

Chaotic Inflation and Fractional Powers

The Dynamical Origin of the Inflaton Potential in Chaotic Inflation ■



Kai Schmitz

Kavli Institute for the Physics and Mathematics of the Universe (WPI)

Todai Institutes for Advanced Study, University of Tokyo, Kashiwa, Japan

Based on arXiv:1211.6241, 1403.4536, 1407.3084 [hep-ph].

In collaboration with Keisuke Harigaya, Masahiro Ibe and T. T. Yanagida.

SUSY 2014 | University of Manchester, UK | July 21, 2014

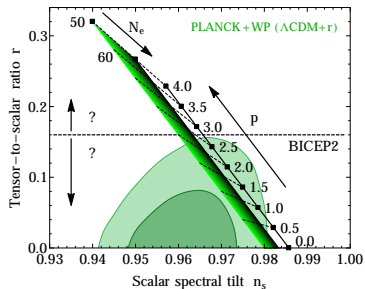
Outline

- 1 A Fractional Power-Law Potential for Chaotic Inflation
- 2 Explicit Realizations in Models of Dynamical SUSY Breaking
- 3 Embedding into Supergravity and Phenomenology
- 4 Conclusions and Outlook

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The Case for Dynamical Chaotic Inflation (DCI)



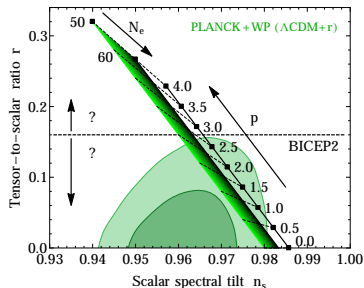
Assume BICEP2 claim survives further scrutiny:

- ▶ Large field excursion. → Chaotic inflation based on a simple monomial potential.

$$V(\phi) \sim M^4 \left(\frac{|\phi|}{M} \right)^p, \quad p \in \mathbb{Q}^+.$$

- ▶ Perturbative QFT: Only $p \in \mathbb{N}^+$ feasible.

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Why such a potential? What determines the power p and the mass scale M ?

- ▶ Axion monodromy in *string theory*:

[Silverstein & Westphal '08] [McAllister, Silverstein & Westphal '08]

$$p = \frac{2}{5}, \frac{2}{3}, 1, 2.$$

- ▶ Fractional powers from first principles!

- ▶ Embedding into *SUGRA* for $p \in \mathbb{N}^+$:

[Kawasaki, Yamaguchi & Yanagida '00] [Kallosh & Linde '10]

$$W = X f(\Phi), \quad \Phi \rightarrow \Phi + i\alpha.$$

- ▶ Only educated guesses of W and K .

Generate fractional power-law potential for chaotic inflation *dynamically* within field theory!

DCI in Strongly Interacting SUSY Gauge Theories (I)

Theoretical ingredients: [Seiberg '94] [Intriligator & Pouliot '95] [Csaki, Schmaltz & Skiba '97]

- 1 Supersymmetry:** Otherwise no control over IR dynamics, quadratic divergences, ...
 - 2 Dynamical SUSY breaking (DSB):** Vacuum energy $V(\phi)$ acts as inflaton potential.
 - 3 S-confinement:** Smooth effective field theory in terms of composites at low energies.
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Dynamical generation of the inflaton potential (in words):

- ▶ In a given DSB model, provide some quark flavors with **inflaton-dependent mass**:

$$W \supset (\lambda \Phi + M_a) P^a \bar{P}^a, \quad M_a \geq 0.$$

- ▶ For $\lambda \Phi \gtrsim \Lambda$, these quark flavors decouple perturbatively, so that at low energies:

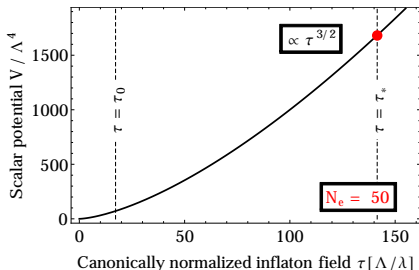
$$\Lambda_{\text{eff}} = \Lambda \left(\frac{\lambda \Phi}{\Lambda} \right)^{p/4}, \quad \frac{p}{4} = \frac{b_{\text{eff}} - b}{b_{\text{eff}}} \Rightarrow W_{\text{eff}} \simeq \Lambda_{\text{eff}}^2 X, \quad V_{\text{eff}} \simeq \Lambda^4 \left(\frac{\lambda \Phi}{\Lambda} \right)^p.$$

