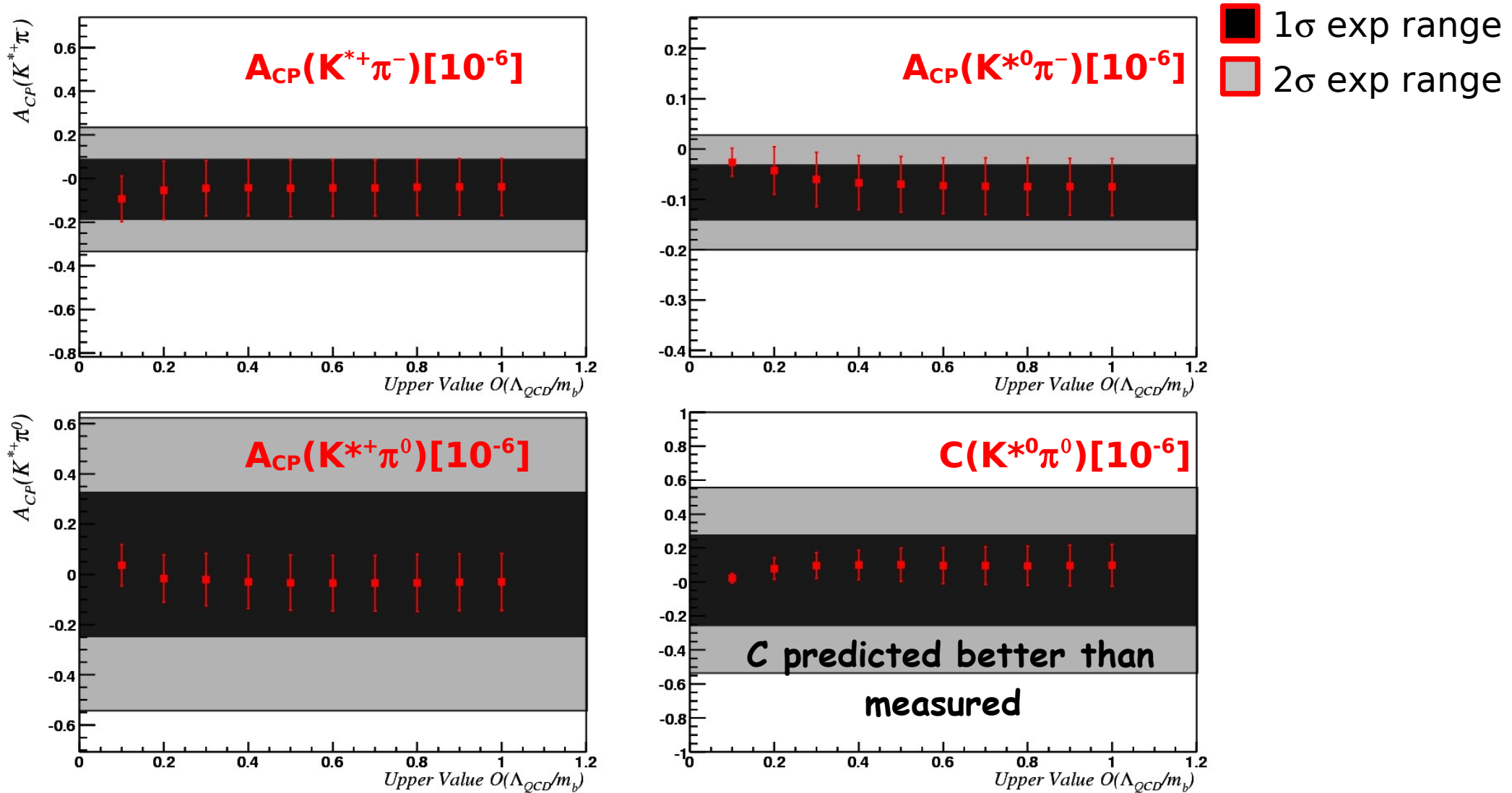
Since  $C$  is better determined by the fit than by the experiment, we have information on it from the other variables +  $SU(2)$  relations (all possible sum rules you can imagine are implemented). We can remove also  $C$  from the set of inputs and look at the agreement in the  $S$  vs  $C$  plane



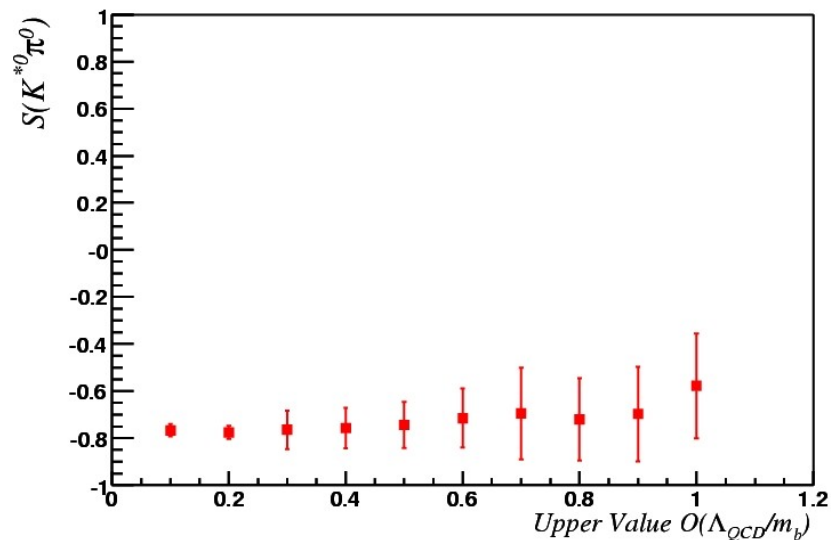
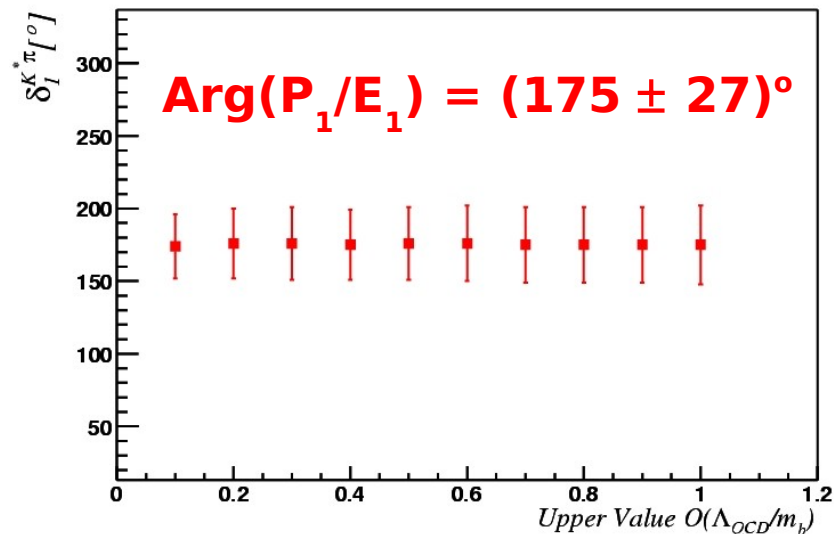
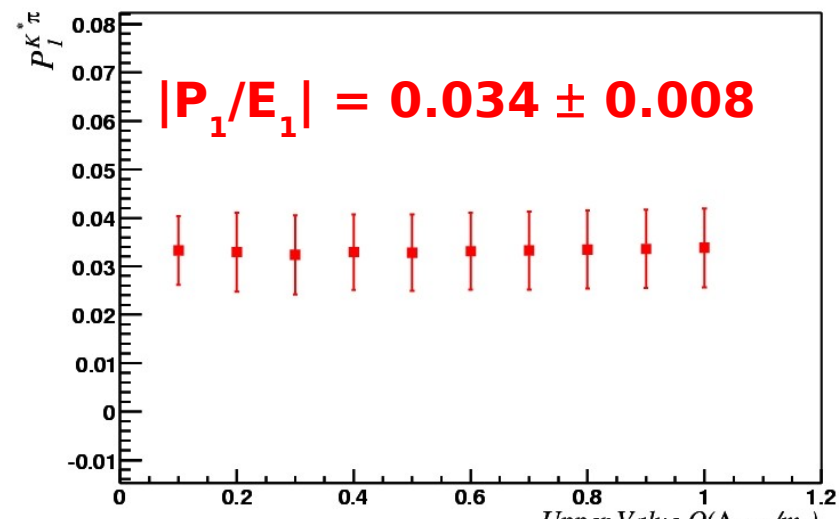
# Result on $B \rightarrow K^* \pi$ (I)



