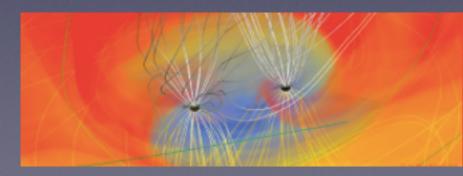
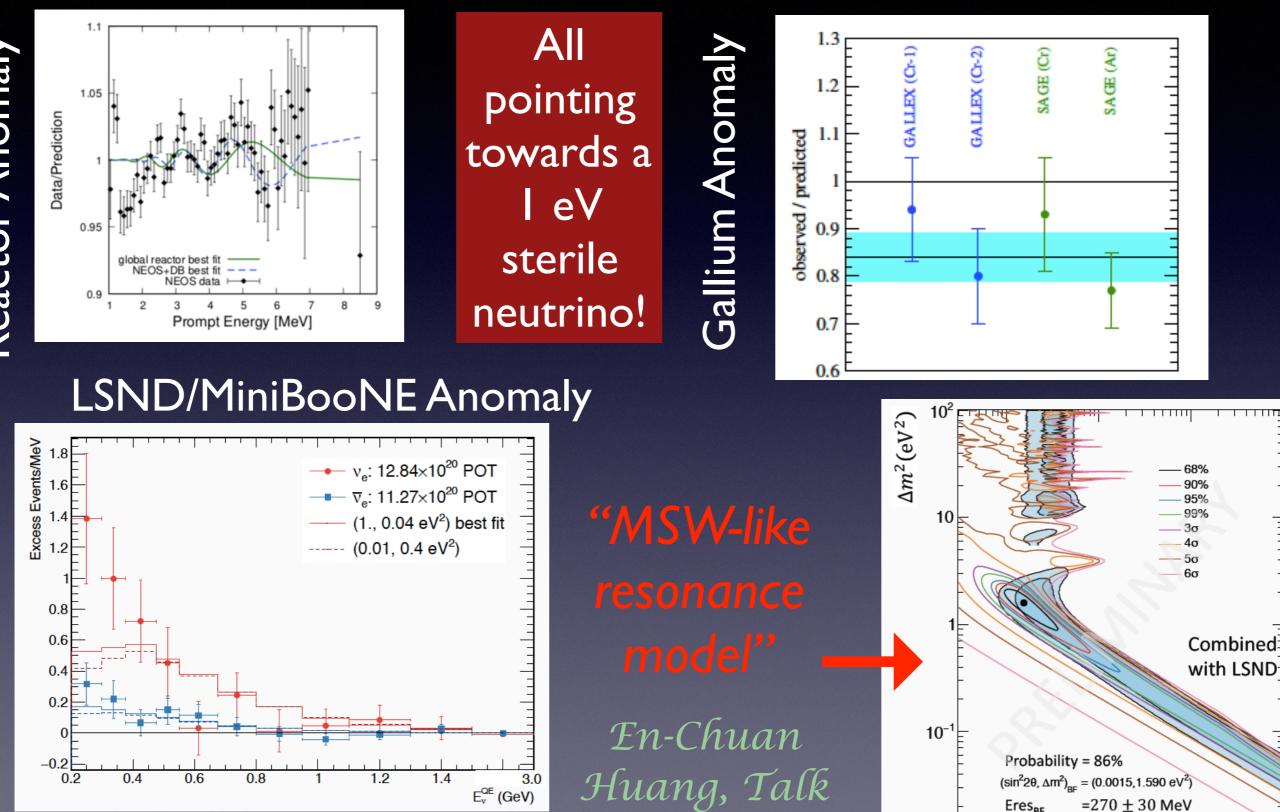
Sterile Neutrinos with Altered Dispersion Relations as an Explanation for MiniBooNE, LSND, Gallium and Reactor Anomalies

