

INFLATION WITH STRONGLY NON-GEODESIC MOTION, NON-GAUSSIANITIES AND SWAMPLAND CONJECTURES

Jacopo Fumagalli

INSTITUTE D'ASTROPHYSIQUE DE PARIS

XXV International Symposium

PASCOS 2019

Particle physics, String theory and Cosmology

1–5 July 2019, Manchester, UK

Based on arXiv 1902.03221

J.F., S. Garcia-Saenz, L. Pinol, S. Renaux-Petel, J. Ronayne

GEODESI



European Research Council
Established by the European Commission



**SORBONNE
UNIVERSITÉ**

Top-Down

↳ Negative curvature
in field space
eg α -ATTRACTORS

UV EMBEDDINGS
&
INFLATION

Swampland
Conjectures

← REQUIRE MULTIFIELD INFLATION
WITH LARGE BENDING
A. Achucarro & G. Palma '18

γ -Problem

← REQUIRE INFLATING ON
STEEP POTENTIALS

Top-Down

↳ Negative curvature
in field space
eg α -ATTRACTORS



NEW
ATTRACTORS

SIDETRACKED, HYPERINFLATION,
ANGULAR INFLATION.....

S. Garcia Saenz, S. Renaux-Petel & J.
Ronayne '18, A. Brown '17, D. Marsh &
T. Bjorkmo '19. P. Christodoulidis, D. Rest, E. Spaliaris

.....

UV EMBEDDINGS
&
INFLATION

Swampland
Conjectures

← REQUIRE MULTIFIELD INFLATION
WITH LARGE BENDING

A. Achucarro & G. Palma '18

γ-Problem

← REQUIRE INFLATING ON
STEEP POTENTIALS

PLAN

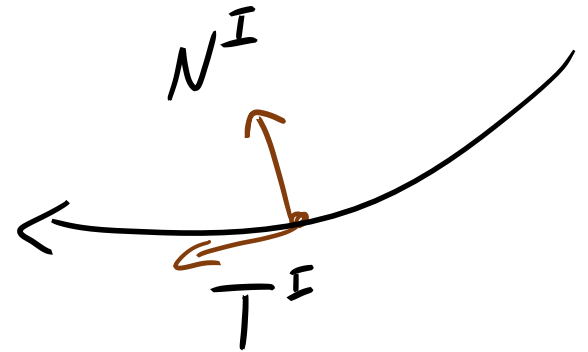
- INFLATION IN NON-GEODESIC MOTION AND NEW ATTRACTORS
- PERTURBATIONS \longrightarrow EFFECTIVE FIELD THEORY WITH IMAGINARY SPEED OF SOUND
- MATCHING THE EFT AND GOING BEYOND THE BISPECTRUM
- MODEL INDEPENDENT BOUNDS
- EXAMPLE: HYPERINFLATION

INFLATION IN NON-GEODESIC MOTION

$$\mathcal{L} = -\frac{1}{2} G_{IJ} \partial\phi^I \partial\phi^J - V(\phi) + M_{\text{pl}}^2 \frac{\mathcal{R}}{2}$$

$$\text{EOM} \equiv D_t \dot{\phi}^I + 3H \dot{\phi}^I + G^{IJ} V_{,J} = 0$$

$$\downarrow$$
$$\ddot{\phi}^I + \Gamma_{JK}^I \dot{\phi}^J \dot{\phi}^K$$



$$\text{Def. } \dot{\sigma}^2 = G_{IJ} \dot{\phi}^I \dot{\phi}^J, \quad T^I = \frac{\dot{\phi}^I}{\dot{\sigma}}, \quad N^I \perp T^I$$

INFLATION IN NON-GEODESIC MOTION

$$N^I(\text{EOM}) \rightarrow D_t T^I = \gamma_{\perp} H N^I \quad \gamma_{\perp} = -\frac{V_N}{H\dot{\sigma}}$$

INFLATION IN NON-GEODESIC MOTION

$$N^I(\text{EOM}) \rightarrow D_t T^I = \gamma_{\perp} H N^I \quad \gamma_{\perp} = -\frac{V_N}{H\dot{\sigma}}$$

$\gamma_{\perp} = 0 \Rightarrow$ GEODESIC MOTION, $D_t T^I = 0$

$\gamma_{\perp} \gg 1 \Rightarrow$ STRONGLY NON-GEODESIC



INFLATION IN NON-GEODESIC MOTION

$$N^I(\text{EOM}) \rightarrow D_t T^I = \gamma_{\perp} H N^I \quad \gamma_{\perp} = -\frac{V_N}{H\dot{\sigma}}$$

$$\gamma_{\perp} = 0 \Rightarrow \text{GEODESIC MOTION, } D_t T^I = 0$$

$$\gamma_{\perp} \gg 1 \Rightarrow \text{STRONGLY NON-GEODESIC}$$



SWAMPLAND REMARK A. Achúcarro & G. Palma '18

$$\epsilon = \frac{\epsilon_V}{\left(1 + \frac{\gamma_{\perp}^2}{g}\right)}, \quad \epsilon_V = \frac{M_{\text{pl}}^2}{2} \left(\frac{V'}{V}\right)^2 \sim \mathcal{O}(1)$$

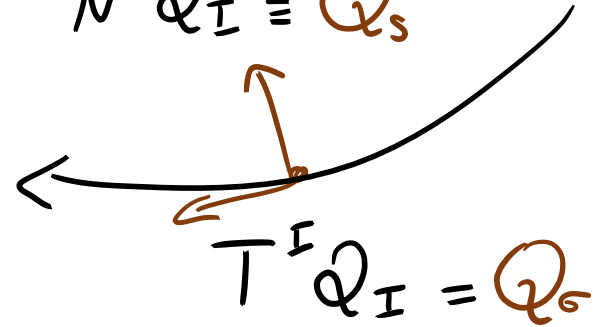
$$\not\Rightarrow \epsilon = -\frac{\dot{H}}{H^2} \sim \mathcal{O}(1)$$