# Result on $B \rightarrow K^* \pi$ (II)



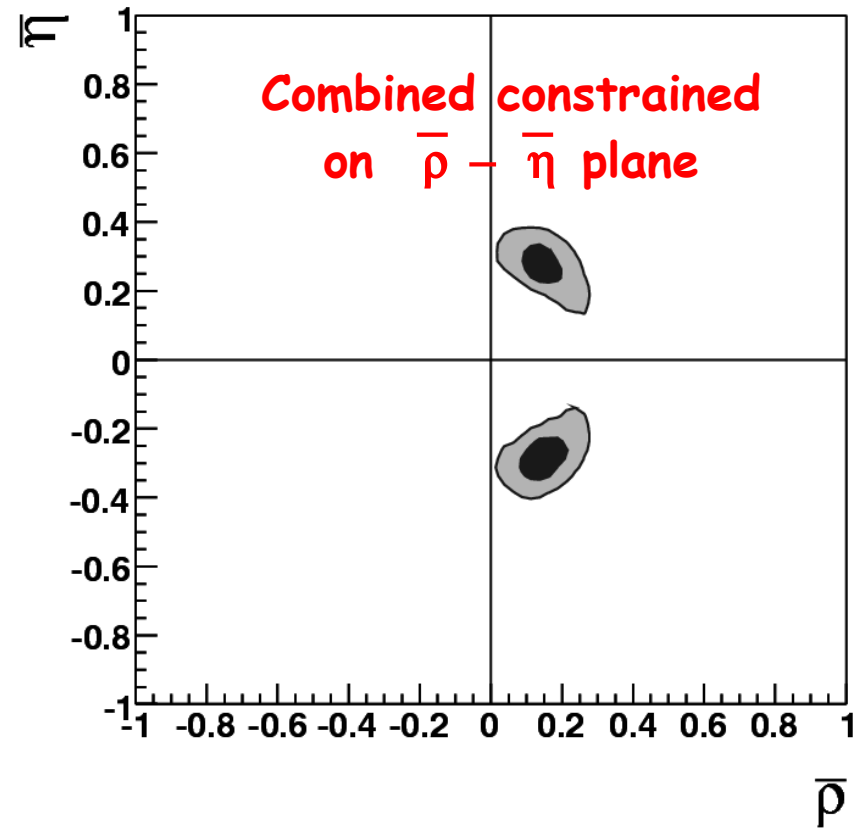
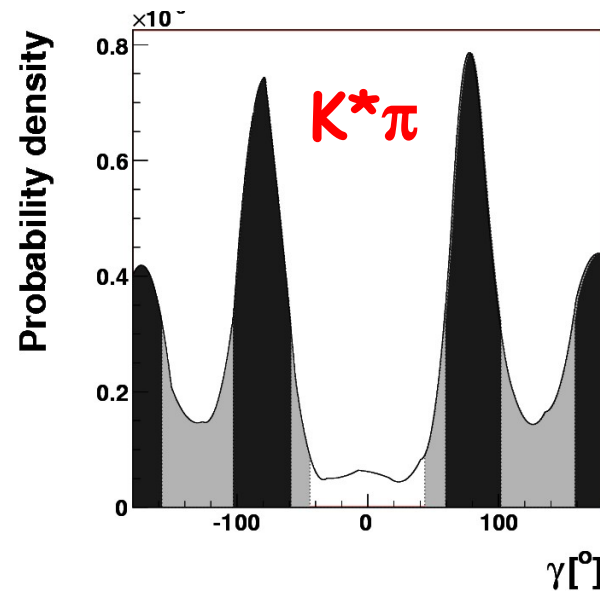
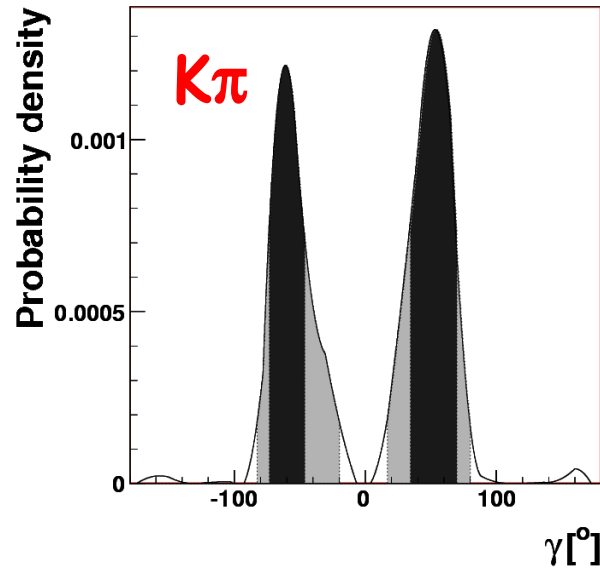
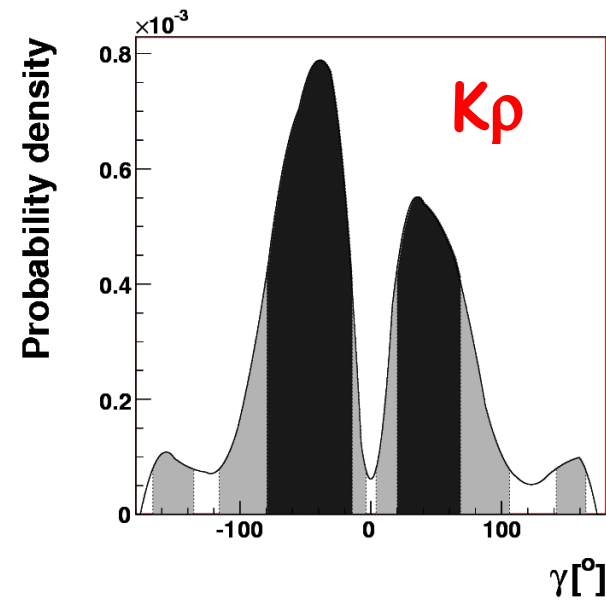
# Result on $B \rightarrow K^* \pi$ (III)