- HEINRICH PÄS tu dortmund



PASCOS 2019 XXV International Symposium





at Netríno2018

10⁻²

 10^{-3}

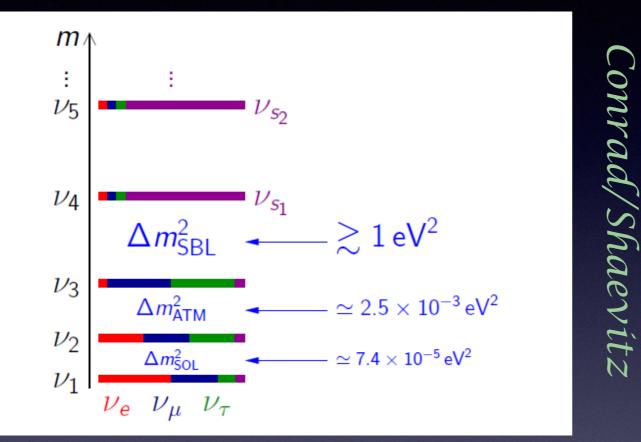
 10^{-2}

1 1 1 1 1 1 1 1 1

10-1

 $sin^2 2\theta$

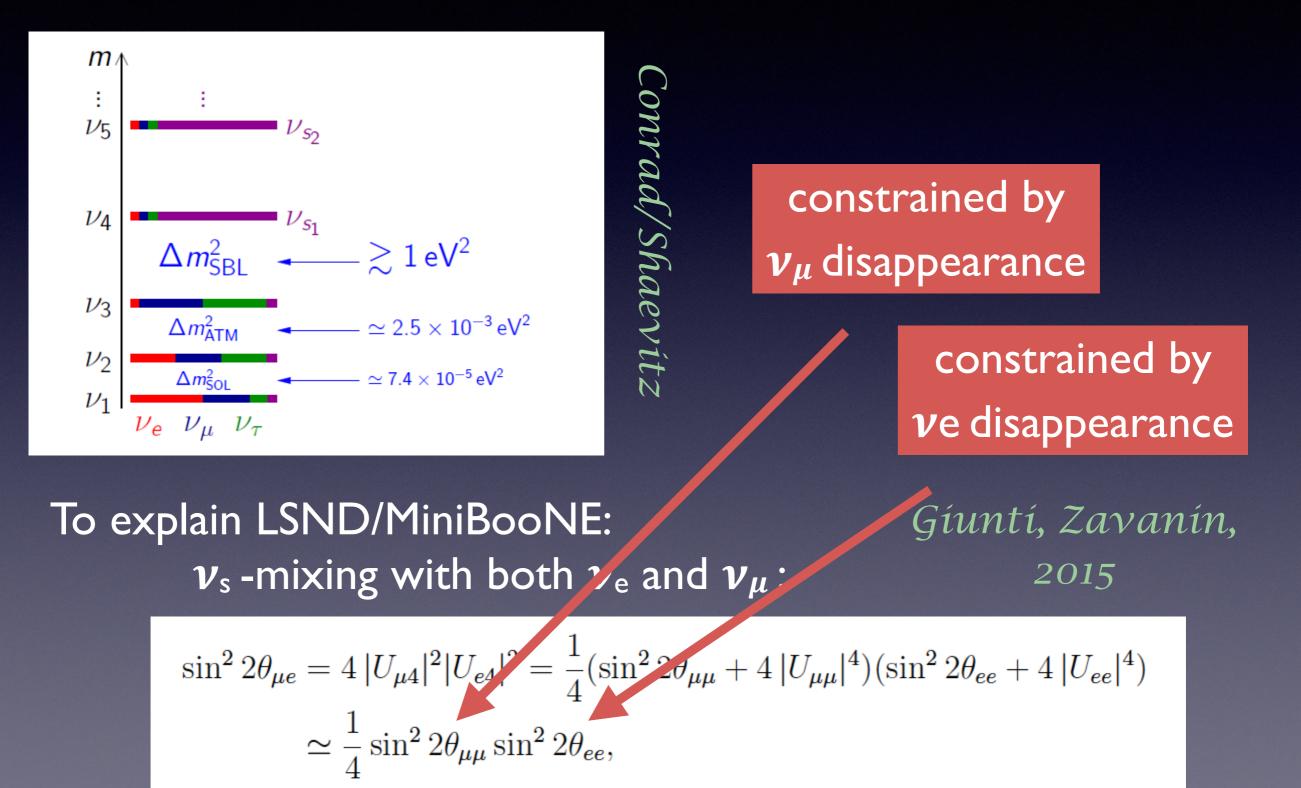
\rightarrow Add a sterile neutrino with $\Delta m^2 \sim 1 eV^2$