MULTI-FIELD PERTURBATIONS

- COVARIANT PERTURBATIONS (FLAT GAUGE)

$$\phi^I - \phi_0^I = \delta\phi^I = \varrho^I - \Gamma_{JK}^I \varrho^J \varrho^K$$

$$N^I \varrho_I \equiv \varrho_S$$

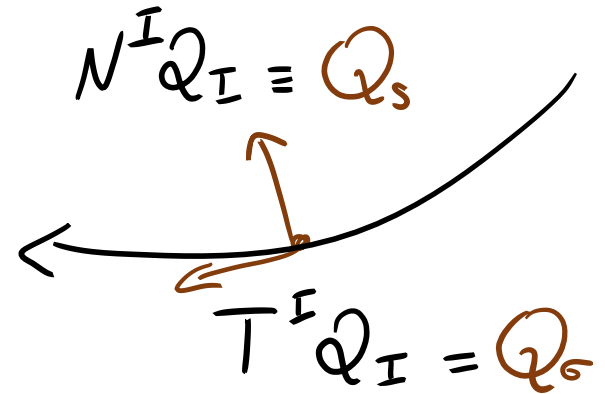


$$T^I \varrho_I = \varrho_G$$

MULTI-FIELD PERTURBATIONS

- COVARIANT PERTURBATIONS (FLAT GAUGE)

$$\phi^I - \phi_0^I = \delta\phi^I = Q^I - \Gamma_{JK}^I Q^J Q^K$$



- EOM ENTROPIC FLUCTUATION

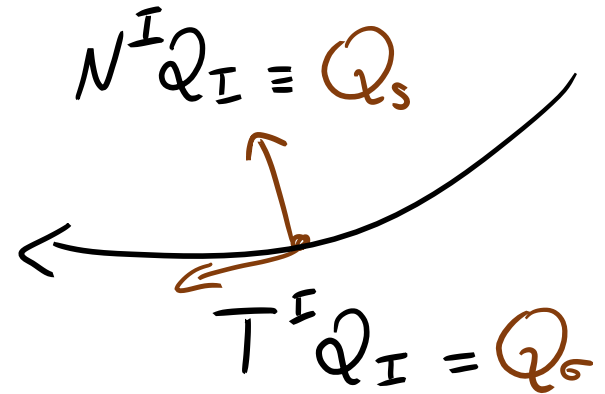
$$\ddot{Q}_s + 3H\dot{Q}_s + \left(\frac{k^2}{a^2} + m_s^2\right)Q_s = -\alpha\gamma_{\perp}\zeta$$

$$V_{,ss} + \epsilon R_{sf} H^2 \mathcal{R}_{pe}^2 - H^2 \gamma_{\perp}^2$$

MULTI-FIELD PERTURBATIONS

- COVARIANT PERTURBATIONS (FLAT GAUGE)

$$\phi^I - \phi_0^I = \delta\phi^I = Q^I - \Gamma_{JK}^I Q^J Q^K$$



- EOM ENTROPIC FLUCTUATION

$$\ddot{Q}_s + 3H\dot{Q}_s + \left(\frac{k^2}{a^2} + m_s^2\right)Q_s = -\alpha \zeta_\perp \dot{\zeta}$$

$$V_{,ss} + \epsilon R_{sf} H^2 \mathcal{R}_{pe}^2 - H^2 \zeta_\perp^2$$

- N.B. $\zeta_\perp^2 \gg 1 \implies m_s^2 < 0 ; |m_s| \gg H$

MULTI-FIELD PERTURBATIONS

$$m_s^2 < 0 \quad \not\Rightarrow \quad \text{UNSTABLE BKG}$$

MULTI-FIELD PERTURBATIONS

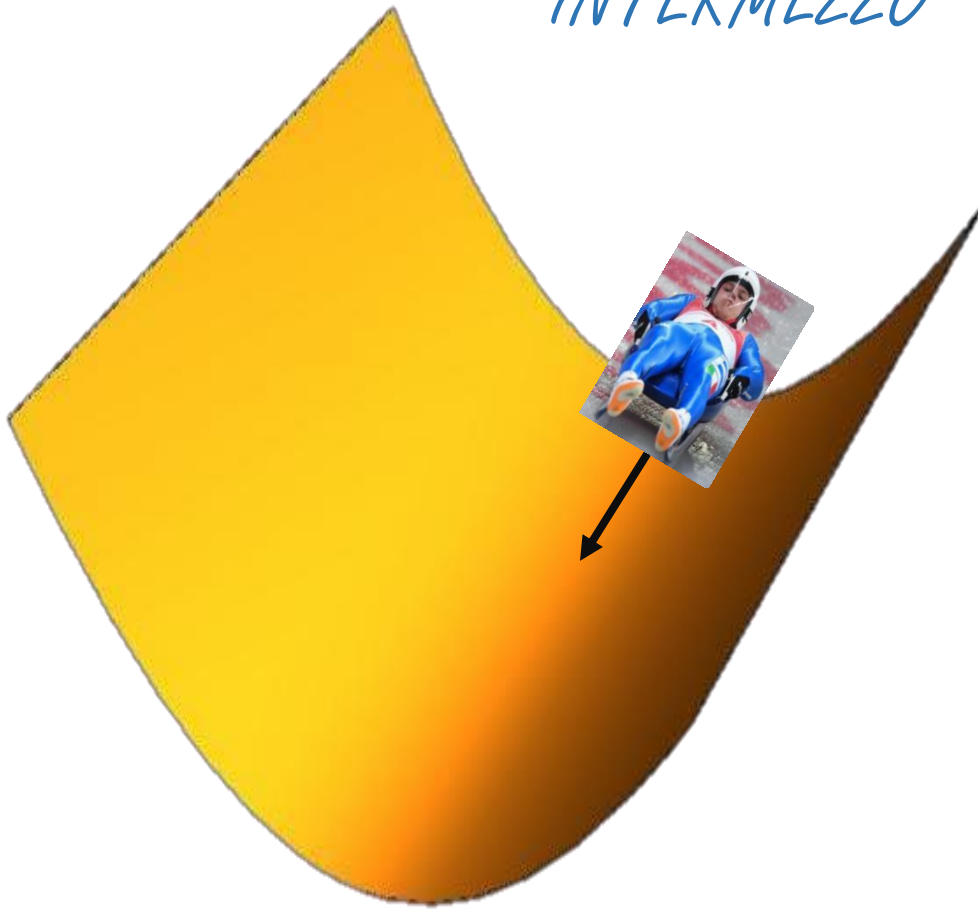
$$m_s^2 < 0 \quad \not\Rightarrow \quad \text{UNSTABLE BKG}$$

