

The 126 GeV Higgs boson mass and naturalness in (deflected) mirage mediation

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LHC results

- discovery of the 126 GeV Higgs boson

the Higgs boson mass in the MSSM

$$m_h^2 \sim m_Z^2 \cos^2 2\beta + \frac{3v_u^2 y_t^4}{8\pi^2} \left[\log \frac{M_{st}^2}{m_t^2} + \frac{2A_t^2}{M_{st}^2} \left(1 - \frac{A_t^2}{12M_{st}^2} \right) \right]$$

→ heavy stop mass ($\gtrsim 1$ TeV)

- no SUSY signatures

→ heavy SUSY particles ($\gtrsim 1$ TeV)

126 GeV Higgs boson

no SUSY signature

high-scale SUSY ?

high-scale SUSY vs naturalness

high-scale SUSY is disfavored from naturalness

➤ EWSB vacuum

$$m_Z^2 \simeq -2\mu(m_Z)^2 + 2|m_{H_u}^2(m_Z)|$$

high-scale SUSY leads the fine-tuning problem

➤ degree of tuning the μ parameter

$$\Delta_\mu^{-1} \times 100 \% \quad \text{where} \quad \Delta_\mu = \left| \frac{\partial \ln m_Z}{\partial \ln \mu_0} \right|$$

e.g.) CMSSM ($A_0 = 0$)

126 GeV Higgs boson \longrightarrow $10^{-3} \%$ fine-tuning
RGE

126 GeV Higgs boson

no SUSY signature

high-scale SUSY ?

126 GeV Higgs boson

no SUSY signature

unnatural SUSY ?

RG effects to soft parameters

this argument is based on RG-effects

➤ naturalness $m_Z^2 \simeq -2\mu(m_Z)^2 + 2|m_{H_u}^2(m_Z)|$

$$m_{H_u}^2(m_Z) \simeq +0.17M_2^2 - 0.20M_2M_3 - 3.09M_3^2 - 0.23m_0^2$$

GUT scale

➤ the Higgs boson mass

$$m_h^2 \sim m_Z^2 \cos^2 2\beta + \frac{3v_u^2 y_t^4}{8\pi^2} \left[\log \frac{M_{st}^2}{m_t^2} + \frac{2A_t^2}{M_{st}^2} \left(1 - \frac{A_t^2}{12M_{st}^2} \right) \right]$$

$$m_{t_R}^2(m_Z) \simeq -0.21M_2^2 + 4.61M_3^2 + 0.19m_0^2$$

$$A_t(m_Z) \simeq -0.21M_2 - 1.90M_3 + 0.18A_0$$

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stop-mixing

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RG effects are dominated by the gluino mass

RG effects to soft parameters

➤ $M_2 \sim M_3$ (universal gaugino masses)

$$\begin{aligned} m_{H_u}^2(m_Z) &\simeq +0.17M_2^2 - 0.20M_2M_3 - 3.09M_3^2 - 0.23m_0^2 \\ &\sim -3.12M_3^2 + \dots \end{aligned}$$

severe fine-tuning when M_{SUSY} increases

$$\begin{aligned} m_{\tilde{t}_R}^2(m_Z) &\simeq -0.21M_2^2 + 4.61M_3^2 + 0.19m_0^2 \sim 4.61M_3^2 + \dots \\ A_t(m_Z) &\simeq -0.21M_2 - 1.90M_3 + 0.18A_0 \sim -1.90M_3 + \dots \end{aligned}$$

$$\longrightarrow \left| \frac{A_t}{M_{stop}} \right| < 1 \quad \text{small stop mixing}$$

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$$m_{\tilde{t}_R}^2(m_Z) \simeq -0.21M_2^2 + 4.61M_3^2 + 0.19m_0^2 \sim 4.61M_3^2 + \dots$$

$$A_t(m_Z) \simeq -0.21M_2 - 1.90M_3 + 0.18A_0 \sim -1.90M_3 + \dots$$

$$\longrightarrow \left| \frac{A_t}{M_{stop}} \right| < 1 \quad \text{small stop mixing}$$

➔ heavy stop mass

RG effects to soft parameters

➤ $M_2 \sim 5 \times M_3$

$$m_{H_u}^2(m_Z) \simeq +0.17M_2^2 - 0.20M_2M_3 - 3.09M_3^2 - 0.23m_0^2 \\ \sim 0 \times M_3^2 + \dots$$

➔ the fine-tuning is relaxed even when M_{SUSY} increases

$$m_{t_R}^2(m_Z) \simeq -0.21M_2^2 + 4.61M_3^2 + 0.19m_0^2 \sim -0.2M_3^2 + \dots$$

$$A_t(m_Z) \simeq -0.21M_2 - 1.90M_3 + 0.18A_0 \sim -3.0M_3 + \dots$$

➔ $\left| \frac{A_t}{M_{stop}} \right| \sim \sqrt{3} - \sqrt{6}$ the Higgs boson mass is increased

more detailed analysis can be seen in ref. [1]

126 GeV Higgs boson

no SUSY signature

unnatural SUSY ?

non-universal gaugino masses \rightarrow **No !!**

How explain the non-universal gaugino masses ?