To explain LSND/MiniBooNE: v_s -mixing with both v_e and v_{μ} :

$$\sin^2 2\theta_{\mu e} = 4 |U_{\mu 4}|^2 |U_{e4}|^2 = \frac{1}{4} (\sin^2 2\theta_{\mu \mu} + 4 |U_{\mu \mu}|^4) (\sin^2 2\theta_{ee} + 4 |U_{ee}|^4)$$
$$\simeq \frac{1}{4} \sin^2 2\theta_{\mu \mu} \sin^2 2\theta_{ee},$$

\rightarrow Add a sterile neutrino with $\Delta m^2 \sim 1 eV^2$



A 4th sterile neutrino ?

Sterile netrinos can have a very different origin than the SM neutrinos (e.g. superpartners of dilaton, radion or other moduli fields, mirror world fermions, etc...)

> There is NO compelling reason to believe that a sterile neutrino should behave just as the SM neutrinos

Sterile neutrino Altered Dispersion Relations

Evidence for light sterile v is partly conflicting!

- May be wrong!
- May hint towards deviations from the usual oscillation mechanism!
- Sterile neutrinos as messengers of exciting new physics?

Attractive candidate: Altered dispersion relations

$$E = p + m^2/2E + new terms$$

→ novel energy dependence!

HP, Pakvasa, Weiler 2005

Evolution equation in flavor space:

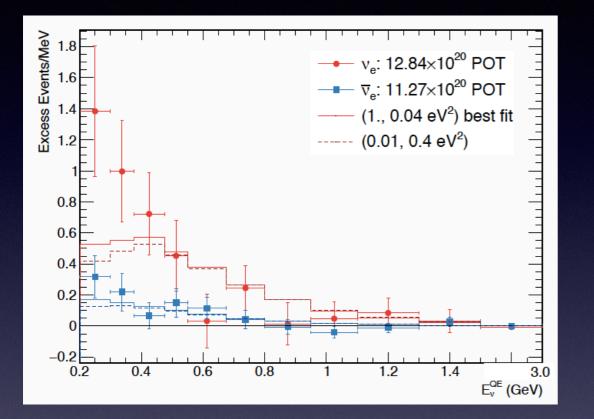
$$i\frac{d}{dt}\left(\begin{array}{c}\nu_a(t)\\\nu_s(t)\end{array}\right) = H_F\left(\begin{array}{c}\nu_a(t)\\\nu_s(t)\end{array}\right)$$

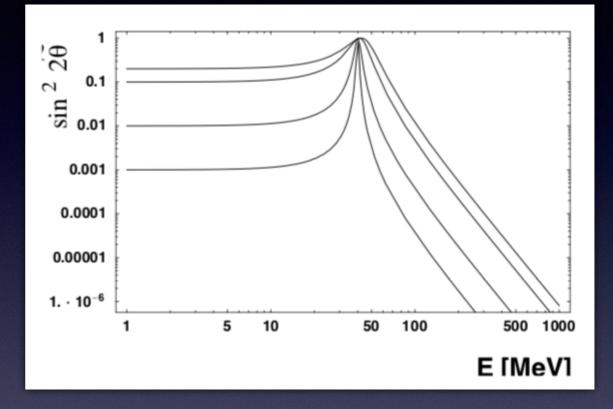
Hamiltonian in the presence of bulk shortcuts:

$$H_F = \pm \frac{\delta m^2}{4E} \begin{pmatrix} \cos 2\theta & -\sin 2\theta \\ -\sin 2\theta & -\cos 2\theta \end{pmatrix} + E \frac{\epsilon}{2} \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}$$