- ✚ The prediction on  $S(K^{*0} \pi^0)$  is stable
- ✚ The error depends on the upper value of the range
- ✚ In a very conservative situation  $(O(\Lambda_{\text{QCD}}/m_b)/E_1 \in [0,1])$  we can still test the SM
- ✚ Limiting factor is still the exp. precision

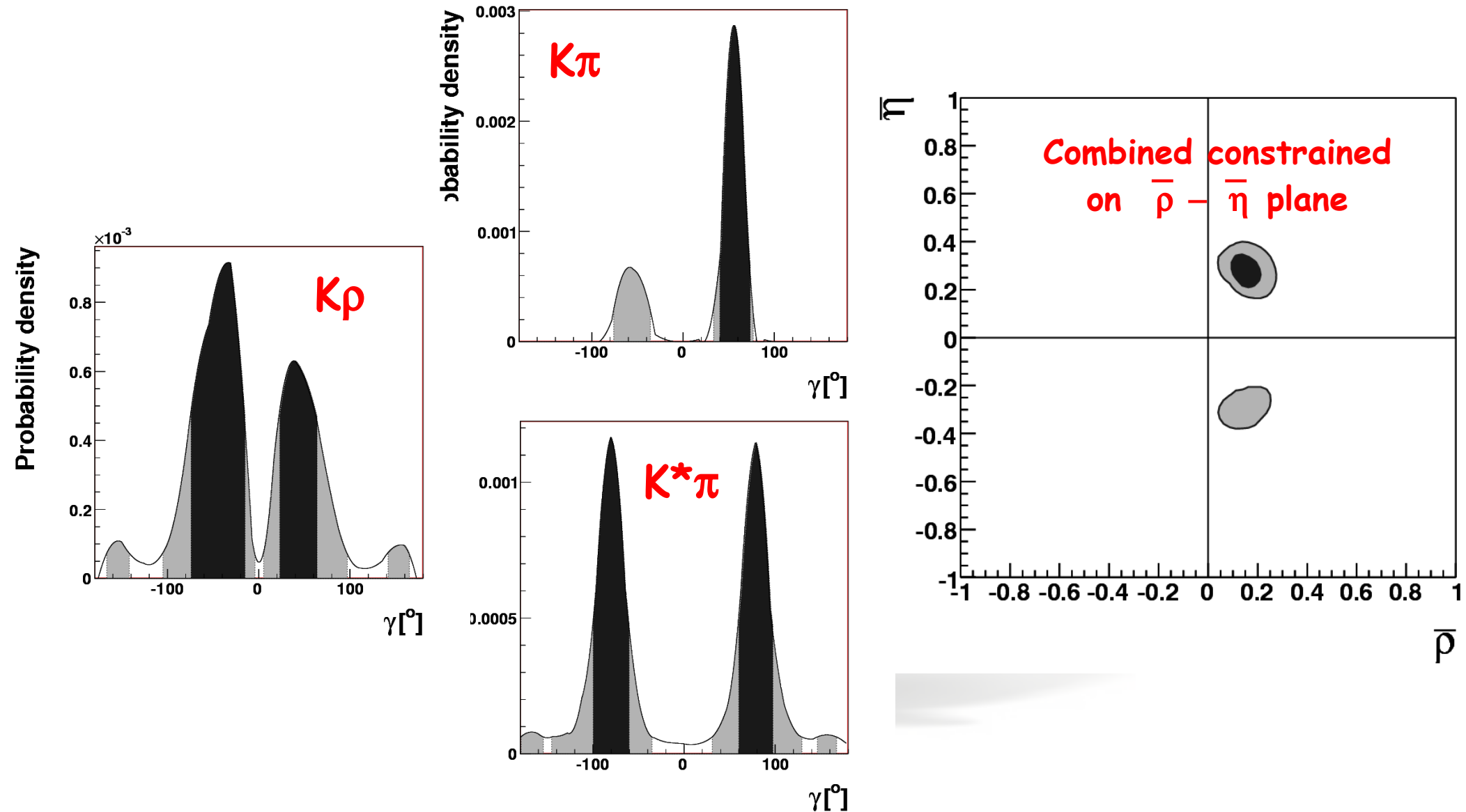
# Determination of $\gamma$ with UL@1.0

Remove the input information on  $\bar{\rho}$  and  $\bar{\eta}$  and fit for them using all BR and direct  $A_{CP}$