- ▶ For $\lambda \Phi \lesssim \Lambda$, s-confined phase with all fields being stabilized around the origin:

$$V \simeq \Lambda^2 |\tilde{\phi}|^2, \quad \tilde{\phi} = f(\phi).$$

Fractional power p . ✓ M identified as Λ . No input scale. ✓ Smooth around $\Phi = 0$. ✓

DCI in Strongly Interacting SUSY Gauge Theories (II)

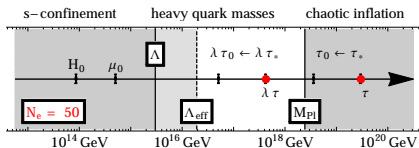
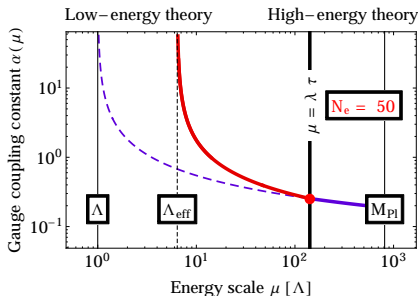
Example: $p = 3/2$ [based on $SP(3)$ theory]

RGE matching at quark mass threshold:

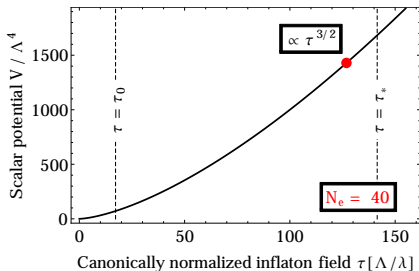
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Energy scales involved in DCI:

Running of the gauge coupling $\alpha = \frac{g^2}{4\pi}$:

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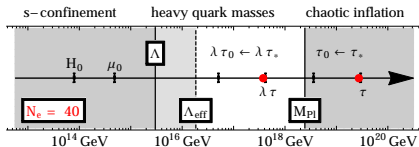
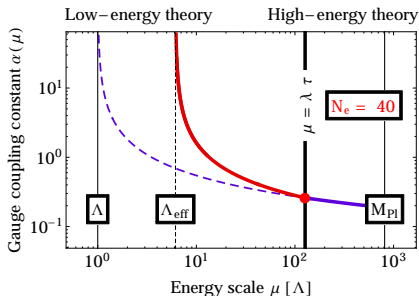
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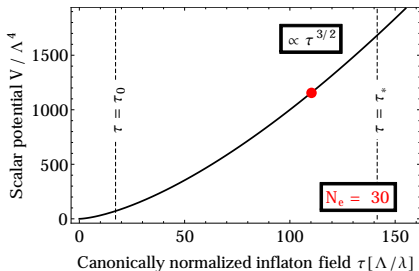
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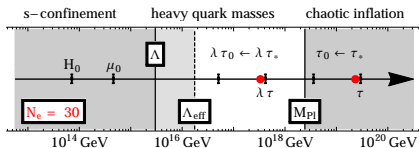
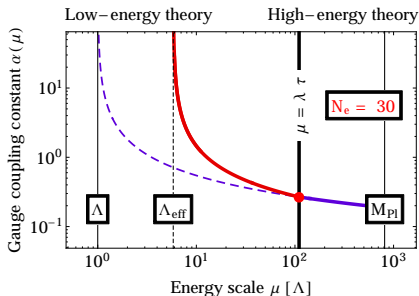
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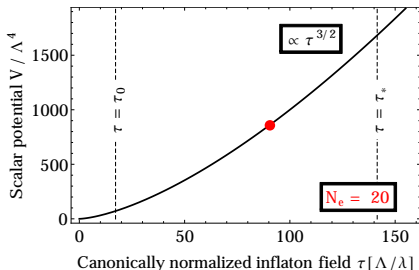
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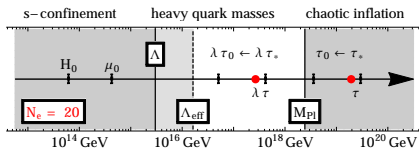
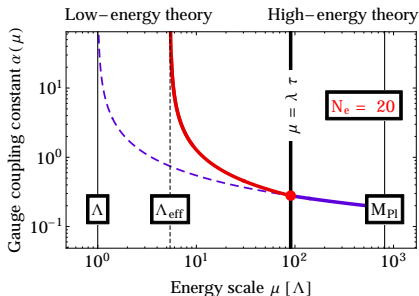
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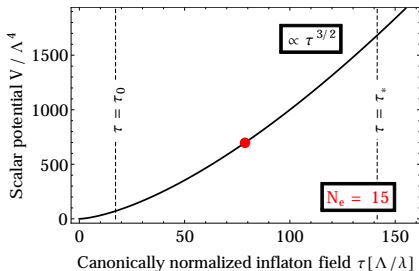
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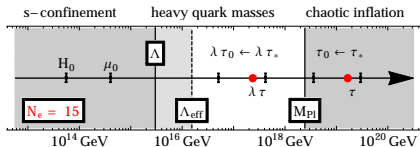
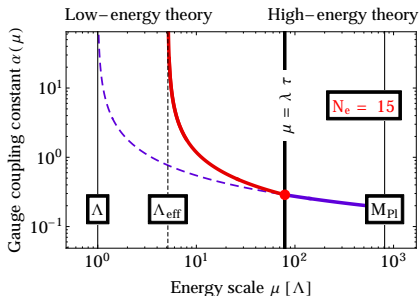
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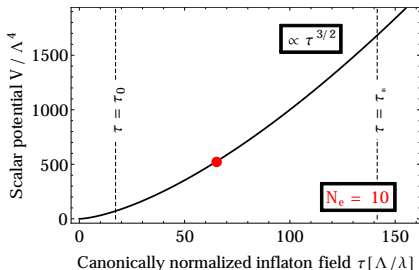
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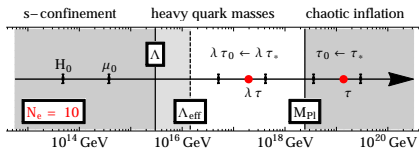
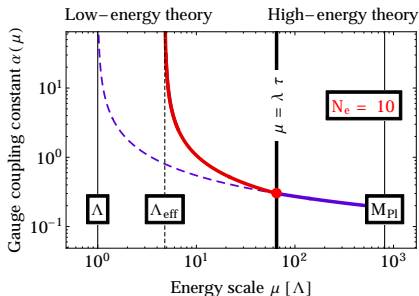
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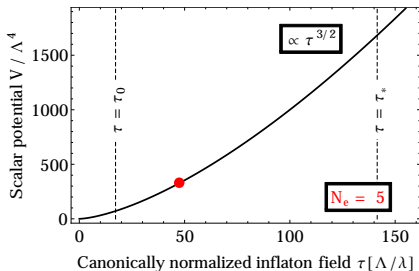
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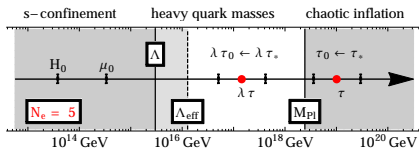
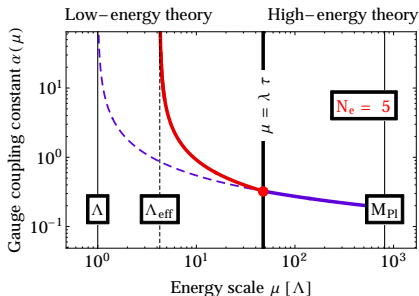
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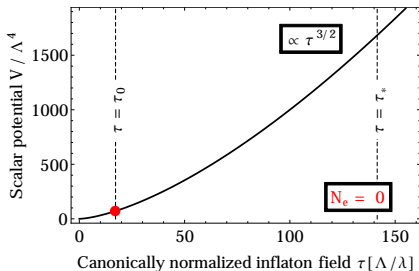
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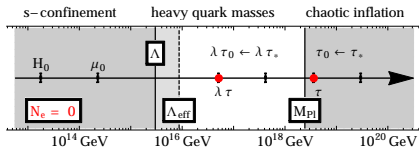
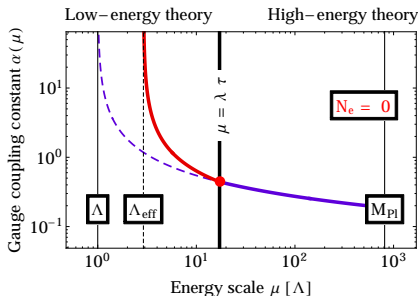
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Minimal Scenario: $SP(N_c)$ Dynamics

Recall ingredient ②: DSB responsible for $V(\phi) \simeq \Lambda_{\text{eff}}^4$ during inflation.

- ▶ Need DSB model that flows to s-confining theory when $\phi \rightarrow 0$.
 - ▶ Or alternatively: s-confining theory flowing to DSB model when $\phi \gg \Lambda$.
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- ▶ Or alternatively: s-confining theory flowing to DSB model when $\phi \gg \Lambda$.

Simplest example: $SP(N_c)$ gauge theory with $2N_f$ quarks and $N_f = N_c + 2$.