• SUPER-HUBBLE $\frac{k}{a} \ll H$; $\dot{\gamma} = 2H^2 \gamma_{\perp} Q_s \rightarrow \text{EOM}[Q_s]$

$$\Rightarrow \ddot{Q}_s + 3H\dot{Q}_s + m_{s,\text{EFF}}^2 Q_s = 0$$

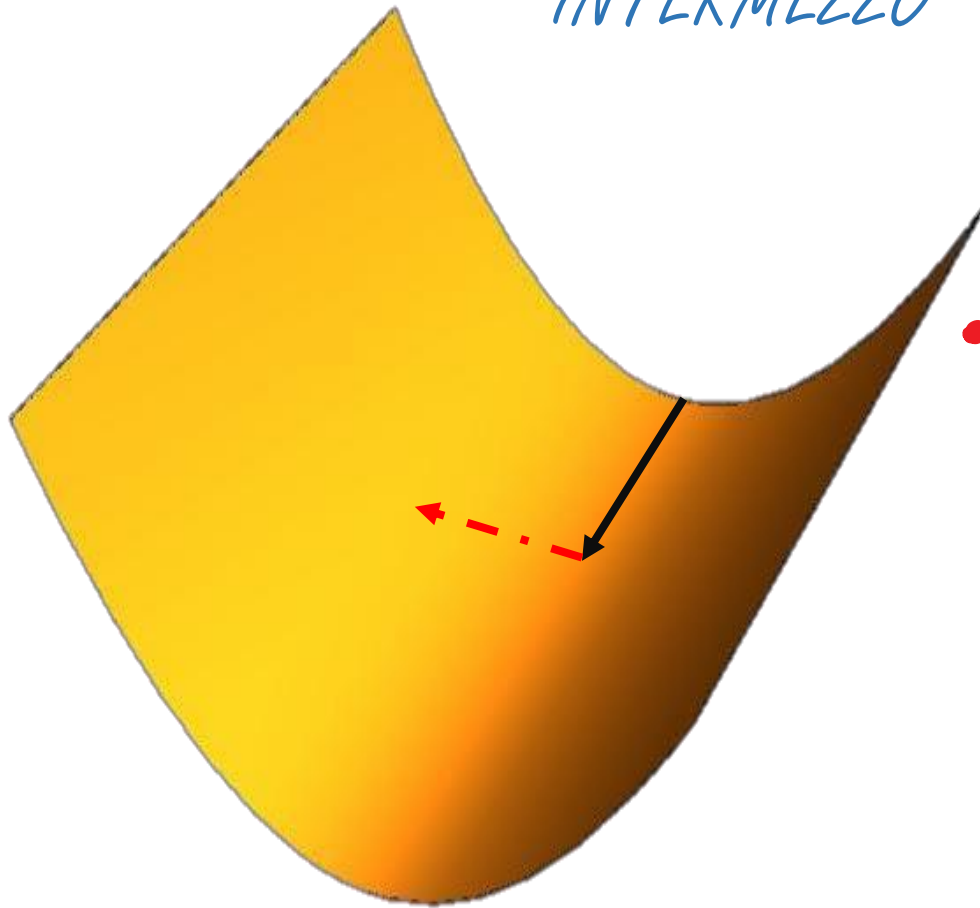
$$m_{s,\text{EFF}}^2 = \underbrace{V_{ss} + \epsilon R_{sf} H^2 \kappa_{pe}^2 - H^2 \gamma_{\perp}^2}_{m_s^2 < 0} + 4H^2 \gamma_{\perp}^2 > 0$$

INTERMEZZO



$$m_{S, \text{EFF}}^2 = V_{SS} + \epsilon R_{sf} H^2 \kappa_{pe}^2 > 0$$

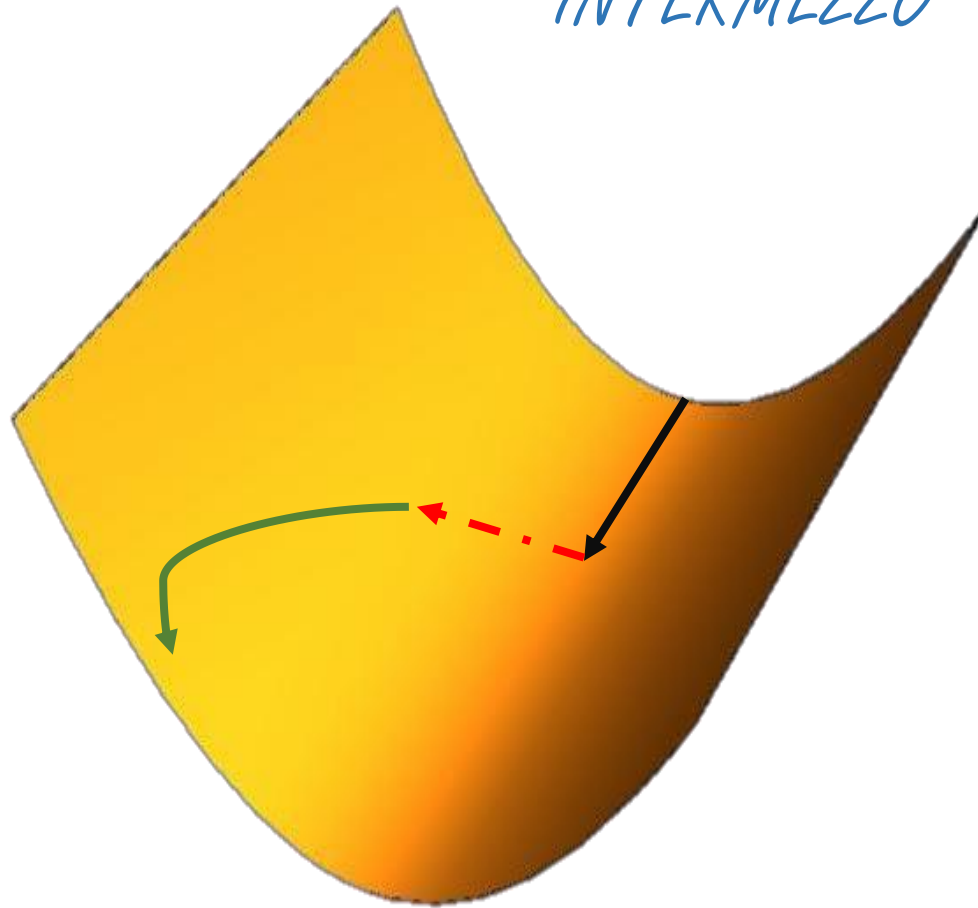
INTERMEZZO



- GEOMETRICAL DESTABILIZATION
S. Renaux-Petel & K. Turzinsky '15

$$m_{S, \text{EFF}}^2 = V_{SS} + \epsilon R_{Sf} H^2 \mathcal{R}_{Pe}^2 < 0$$

INTERMEZZO



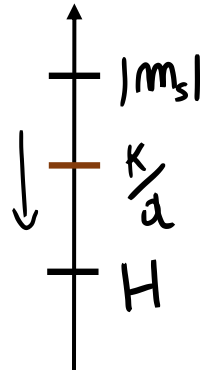
- *SIDETRACKED*
S. Garcia Saenz, S. Renaux-Petel & J. Ronayne '18
- HYPERINFLATION*
A. Brown '17
D. Marsh & T. Bjorkmo '19
- ANGULAR INFLATION*
P. Christodoulidis, D. Roest, E. Sfakianakis
-

$$m_{S, \text{EFF}}^2 = \underbrace{V_{SS} + \epsilon R_{sf} H^2 \kappa_{pe}^2 - H^2 \gamma_{\perp}^2}_{m_s^2 < 0} + 4 H^2 \gamma_{\perp}^2 > 0$$