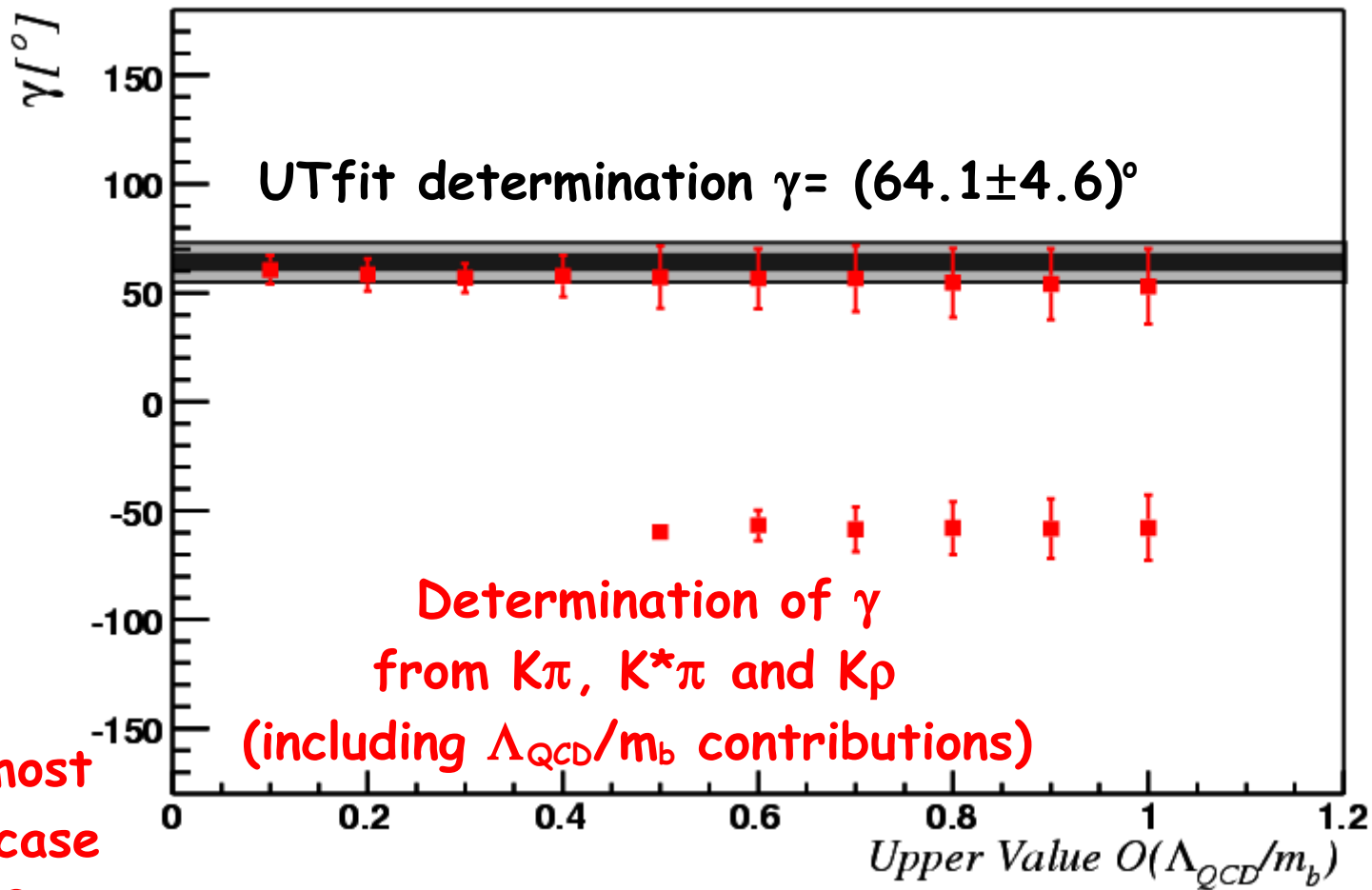




# Determination of $\gamma$ with UL@0.5



# $\gamma$ as a function of the UL



Even in the most conservative case we have a  $10^\circ$  determination of  $\gamma$  from charmless B decays

$$\begin{aligned} \gamma &= (-65 \pm 12)^\circ \cup (63 \pm 9)^\circ \text{ for UL@1.0} \\ \gamma &= (-61 \pm 3)^\circ \cup (60 \pm 10)^\circ \text{ for UL@0.5} \end{aligned}$$

# Conclusions

- + The zoology of  $b \rightarrow s$  transitions looks more rich than a land full of trees and charming penguins
  - + The perturbative calculations can describe the main features of the decays but additional effort is needed to match the experimental precision
  - + We are not sensitive yet to  $\Lambda_{\text{QCD}}/m_b$  CKM suppressed corrections, which have impact on the prediction of  $S$  in NP sensitive modes
  - + We can still obtain some information from data, but the upper value of the allowed range is needed as external input
    - ▶ too low values produce deviation from data ( $A_{\text{CP}}(K^+\pi^0)$ )
    - ▶  $UV \sim 1$  (ignoring  $\Lambda_{\text{QCD}}/m_b$  hierarchy) reduces predictive power
  - + Still, experimental measurements of  $S$  are the limiting factor of a meaningful SM test
  - + With  $\Lambda_{\text{QCD}}/m_b$  in  $[0.0, 0.5]E_1$ 
    - ▶ Good agreement with data
    - ▶ Confirmed by  $B \rightarrow KK$  data (no CKM suppression)
- See backup slides  
(or ask) for  
 $\phi K, \eta K, \omega K$   
and  $B_s \rightarrow KK$**