- ▶ For every flat direction in moduli space, introduce one singlet field Z_{IJ} :

$$W = \lambda_{IJ} Z_{IJ} Q^I Q^J \quad \rightarrow \quad W = \lambda_{ij} Z_{ij} Q^i Q^j + \lambda \Phi P \bar{P} + \dots$$

- ▶ Identify one of the singlets as inflaton field Φ , e.g. because $[\Phi]_R = 0$.
- ▶ For $\lambda \Phi \lesssim \Lambda$, s-confined phase, dynamical superpotential, SUSY vacuum at origin.
- ▶ For $\lambda \Phi \gtrsim \Lambda$, (P, \bar{P}) flavor decouples, deformed moduli constraint, SUSY broken

[Izawa & Yanagida '96] [Intriligator & Thomas '96]

$$W \simeq \lambda_{ij} \Lambda_{\text{eff}} Z_{ij} M^{ij}, \quad \text{Pf}(M) = \Lambda_{\text{eff}}^{N_c+1}, \quad M^{ij} \simeq Q^i Q^j / \Lambda_{\text{eff}}, \quad \Lambda_{\text{eff}} \simeq \Lambda \left(\frac{\lambda \Phi}{\Lambda} \right)^{p/4}.$$

Minimize potential with respect to meson fields $M^{ij} \rightarrow$ low-energy effective theory:

$$W_{\text{eff}} \simeq \Lambda_{\text{eff}} X, \quad X \propto \sum Z_{ij} \quad V_{\text{eff}}(\phi) \simeq (N_c + 1) \Lambda_{\text{eff}}^4(\phi), \quad p = \frac{4(b_{\text{eff}} - b)}{b_{\text{eff}}} = \frac{1}{2N_c + 1}$$

Generalization along Two Different Directions

- 1 Additional massive matter fields that decouple at energies $\mu \gtrsim \Lambda$.
 - 2 Vacuum energy provided by alternative models of dynamical SUSY breaking.
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1 For example, $SP(N_c)$ gauge theory with $2N_f$ quarks and $N_f = N_c + 2 + N_m$.

- ▶ To retain s-confinement, all extra flavors must decouple above the dynamical scale,

$$W \supset (\lambda \Phi + M_a) P^a \bar{P}^a, \quad a = 1, \dots, N_m, \quad \mathcal{O}(\Lambda) \lesssim M_a \lesssim \mathcal{O}(M_{\text{Pl}}).$$

- ▶ Extra contribution to the high-energy beta-function coefficient changes the power p ,

$$b = 3(N_c + 1) - (N_c + 2 + N_m), \quad b_{\text{eff}} = 3(N_c + 1) - (N_c + 1), \quad p = \frac{2(1 + N_m)}{N_c + 1}.$$

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2 Seek alternative s-confining theories that can be transformed into DSB models:

Gauge group	S-confining phase	SUSY-breaking phase	Power p
$SP(N_c)$	$Q^i, i = 1, \dots, 2(N_c + 2)$	$Q^i, i = 1, \dots, 2(N_c + 1)$	$2/(N_c + 1)$
$SO(10)$	$\mathbf{16}_{0,1}, \bar{\mathbf{16}}_1, \mathbf{10}_{1,2,3}$	$\mathbf{16}_0$	$14/11$
$SU(5)$	$\mathbf{5}_{1,\dots,4}, \mathbf{5}_{0,\dots,4}^*, \mathbf{10}$	$\mathbf{5}_0^*, \mathbf{10}$	$16/13$
$SU(3) \times SU(2)$	$q, \bar{u}, \bar{d}, \ell, U, \bar{U}, D, \bar{D}, L, \bar{L}$	$q, \bar{u}, \bar{d}, \ell$	$8/7$
$SU(N_c)$	$Q^i, \bar{Q}^j, i = 1, \dots, N_c + 1$	$Q^i, \bar{Q}^j, i = 1, \dots, N_c$	1

[Iizawa & Yanagida '96] [Intriligator & Thomas '96] [Affleck, Dine & Seiberg '84] [Affleck, Dine & Seiberg '85] [Seiberg '85]

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Constraints on Parameter Space

$V \simeq \Lambda^4 \left(\frac{\lambda \Phi}{\Lambda} \right)^p \rightarrow$ two parameters: dynamical scale Λ and inflaton Yukawa coupling λ .