SINGLE FIELD EFFECTIVE FIELD THEORY

- INTEGRATING OUT THE ENTROPIC MODE $|m_s^2| \gg \frac{k^2}{a^2}$

$$Q_s^{\text{EFT}} = -\frac{2\delta\gamma_{\perp}}{m_s^2} \dot{\gamma} \longrightarrow \int_2^{\text{DEFT}} [\dot{\gamma}] = \int dt d^3x \epsilon a^2 \left[\frac{(\dot{\gamma})^2}{c_s^2} - (\partial_i \gamma)^2 \right]$$



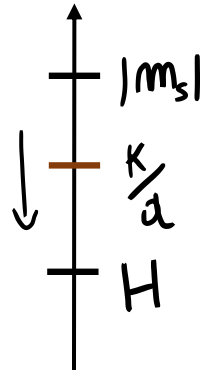
$$\bullet \frac{1}{c_s^2} = 1 + \frac{4H^2 m_{\perp}^2}{m_s^2}$$

A. Achúcarro et al. '11

SINGLE FIELD EFFECTIVE FIELD THEORY

- INTEGRATING OUT THE ENTROPIC MODE $|m_s^2| \gg \frac{k^2}{a^2}$

$$Q_s^{\text{EFT}} = -\frac{2\delta\gamma_{\perp}}{m_s^2} \dot{\gamma} \longrightarrow \int_2^{\text{EFT}} [\dot{\gamma}] = \int dt d^3x \epsilon a^2 \left[\frac{(\dot{\gamma})^2}{c_s^2} - (\partial_i \gamma)^2 \right]$$



$$\bullet \frac{1}{c_s^2} = 1 + \frac{4H^2 \gamma_{\perp}^2}{m_s^2} = \frac{m_{s,\text{eff}}^2}{m_s^2} > 0 \iff \text{Stable Background}$$

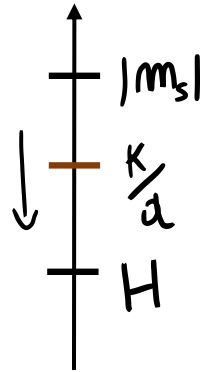
$$m_s^2 < 0 \iff \gamma_{\perp}^2 \gg 1$$

A. Achúcarro et al. '11

SINGLE FIELD EFFECTIVE FIELD THEORY

- INTEGRATING OUT THE ENTROPIC MODE $|m_s^2| \gg \frac{k^2}{a^2}$

$$Q_s^{\text{EFT}} = -\frac{2\delta\gamma_{\perp}}{m_s^2} \dot{\gamma} \longrightarrow \int_2^{\text{EFT}} [\dot{\gamma}] = \int dt d^3x \epsilon a^2 \left[\frac{(\dot{\gamma})^2}{c_s^2} - (\partial_i \gamma)^2 \right]$$



$$\bullet \frac{1}{c_s^2} = 1 + \frac{4H^2 \gamma_{\perp}^2}{m_s^2} = \frac{m_{s,\text{eff}}^2}{m_s^2} > 0 \iff \text{Stable Background}$$