**In this picture,  $S(K^0\pi^0)$  emerges as the most predictive test of the SM**

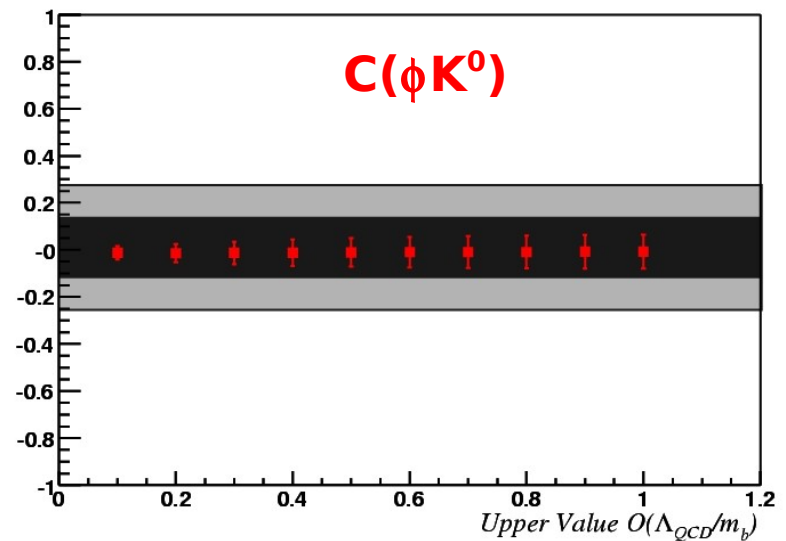
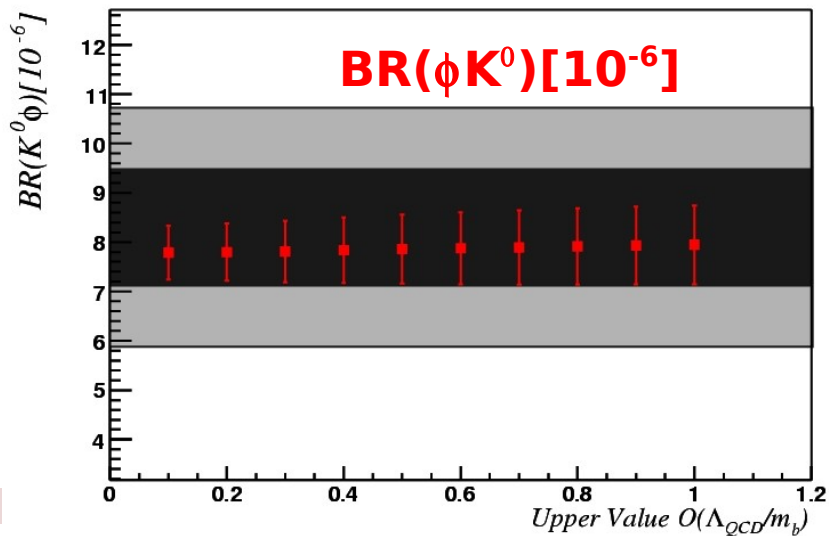
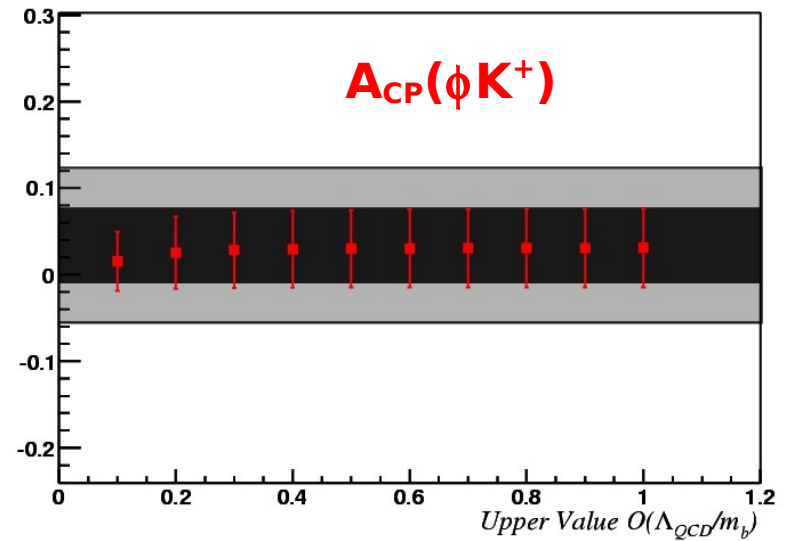
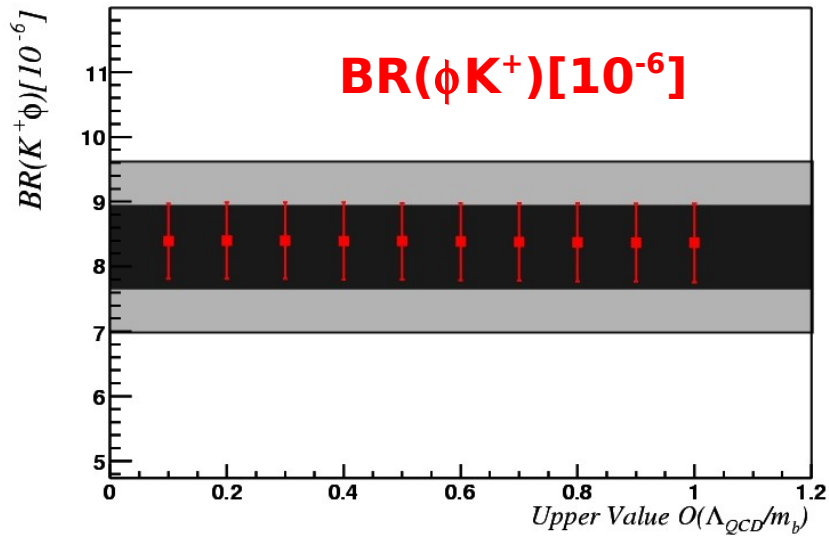
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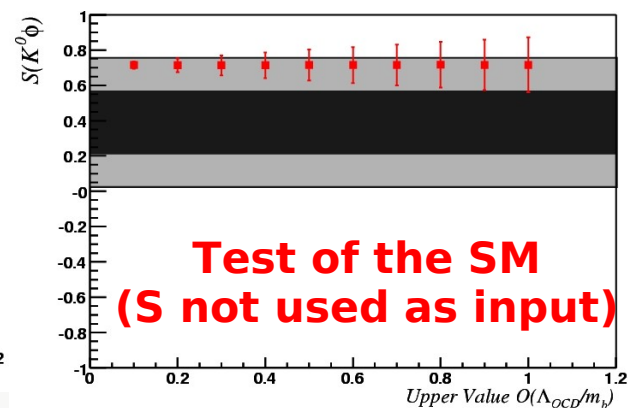
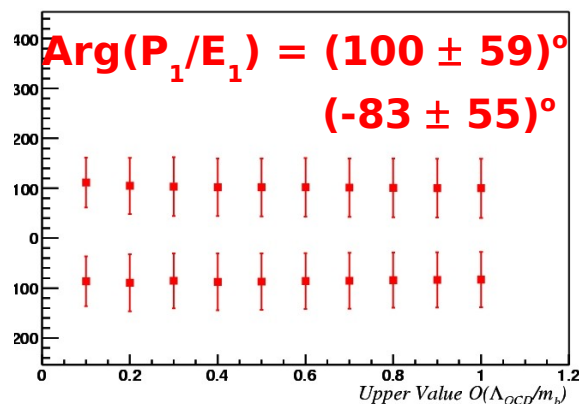
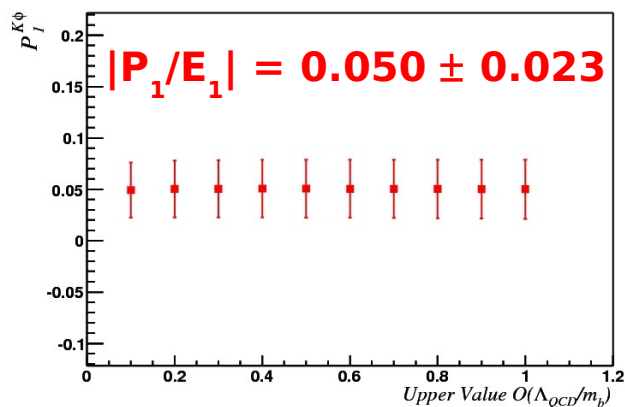
# Backup Slides

# Result on $B \rightarrow \phi K$ (I)

- $1\sigma$  exp range
- $2\sigma$  exp range

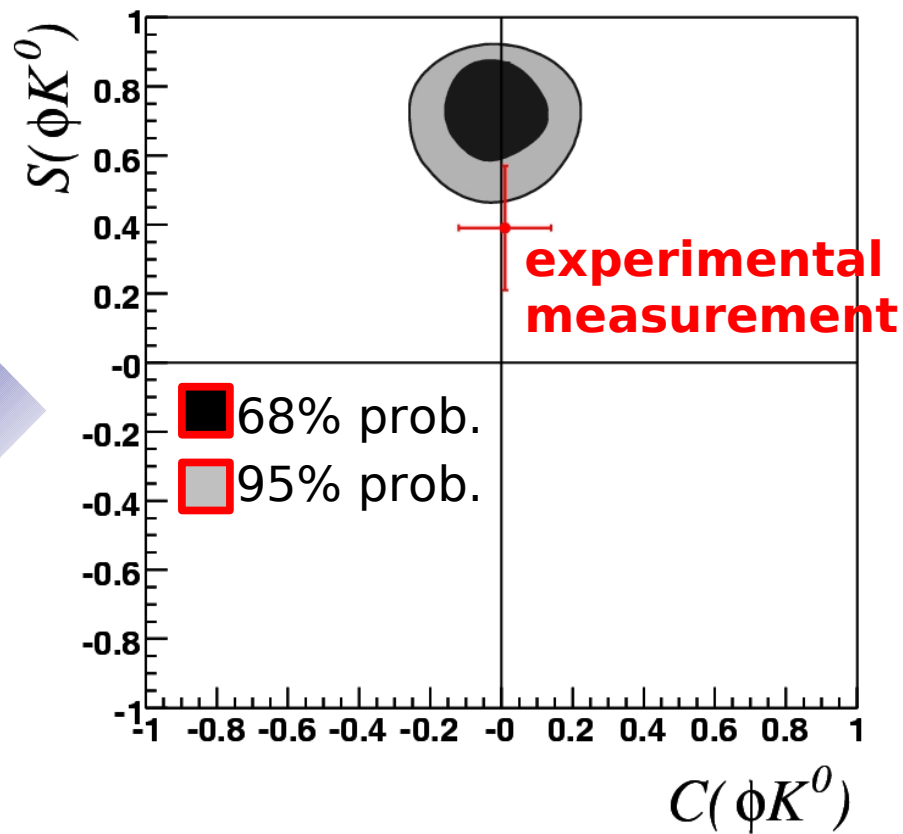
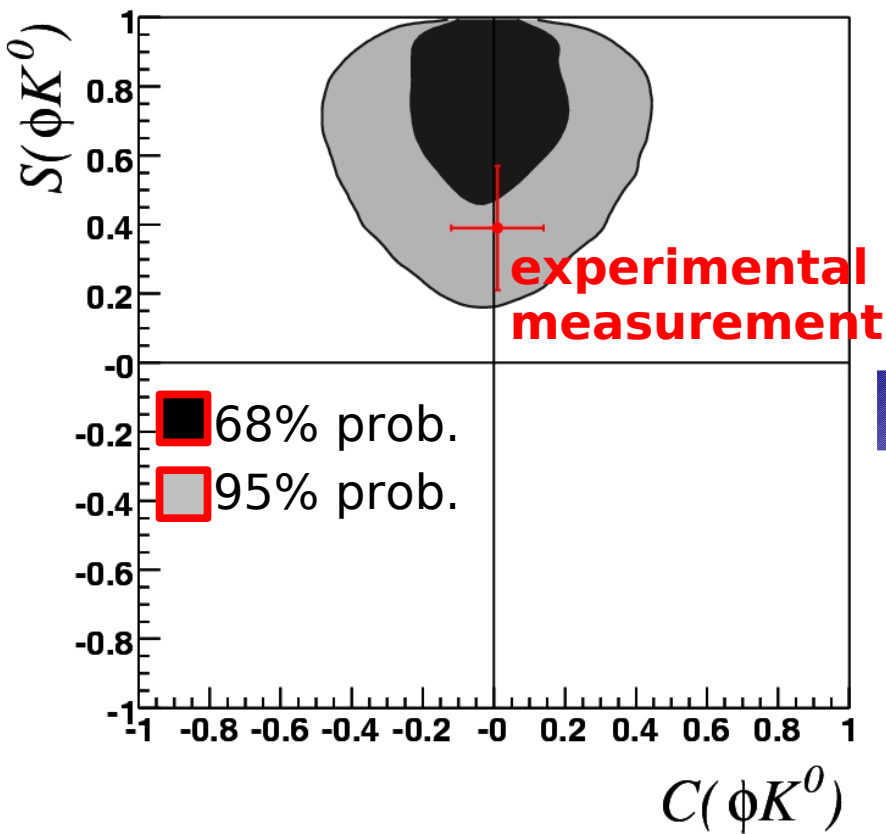