 \Rightarrow A Resonance exists at $E_{\rm res} = \sqrt{\frac{\delta m^2 \cos 2\theta}{2 \epsilon}}$

 \rightarrow choose $E_{\rm res}$ =30-400 MeV $\leftrightarrow \epsilon \simeq 10^{-18} - 10^{-16}$ (Päs, Pakvasa, Weiler, 2005)

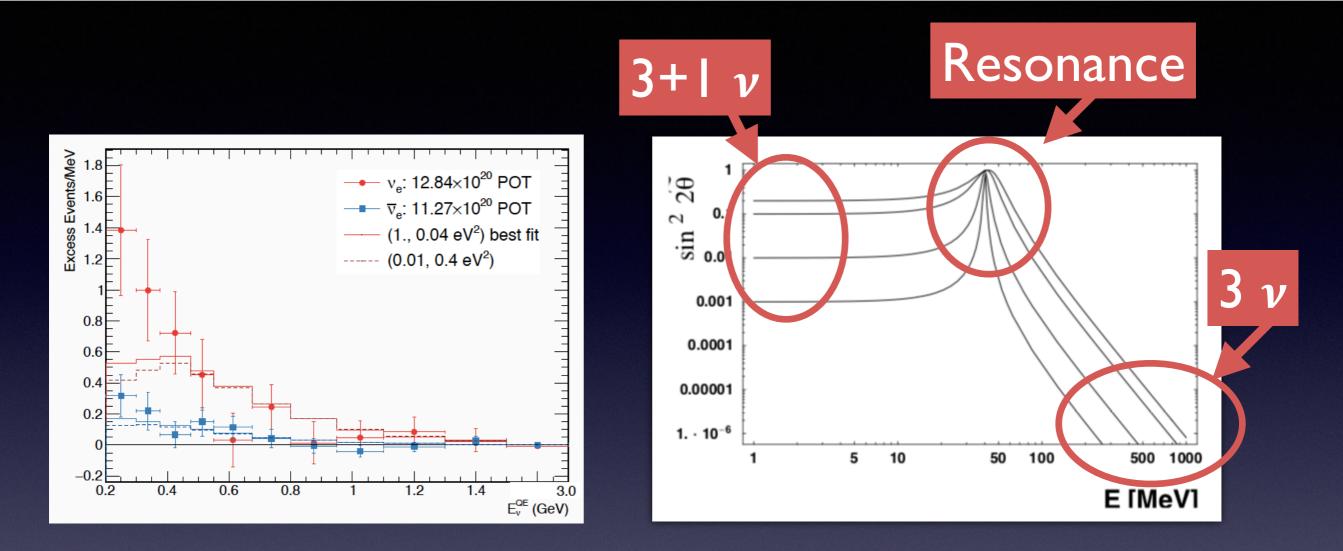




MíníB00NE 2018

HP, Pakvasa, Weiler 2005

→ resonance-like features!



MíníB00NE 2018

HP, Pakvasa, Weiler 2005

→ resonance-like features!