$$m_s^2 < 0 \iff \gamma_{\perp}^2 \gg 1$$

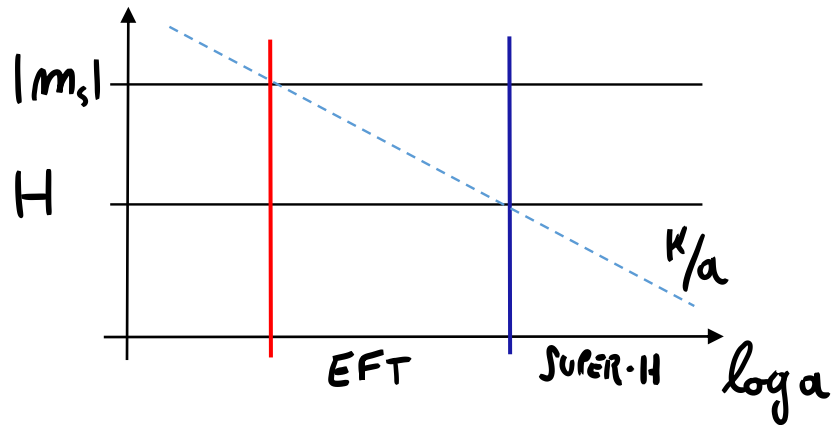
A. Achúcarro et al. '11

$$\implies \boxed{c_s^2 < 0 \quad \text{EFT WITH IMAGINARY SPEED OF SOUND}}$$

S. Garcia Saenz, S. Renaux-Petel '18
 J.F., S. Garcia Saenz, L. Pinol, S. Renaux-Petel, J. Ronayne '19

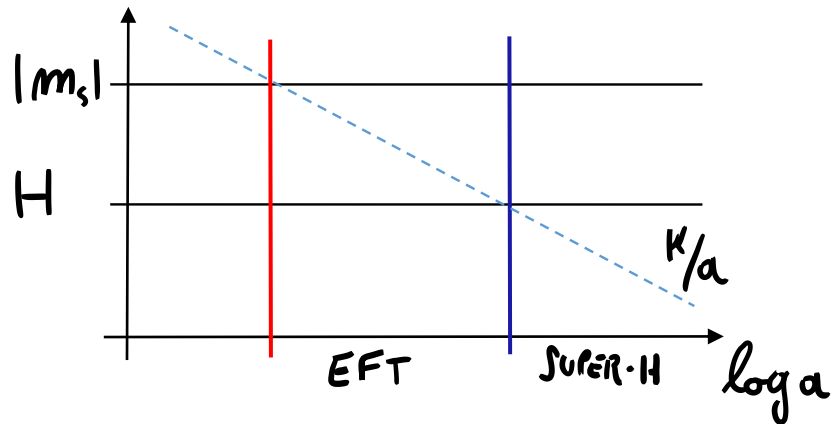
SINGLE FIELD EFFECTIVE FIELD THEORY

TRANSIENT INSTABILITY



SINGLE FIELD EFFECTIVE FIELD THEORY

TRANSIENT INSTABILITY



Validity of the EFT

$$\frac{|c_s| k}{a} = x H, \quad x \leq \frac{|m_s| |c_s|}{H}$$

$$N_{\text{S.H.}} - N_{\text{EFT}} = \ln x$$

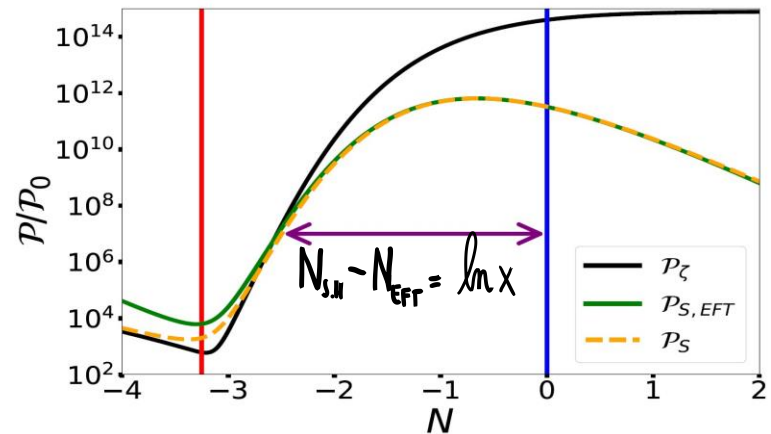
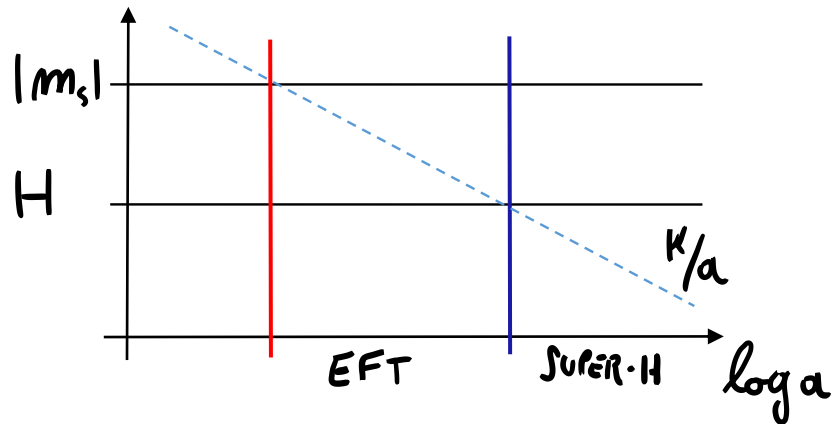
• MODE FUNCTION SOLUTION

$$\zeta_s(\tau) = \left(\frac{2\pi}{k^3}\right)^{\frac{1}{2}} \alpha \cdot \left[e^{k|c_s|\tau + x} \left(k|c_s|\tau - 1\right) - \beta e^{i\tau} e^{-(k|c_s|\tau + x)} \left(k|c_s|\tau + 1\right) \right]$$

Overall amplitude
Growing mode
Decaying mode

SINGLE FIELD EFFECTIVE FIELD THEORY

TRANSIENT INSTABILITY



- MODE FUNCTION SOLUTION $\longrightarrow P_i = \alpha^2 e^{2x}$

$$\zeta(\tau) = \left(\frac{2\pi}{k^3}\right)^{\frac{1}{2}} \alpha \cdot \left[e^{k|c_s|\tau + x} (k|c_s|\tau - 1) - \rho e^{i\tau} e^{-(k|c_s|\tau + x)} (k|c_s|\tau + 1) \right]$$

Overall amplitude
Growing mode
Decaying mode

ROADMAP: USEFULNESS OF THE EFT

EFT \rightarrow $\frac{\langle \gamma^3 \rangle}{\langle \gamma^2 \rangle^2} \sim f_{NL}(A, X)$ **MATCHING**

NUMERICS

PyTransport 2.0

D. Mubryne and J. Ronayne

$\int_{\text{EFF}}^3 = \int dt d^3x \frac{a \epsilon M_{Pl}^2}{H} \left(\frac{1}{c_s^2} - 1 \right) \left[\dot{\gamma}^i (\partial_i \gamma)^2 + \frac{A}{c_s^2} \gamma'^3 \right]$

WHY THE EFFECTIVE FIELD THEORY?

↳ Understanding enhancement of certain shapes (e.g. flattened NG)

↳ More importantly, beyond 3pt functions:

Estimate higher order correlators $\langle \gamma^m \rangle$ and set perturbative bounds

THEORETICAL CONSISTENCY BOUNDS

J.F., S. Garcia Saenz,
L. Pinol, S. Renaux-
Petel, J. Ronayne '19

$$\langle \zeta^m \rangle = \langle 0 | \overline{T} \left(e^{i \int \mathcal{H} dt} \right) \zeta_{\mathbb{I}} T \left(e^{-i \int \mathcal{H} dt} \right) | 0 \rangle$$

$$\supset \text{Diagram} \sim (\mathcal{H}^{(3)})^{m-2}$$

THEORETICAL CONSISTENCY BOUNDS

J.F., S. Garcia Saenz,
L. Pinol, S. Renaux-
Petel, J. Ronayne '19

$$\langle \zeta^n \rangle = \langle 0 | \overline{T} \left(e^{i \int H dt} \right) \zeta_I T \left(e^{-i \int H dt} \right) | 0 \rangle$$

$$\supset \text{[Feynman diagram: a tree with } n \text{ external legs]} \sim (H^{(3)})^{n-2}$$

ENHANCEMENT OF THE POWER SPECTRUM

$$\Rightarrow f_{NL}^{n-2} \sim \frac{\langle \zeta^n \rangle}{\langle \zeta^2 \rangle^{n-1}} = \left[\frac{P_k}{P_0} \times \left(1 + \frac{1}{|c_s|^2} \right) \right]^{n-2}$$

PERTURBATIVITY VIOLATION Shandera et al. '08 '10 Bartolo et al '10

IF

$$f_{NL}^{n-2} \times P_k^{\frac{n-2}{2}} < 1 \Rightarrow$$

$$\frac{P_k}{P_0} \left(1 + \frac{1}{|c_s|^2} \right) \gtrsim 10^9$$

MODEL
INDEPENDENT
THEORETICAL
BOUND

HYPERINFLATION $\sim 10^{15}$

THEORETICAL CONSISTENCY BOUNDS

J.F., S. Garcia Saenz,
L. Pinol, S. Renaux-
Petel, J. Ronayne '19

$$\langle \zeta^m \rangle = \langle 0 | \overline{T} \left(e^{i \int H dt} \right) \zeta_I T \left(e^{-i \int H dt} \right) | 0 \rangle$$