# Result on $B \rightarrow \phi K$ (II)



$O(\Lambda_{QCD}/m_b)/E_1 \in [0, 1]$

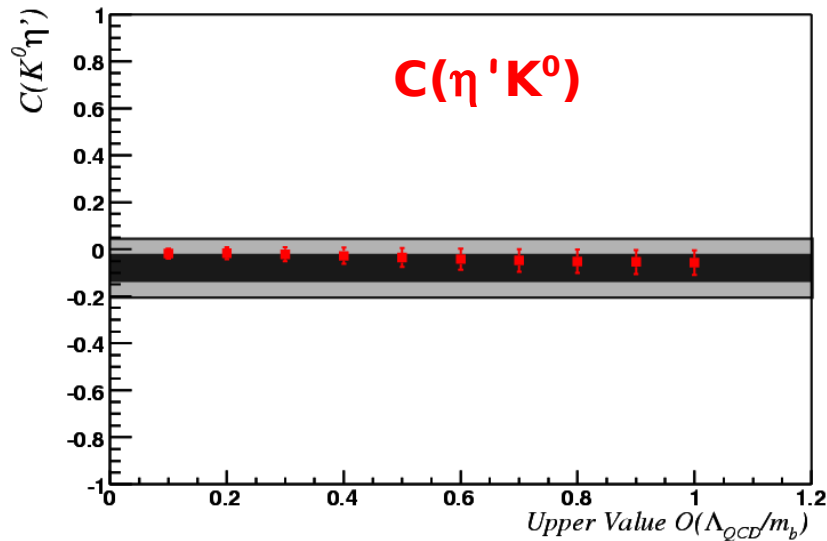
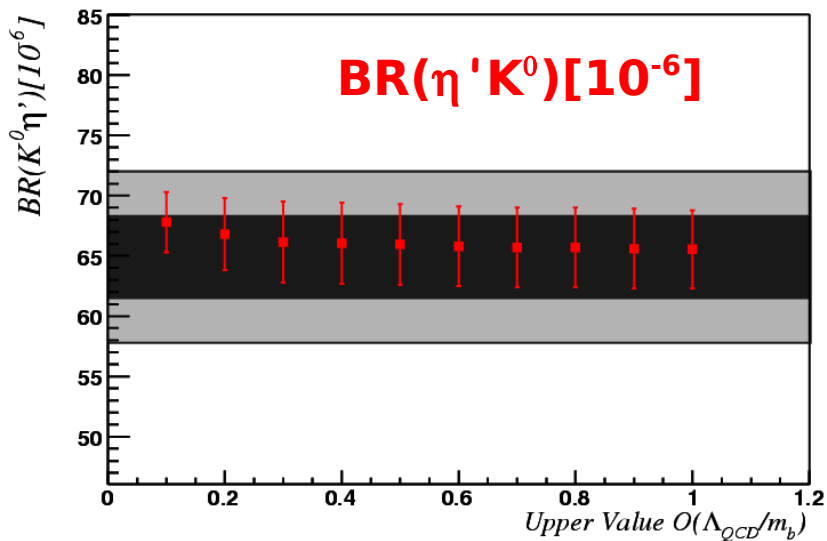
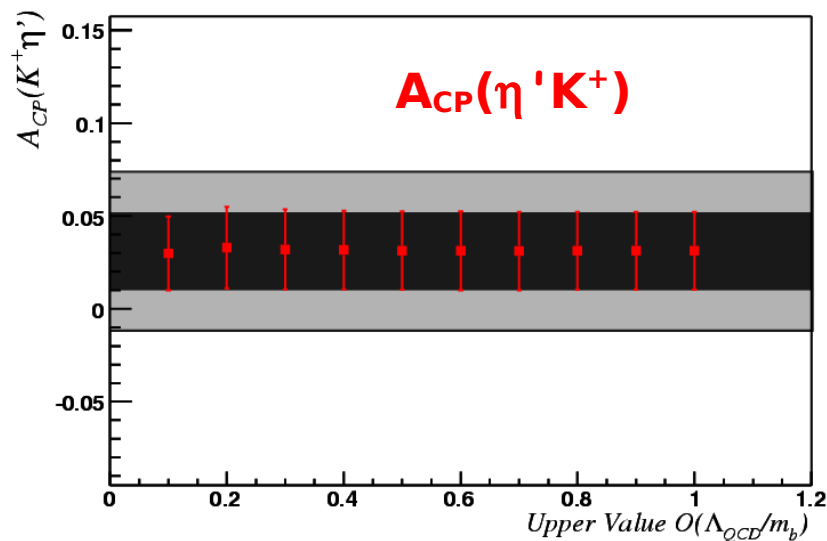
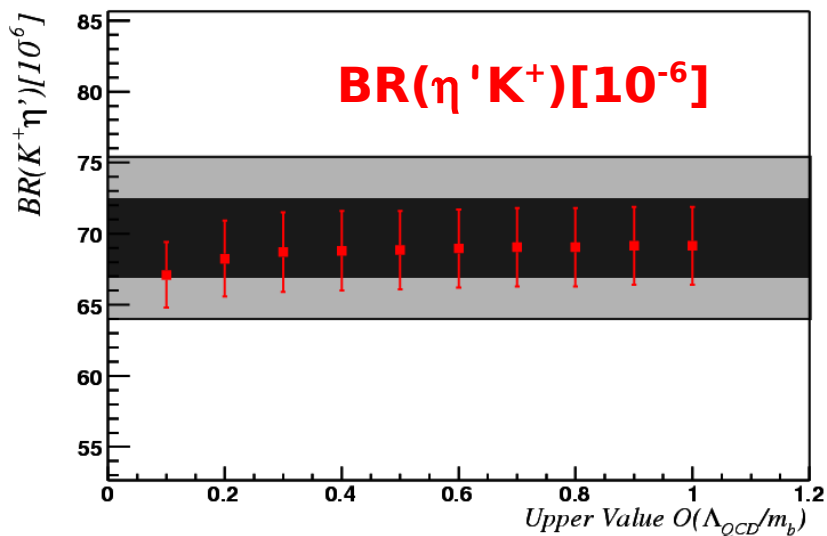
$O(\Lambda_{QCD}/m_b)/E_1 \in [0, 0.5]$