Eta problem in supergravity:

- ▶ Too large η because of e^K corrections.
- ▶ **Shift symmetry** in the direction of Φ :

$$\Phi \rightarrow \Phi + icM_{\text{Pl}}, \quad c \in \mathbb{R}.$$

- ▶ $\tau \equiv \sqrt{2} \text{Im} \{ \Phi \}$ is the actual inflaton.
- ▶ Shift symmetry explicitly broken in the Kähler potential at one-loop level:

$$K_{\text{eff}} \supset \frac{\lambda^2}{16\pi^2} |\Phi|^2 \ln \left(\frac{\mu^2}{M_{\text{Pl}}^2} \right).$$

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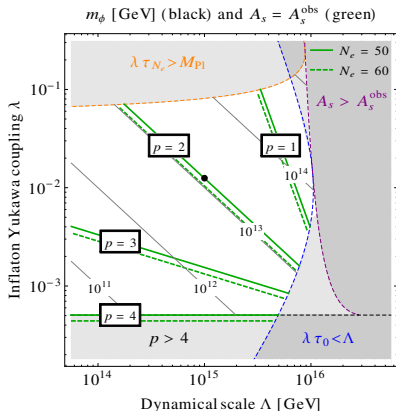
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$$\Lambda \sim \Lambda_{\text{GUT}}, \quad \checkmark \quad \lambda \sim 10^{-3} \dots 10^{-1}. \quad \checkmark$$

Normalization of the power spectrum and bounds imposed for consistency:

$$A_s \equiv A_s^{\text{obs}}, \quad \Lambda \lesssim \lambda \tau \lesssim M_{\text{Pl}}.$$

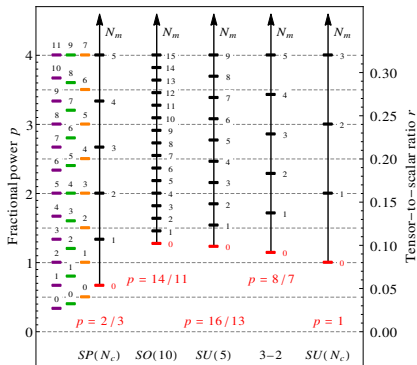


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Precise determination of p would allow to identify the dynamics of the inflaton sector!

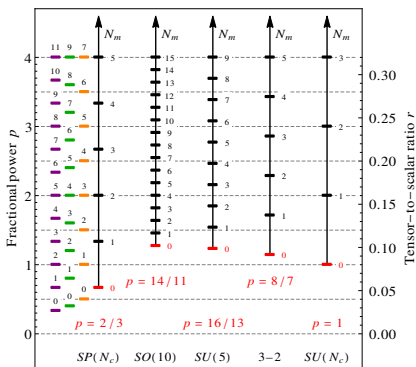


Virtues of dynamical chaotic inflation:

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- ▶ Energy scale of inflation generated via dimensional transmutation; thus, natural reason why $V^{1/4} \ll M_{\text{Pl}}$.
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Next steps towards a theory of DCI:

- ▶ Conformal window or DSB as alternatives to s-confinement.
- ▶ DSB during inflation in meta-stable vacuum or *in the conformal window*.
[Intriligator, Seiberg & Shih '06] [Yanagida et al. '09]
- ▶ Embed DCI into string theory / explain origin of the shift symmetry for Φ .

Fascinating picture: Inflation as a mere consequence of strong supersymmetric gauge dynamics shortly below the Planck scale! Calls for further exploration!

Thank you for your attention!

Supplementary Material

Currently Most Attractive Models: Chaotic Inflation

Fractional power-law potential:

- ▶ Axion monodromy in string theory.

[Silverstein & Westphal '08]

$$V(\phi) \propto \phi^{2/3}, \quad \phi^{2/5}$$

- ▶ Strong gauge dynamics in field theory.

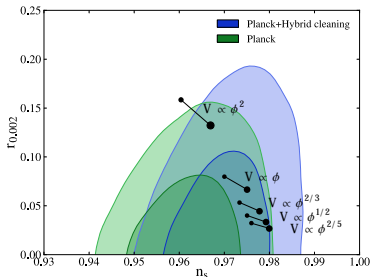
[Harigaya, Ibe, K.S. & Yanagida '13]

$$V(\phi) \propto \phi^{2/(N_c+1)}, \quad G = Sp(N_c)$$

- ▶ Will be probed this year by PLANCK polarization data.

Re-analysed PLANCK data corrected for systematics in the 217 GHz map:

[Spergel, Flauger & Hlozek '13]



Common features of Starobinsky & chaotic Inflation:

- ▶ Large-field models of inflation. \Rightarrow Very sensitive to higher-dim. SUGRA corrections.
- ▶ Particular value of n_s singled out due to particularly shaped scalar potential.

Hybrid inflation is a small-field model, in which n_s is *a priori* undetermined!