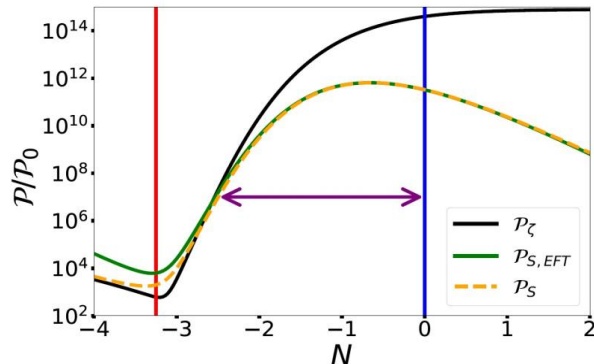
$$\supset \text{[Feynman diagram: a tree with m external legs]} \sim (H^{(3)})^{m-2}$$

ENHANCEMENT OF THE POWER SPECTRUM

$$\Rightarrow f_{NL}^{m-2} \sim \frac{\langle \zeta^m \rangle}{\langle \zeta^2 \rangle^{m-1}} = \left[\frac{P_k}{P_0} \times \left(1 + \frac{1}{|c_s|^2} \right) \right]^{m-2}$$

PERTURBATIVITY VIOLATION

Shandera et al. '08 '10 Bartolo et al '10



$$\frac{P_k}{P_0} \left(1 + \frac{1}{|c_s|^2} \right) \gtrsim 10^9$$

MODEL INDEPENDENT THEORETICAL BOUND

HYPERINFLATION $\sim 10^{15}$

CONCLUSIONS

- Multi-field inflation with *internal geometry* naturally arises in *UV complete theories*.
- *Large bending* may be required to satisfy *the refined de Sitter Swampland conjecture*.
- Perturbative analysis reveals a *transient tachyonic instability* EFT with $C_S^2 < 0$.
 - Exponential enhancement of P_ζ
 - Large flattened bispectrum f_{NL}^{flat} Match with numerics
 - Stringent bound on the enhancement of P_ζ
 - E.g. Hyperinflation is ruled out

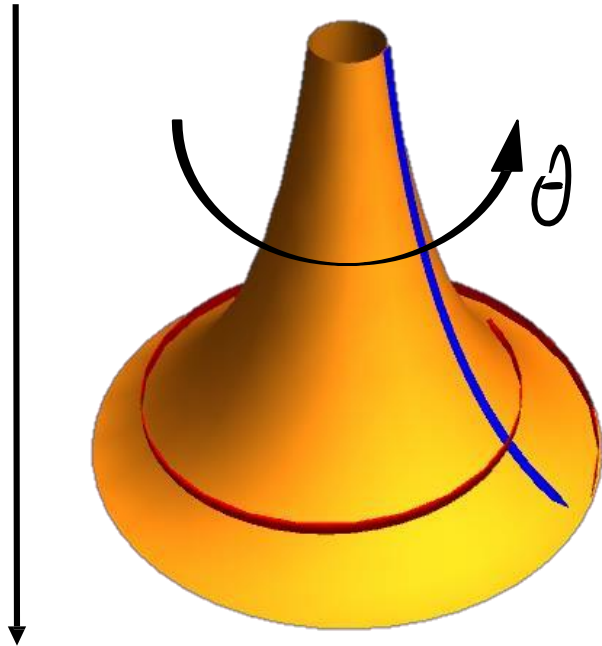
Example: HYPERINFLATION A. Brown '17

Inflating on a hyperbolic geometry $R_{\text{eff}} = -2/\ell^2$

$$G_{ij} d\phi^i d\phi^j = a^2 \left(d\phi^2 + \ell^2 \frac{\sin^2(\phi)}{\ell} d\theta^2 \right) \quad ; \quad V = V(\phi)$$

Conserved quantity

$$J = a^3 \cdot (\text{ANGULAR MOMENTUM})$$



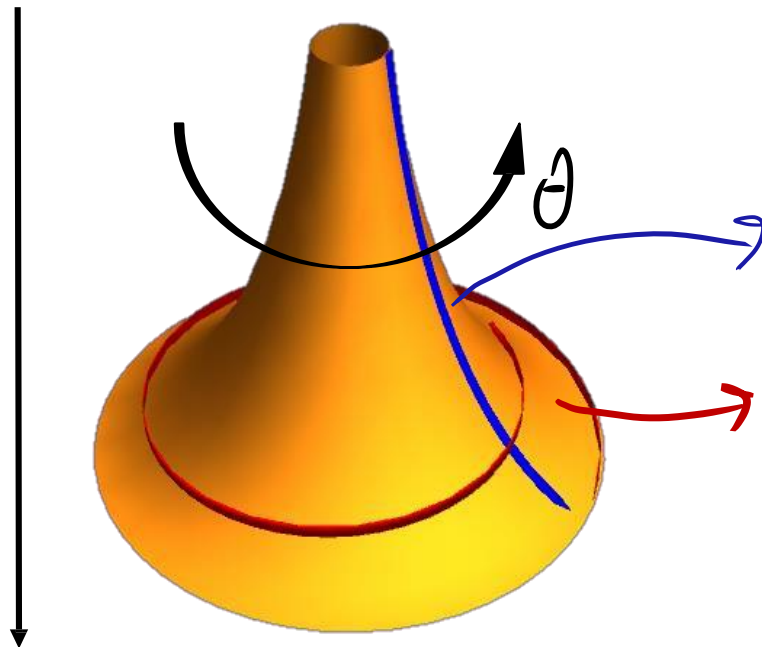
Example: HYPERINFLATION A. Brown '17

Inflating on a hyperbolic geometry $R_{\text{eff}} = -2/M^2$

$$G_{IJ} d\phi^I d\phi^J = d\phi^2 + M^2 \sin^2(\frac{\phi}{M}) d\theta^2 \quad ; \quad V = V(\phi)$$

Conserved quantity

$$J = a^3 \cdot (\text{ANGULAR MOMENTUM}) \dot{\theta}$$



- Radial geodesic $J=0$
unstable for steep potential
- New attractor: spiraling $J \neq 0$

$$\frac{3\pi}{M_{\text{Pl}}} < M_{\text{Pl}} \frac{V'}{V} \ll \frac{M_{\text{Pl}}}{M}, \quad M \frac{|V''|}{V'} \ll 1$$

\longrightarrow Inflate without slow-rolling & large bending
 e.g. $V = \frac{m^2 \phi^2}{2}$, $M = 10^{-2} M_{\text{Pl}} \implies \eta_{\perp}^2 \sim 100, \quad \frac{m_s^2}{H^2} \sim -200 \implies C_s^2 \approx -1$