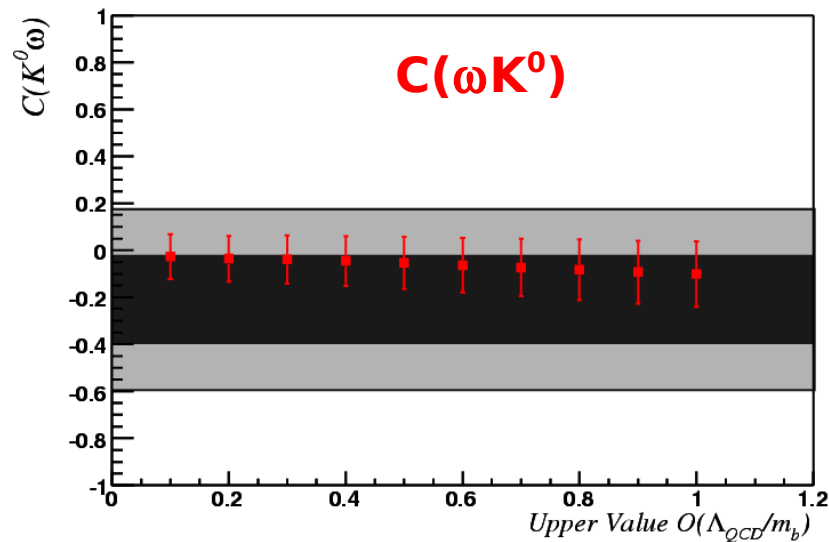
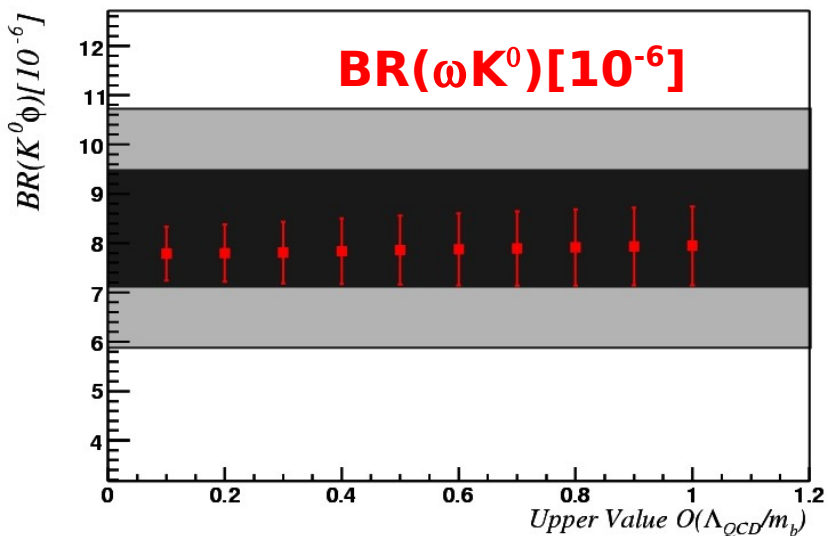
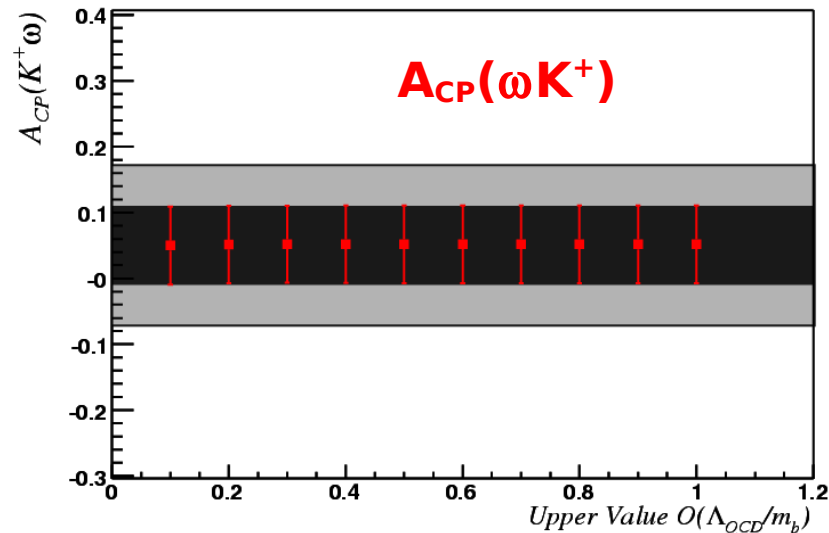
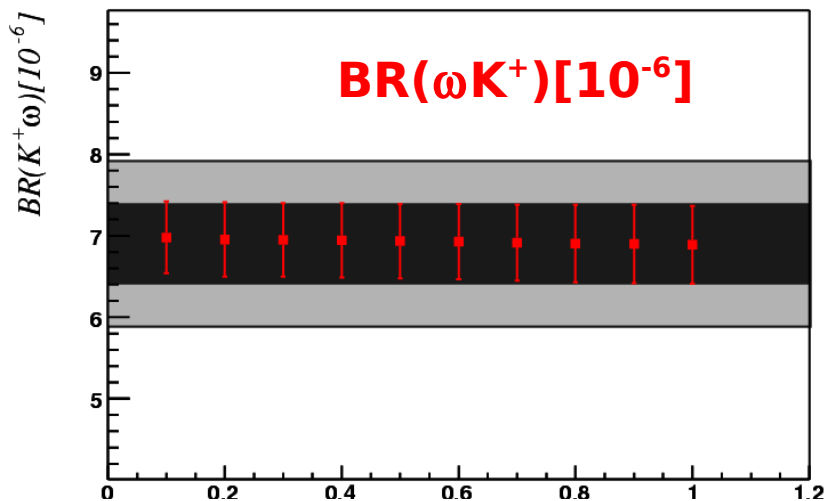
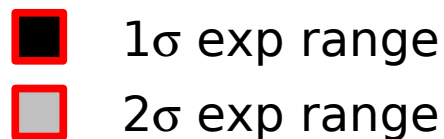
EPS200