[P. Huber, 2007]

"A níce model but unfortunately ít doesn't work"

P. Huber

Altered dispersion relations with 3+1 neutrinos

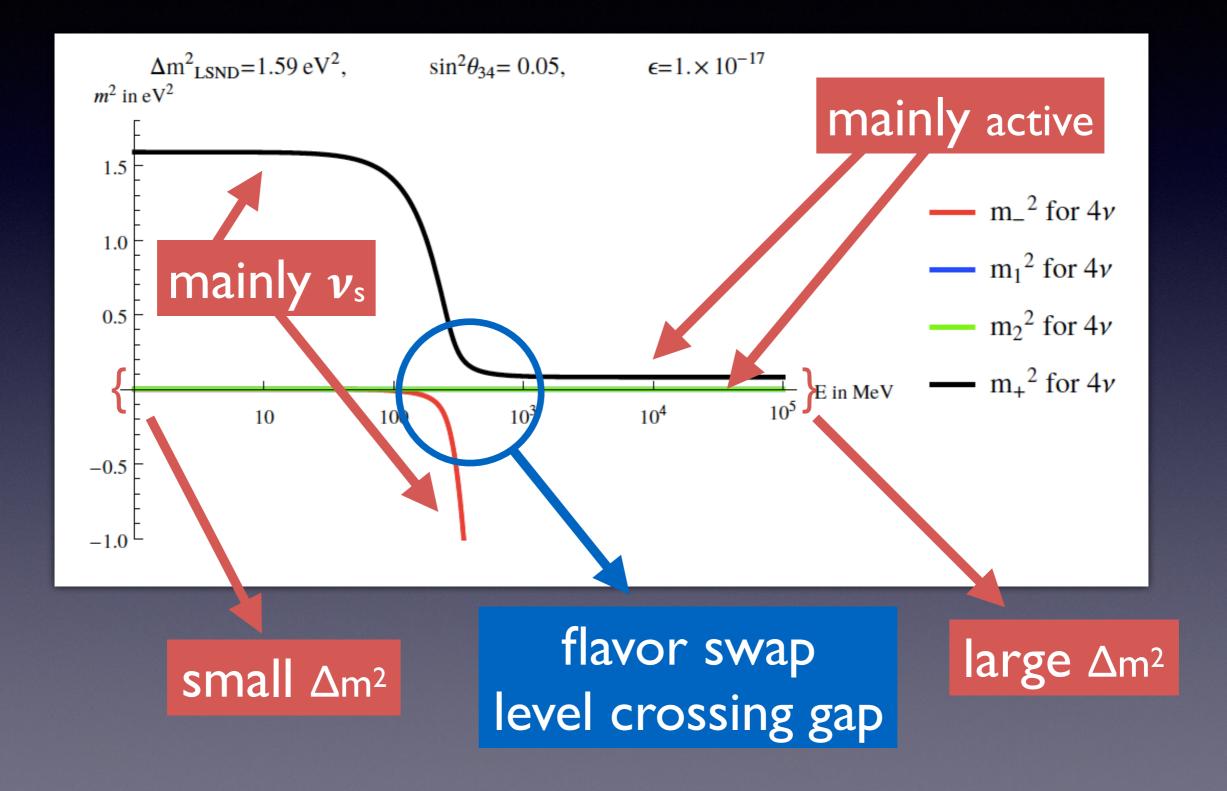
$$\lambda_{4/3} \equiv \lambda_{\pm} = \frac{\Delta}{4E} \left(1 - \cos 2\theta_{34} \left(\frac{E}{E_{\rm R}} \right)^2 \pm \sqrt{\sin^2 2\theta_{34} + \cos^2 2\theta_{34}} \left[1 - \left(\frac{E}{E_{\rm R}} \right)^2 \right]^2 \right)$$

- 2 mass eigenstates with large effective masses!
- Large Δm^2 's [Marfatía, HP, Pakvasa, Weiler, 2012]
- Fast oscillations:

$$P(\nu_a \to \nu_b) = 4 V_{a3}^2 V_{b3}^2 \times \begin{cases} -\sin^2 \left(\frac{L(\lambda_+ - \lambda_-)}{2}\right) & \sin^2 \tilde{\theta} \cos^2 \tilde{\theta} \\ +\sin^2 \left(\frac{L\lambda_+}{2}\right) & \sin^2 \tilde{\theta} \\ +\sin^2 \left(\frac{L\lambda_-}{2}\right) & \cos^2 \tilde{\theta} . \end{cases}$$