MATCHING THE EFT

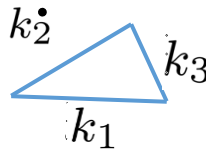
NUMERICS

PyTransport 2.0

D. Mulryne and J. Rouayne '16

Full Bispectrum $\langle \gamma_{n_1} \gamma_{n_2} \gamma_{n_3} \rangle$

MATCHING



EFT

$C_{\xi}^2 \approx -1$

$$\frac{\langle \gamma^3 \rangle}{\langle \gamma^2 \rangle^2} \sim f_{NL}(A, x)$$

MATCHING THE EFT

J.F., S. Garcia Saenz,
L. Pinol, S. Renaux-
Petel, J. Ronayne '19

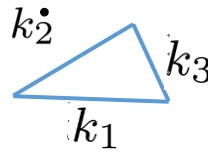
NUMERICS

PyTransport 2.0

D. Mulryne and J. Ronayne '16

Full Bispectrum $\langle \zeta_{n_1} \zeta_{n_2} \zeta_{n_3} \rangle$

MATCHING

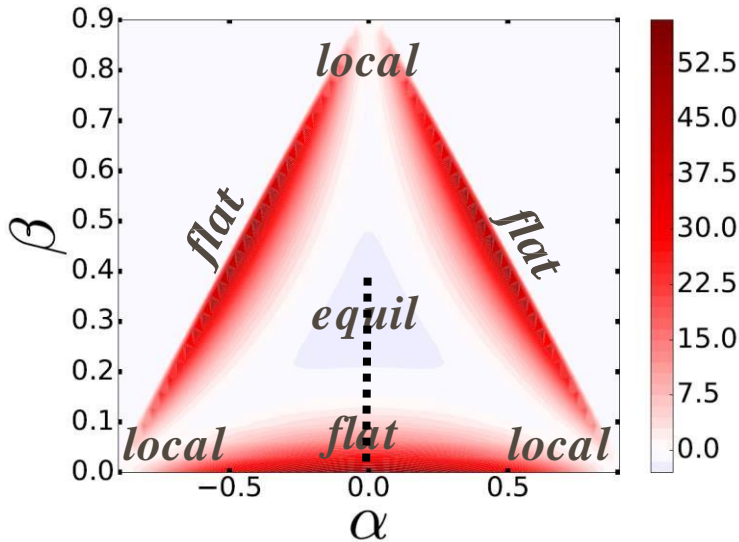


EFT

$$C_{\zeta}^2 \approx -1$$

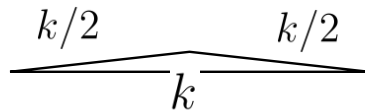
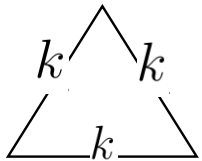
$$\frac{\langle \zeta^3 \rangle}{\langle \zeta^2 \rangle^2} \sim f_{NL}(A, x)$$

~ 10.5
 ~ -0.3



$$f_{NL}^{eq} = -2.0,$$

$$f_{NL}^{flat} = 53.8$$



MATCHING THE EFT

J.F., S. Garcia Saenz,
L. Pinol, S. Renaux-
Petel, J. Ronayne '19

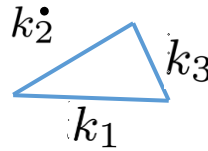
NUMERICS

PyTransport 2.0

D. Mulryne and J. Ronayne '16

Full Bispectrum $\langle \gamma_{n_1} \gamma_{n_2} \gamma_{n_3} \rangle$

MATCHING



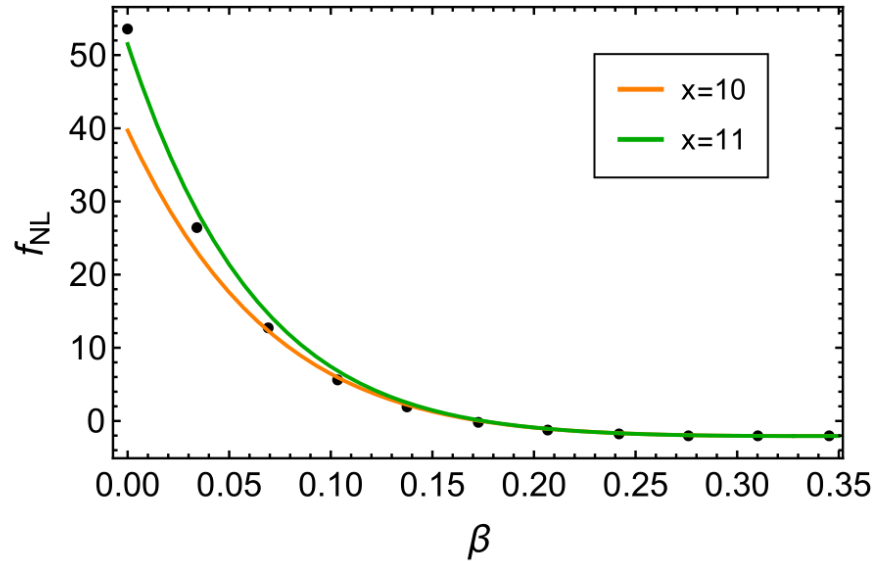
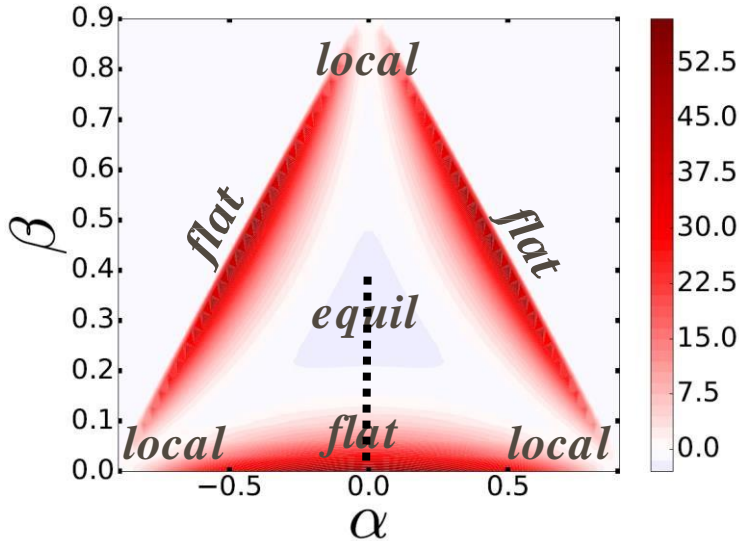
EFT

$$C_{\xi}^2 \approx -1$$

$$\frac{\langle \gamma^3 \rangle}{\langle \gamma^2 \rangle^2} \sim f_{NL}(A, x)$$