# Result on $B \rightarrow \eta' K$ and $B \rightarrow \omega K$ (I)

- $1\sigma$  exp range
- $2\sigma$  exp range



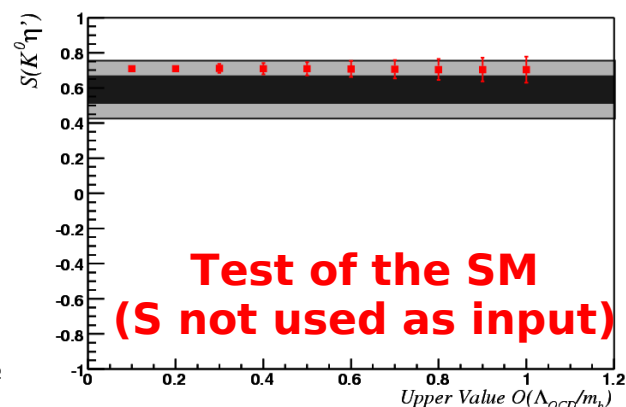
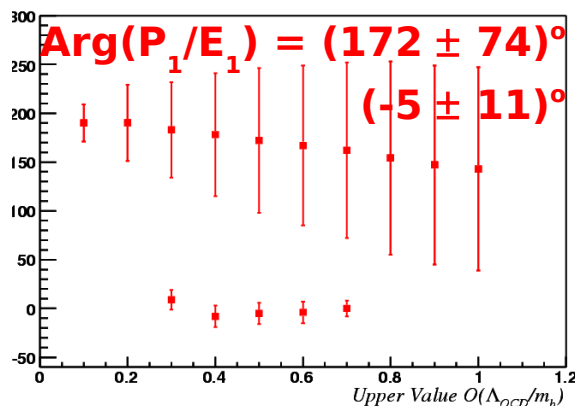
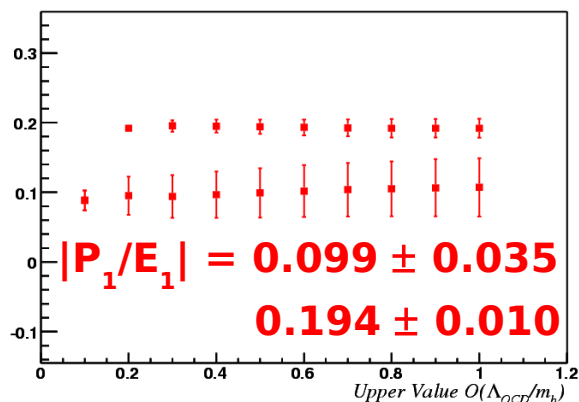
# Result on $B \rightarrow \eta' K$ and $B \rightarrow \omega K$ (II)



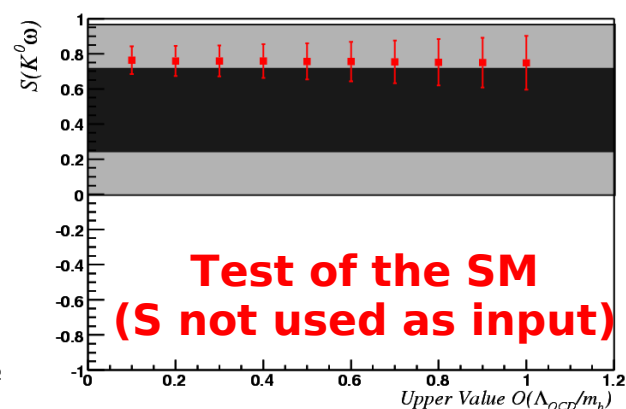
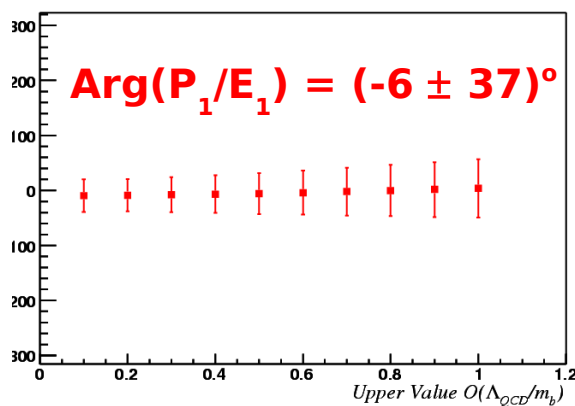
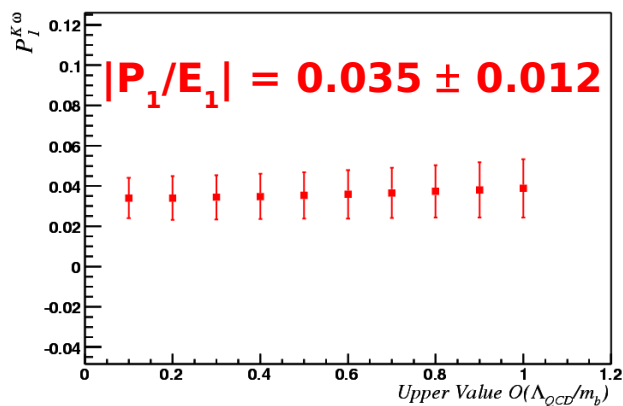


# Result on $B \rightarrow \eta' K$ and $B \rightarrow \omega K$ (III)

## $B \rightarrow \eta' K$

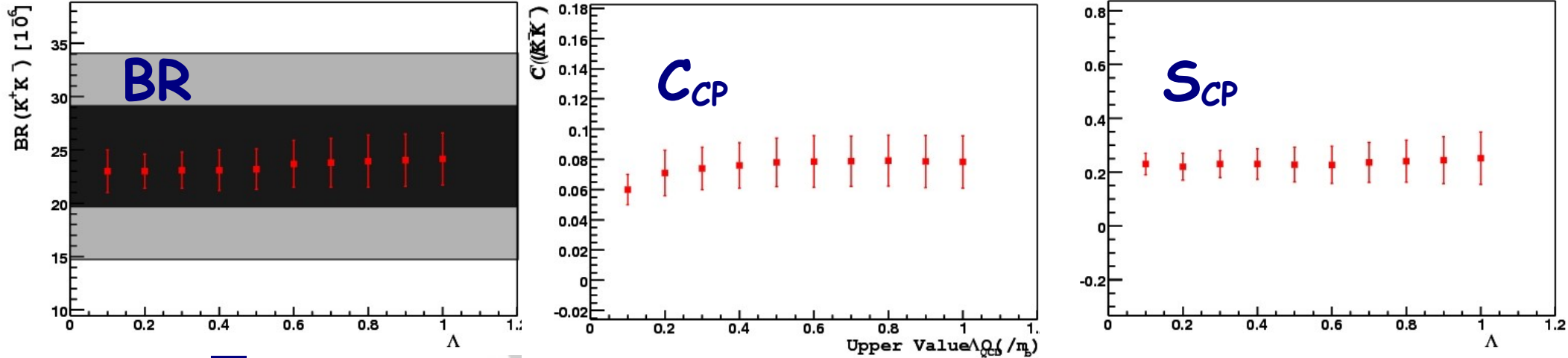


## $B \rightarrow \omega K$



# SU(3) Predictions on $B_s \rightarrow KK$

## $B_s \rightarrow K^+K^-$



## $B_s \rightarrow K^0\bar{K}^0$