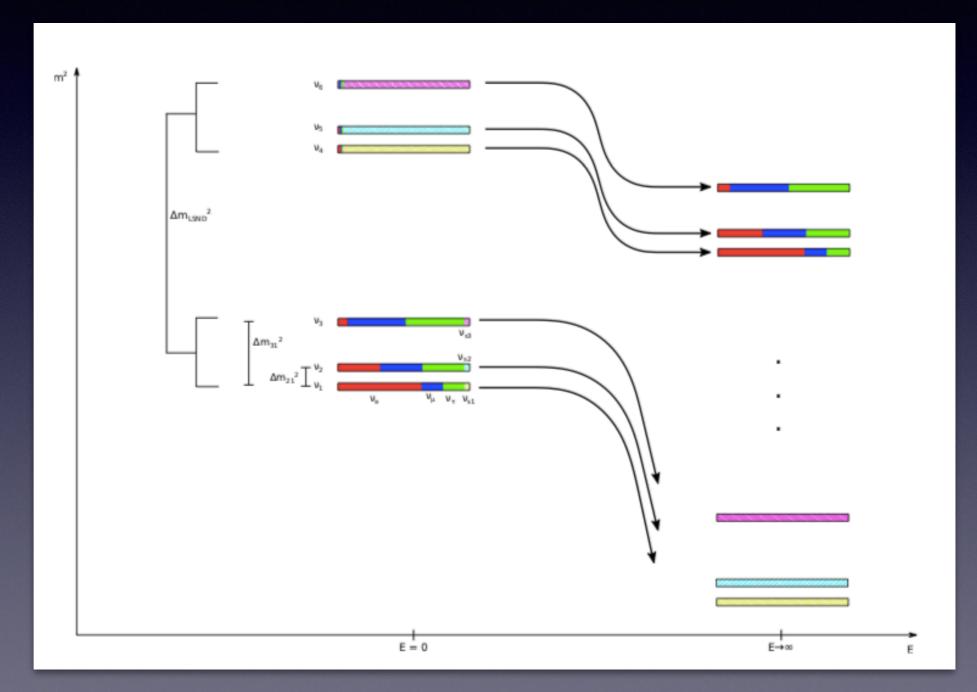
Sterile neutrino Lorentz violation excluded by atmospheric v L/E !!! [P. Huber, 2007]

Altered dispersion relations with 3+1 neutrinos

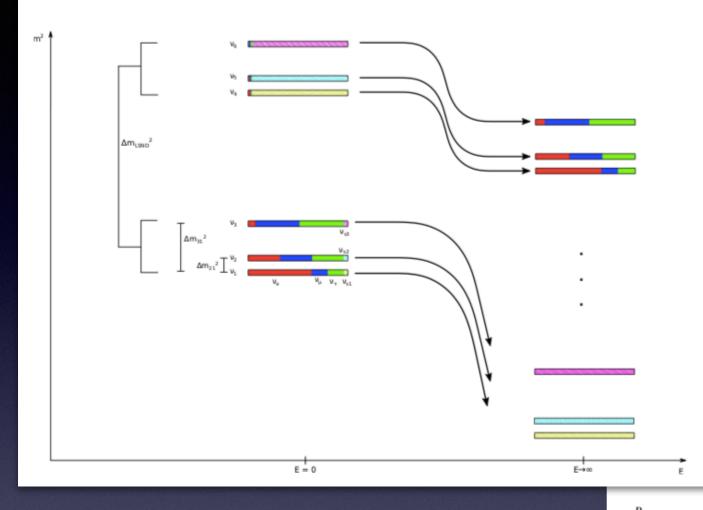


Solution ?