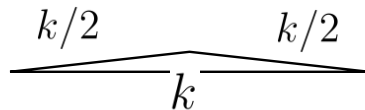
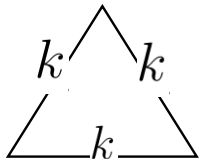
~ 10.5

~ -0.3



$$f_{NL}^{eq} = -2.0,$$

$$f_{NL}^{flat} = 53.8$$



MATCHING THE EFT

J.F., S. Garcia Saenz,
L. Pinol, S. Renaux-
Petel, J. Ronayne '19

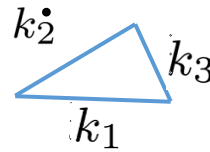
NUMERICS

PyTransport 2.0

D. Mulryne and J. Ronayne '16

Full Bispectrum $\langle \gamma_{n_1} \gamma_{n_2} \gamma_{n_3} \rangle$

MATCHING



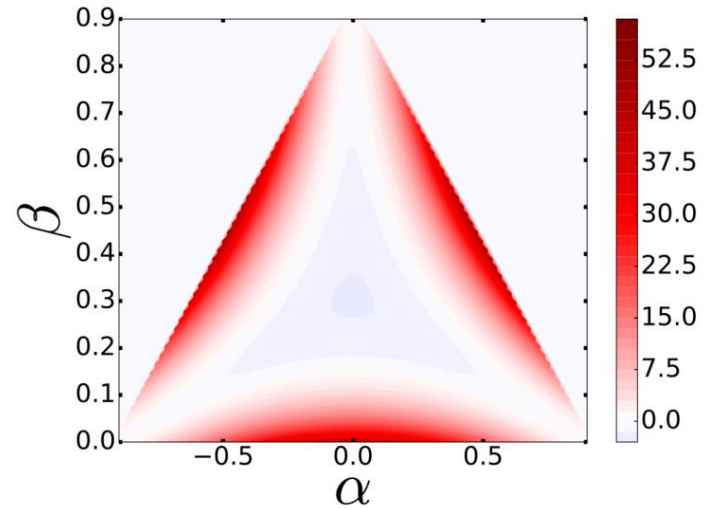
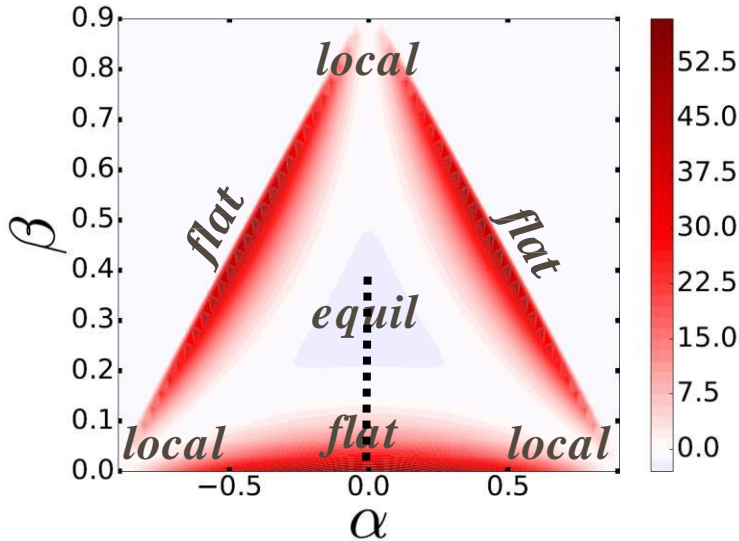
EFT

$$C_{\xi}^2 \approx -1$$

$$\frac{\langle \gamma^3 \rangle}{\langle \gamma^2 \rangle^2} \sim f_{NL}(A, x)$$

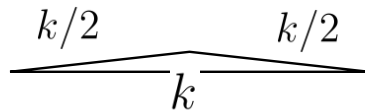
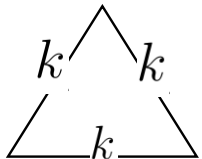
~ 10.5

~ -0.3



$$f_{NL}^{eq} = -2.0,$$

$$f_{NL}^{flat} = 53.8$$



THEORETICAL CONSISTENCY BOUNDS

$$\langle \zeta^m \rangle = \langle 0 | \bar{T}(e^{i \int H dt}) \zeta_I T(e^{-i \int H dt}) | 0 \rangle$$

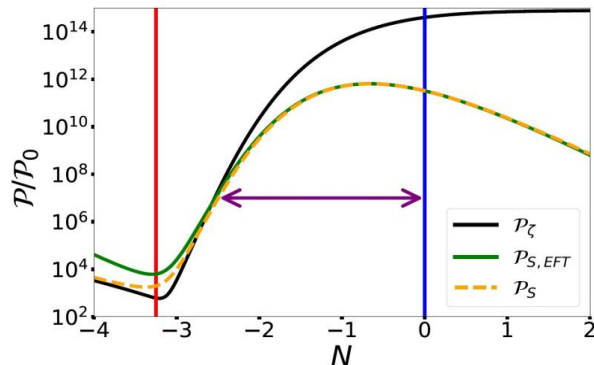
$$\supset \text{Feynman diagram} \sim (\mathcal{H}^{(3)})^{m-2}$$

ENHANCEMENT OF THE POWER SPECTRUM

$$\Rightarrow f_{NL}^{m-2} \sim \frac{\langle \zeta^m \rangle}{\langle \zeta^2 \rangle^{m-1}} = \left[\frac{P_k}{P_0} \times \left(1 + \frac{1}{|c_s|^2} \right) \right]^{m-2}$$

PERTURBATIVITY VIOLATION

Shandera et al. '08 '10 Bartolo et al '10



$$\frac{P_k}{P_0} \left(1 + \frac{1}{|c_s|^2} \right) \gtrsim 10^9$$

MODEL INDEPENDENT THEORETICAL BOUND

HYPERINFLATION $\sim 10^{15}$

CONCLUSIONS

- Multi-field inflation with *internal geometry* naturally arises in *UV complete theories*.
- *Large bending* may be required to satisfy *the refined de Sitter Swampland conjecture*.
- Perturbative analysis reveals a *transient tachyonic instability* EFT with $C_S^2 < 0$.
 - Exponential enhancement of P_ζ
 - Large flattened bispectrum f_{NL}^{flat} Match with numerics
 - Stringent bound on the enhancement of P_ζ
 - E.g. Hyperinflation is ruled out