\rightarrow Promote the model to 3+3

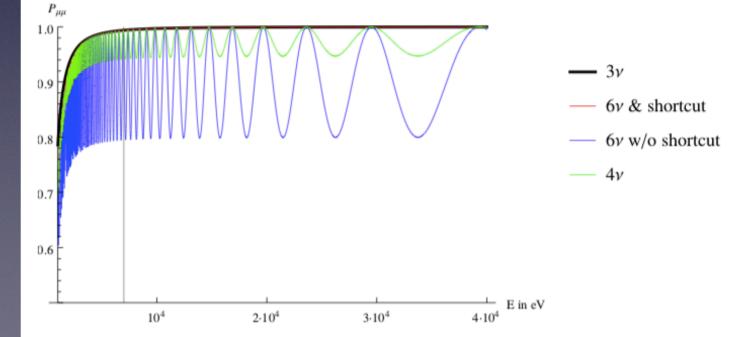


Döríng, HP, Síckíng, Weíler, arXív:1808.07460



In a 3+3 model with democratic mixing the sterile neutrino decouples completely at high energies

Probabilities @ MINOS FarDetector $\Delta m^2_{LSND}=1.59 \text{ eV}^2, \text{ sin}^2\theta = 0.0528,$ $\epsilon = 5. \times 10^{-17}, \kappa = \xi = 100 = 5. \times 10^{-17}, \text{ L} = 735 \text{ km}$



Döríng, HP, Síckíng, Weíler, arXív:1808.07460

Altered dispersion relations with 3+3 neutrinos

Necessary: Different resonance energies for 3 sterile neutrinos to avoid unitarity cancellation!

Döríng, HP, Síckíng, Weíler, arXív: 1808.07460

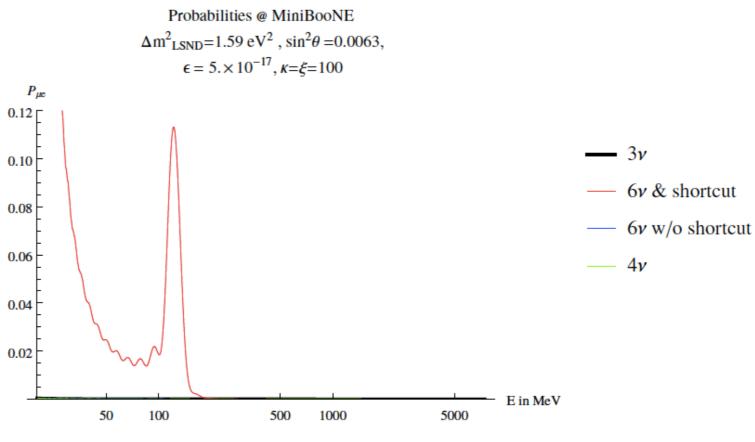
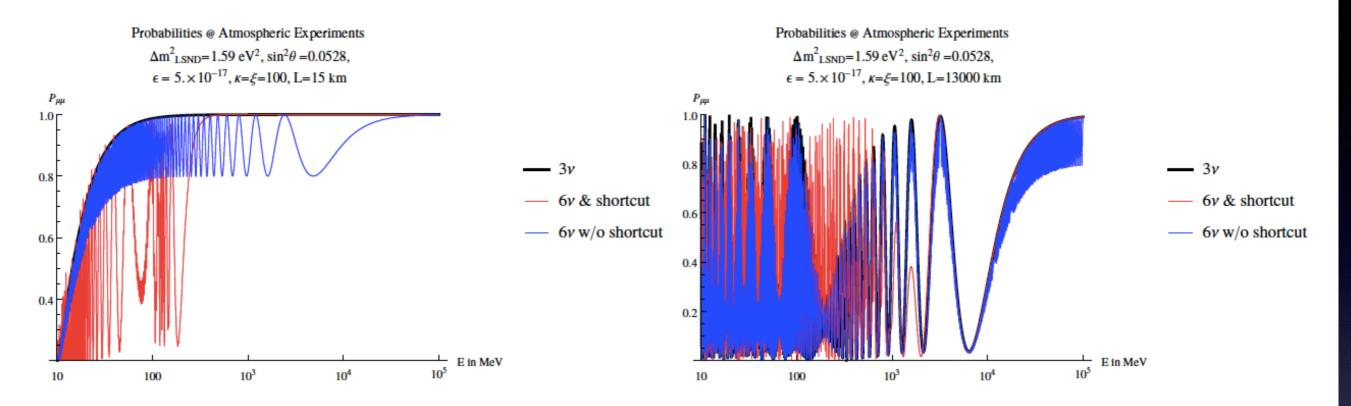


Figure: Probabilities *P* at MiniBooNE in the $\nu_{\mu} \rightarrow \nu_{e}$ channel.

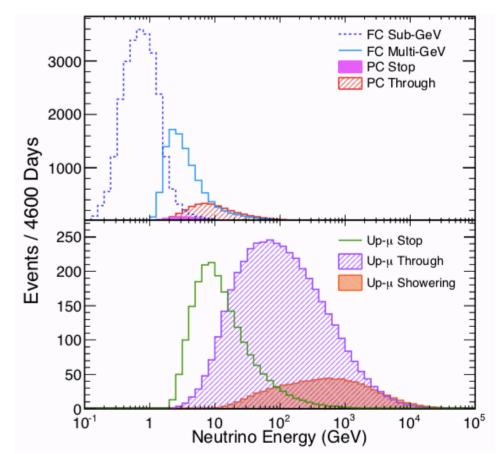
sub-GeV Super-K data?



(a)Downward going neutrinos

(b)Upward going neutrinos

Döríng, HP, Síckíng, Weiler, arXív:1808.07460



Model Building ?

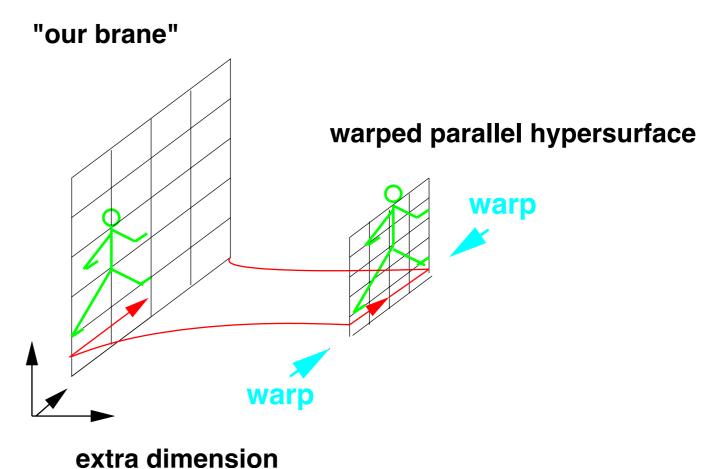
Consider an asymmetrically warped extra dimension with a sterile neutrino in the bulk

$$ds^2 = dt^2 - e^{-2k|u|} dx^2 - du^2$$

Chung, Freese, 1999, 2000 Csakí, Erlích, Grojean, 2001

- Shrinks space parallel to the brane
- Allows for shortcuts in the extra dimension!

[Döríng, HP, arXív:1808.07734



$$p_{\mathbf{x}/\boldsymbol{\xi}} \approx E - \frac{\kappa^2}{2E} + \frac{E^2}{2E}\tilde{I}_{00} + \mathcal{O}(\tilde{I}_{00}^2) + \mathcal{O}(\kappa^2\tilde{I}_{00})$$

Summary

- Chances are high that MiniBooNE and other anomalies due to backgrounds or systematics
- Partially in conflict with other bounds
- But keep in mind: "1 eV² miracle"
- ADR v oscillations: energy dependence & resonant conversion
- ADR 3+3 v model seems to explain all data !
- 4D EFT limit of v mass models in asymmetrically warped extra dimensions
- Potential to alleviate cosmological bounds