SUSY breaking mediation

soft parameters are determined by mediation mechanisms

➤ SUSY breaking mediations

- gravity mediation ••• higher dimensional interactions
- anomaly mediation ••• super-Weyl anomaly
- gauge mediation ••• gauge interactions

➤ assumptions

- single modulus T is a mediator for the gravity mediation
- minimal gauge mediation with N_{mess} pairs of messengers

$$W_{\text{mess}} = W_1(X) + \lambda X \Psi \bar{\Psi}$$

X : SM gauge singlet

Ψ : **5** of $SU(5)$

gaugino mass ratios

➤ gravity mediation

$$\int d^2\theta T \times \mathcal{W}^a \mathcal{W}^a \quad \longrightarrow \quad M_a(M_{\text{GUT}}) = \frac{F^T}{T + \bar{T}} \equiv m_0$$

gauge coupling unification

universal gaugino masses

gaugino mass ratios

➤ gravity mediation $\longrightarrow M_1 : M_2 : M_3 = 1 : 1 : 1$

gaugino mass ratios

➤ gravity mediation $\longrightarrow M_1 : M_2 : M_3 = 1 : 1 : 1$

➤ anomaly mediation

$$M_a(M_{\text{GUT}}) = \frac{g_0^2}{16\pi^2} b_a \frac{F^C}{C} \equiv g_0^2 b_a \times \alpha_m m_0$$

where $(b_1, b_2, b_3) = (\frac{33}{5}, 1, -3)$ for $U(1)_Y, SU(2)_L, SU(3)_C$

$$\longrightarrow M_1 : M_2 : M_3 = \frac{33}{5} : 1 : -3$$

gaugino mass ratios

➤ gravity mediation $\longrightarrow M_1 : M_2 : M_3 = 1 : 1 : 1$

➤ anomaly mediation $\longrightarrow M_1 : M_2 : M_3 = \frac{33}{5} : 1 : -3$

gaugino mass ratios

➤ gravity mediation $\longrightarrow M_1 : M_2 : M_3 = 1 : 1 : 1$

➤ anomaly mediation $\longrightarrow M_1 : M_2 : M_3 = \frac{33}{5} : 1 : -3$

➤ gauge mediation

$$M_a(M_{\text{mess}}) = N_{\text{mess}} \frac{g_a^2(M_{\text{mess}})}{16\pi^2} \frac{F^X}{X} \equiv N_{\text{mess}} g_a^2(M_{\text{mess}}) \alpha_g \alpha_m m_0$$

where $M_{\text{mess}} = \lambda X$

$\longrightarrow M_3 > M_2$ at any scale

gaugino mass ratios

➤ gravity mediation $\longrightarrow M_1 : M_2 : M_3 = 1 : 1 : 1$

➤ anomaly mediation $\longrightarrow M_1 : M_2 : M_3 = \frac{33}{5} : 1 : -3$

➤ gauge mediation $\longrightarrow M_3 > M_2$ at any scale

gaugino mass ratios

- gravity mediation $\longrightarrow M_1 : M_2 : M_3 = 1 : 1 : 1$
- anomaly mediation $\longrightarrow M_1 : M_2 : M_3 = \frac{33}{5} : 1 : -3$
- gauge mediation $\longrightarrow M_3 > M_2$ at any scale

large M_2/M_3 can't be obtained

(deflected) mirage mediation

we should consider mixed mediations

➤ parameterization

m_0 ... gravity mediated contributions

α_m ... anomaly / gravity

α_g ... gauge / anomaly

➤ mixed mediations

$\alpha_m \simeq \mathcal{O}(1), \alpha_g \simeq 0$ \longrightarrow mirage mediation [2]

$\alpha_m, \alpha_g \simeq \mathcal{O}(1)$ \longrightarrow deflected mirage mediation [3]

[2] K. Choi, K. S. Jeong, T. Kobayashi and K. -i. Okumura, Phys. Rev. D **75**, 095012 (2007).

R. Kitano and Y. Nomura, Phys. Lett. B **631** (2005) 58.

[3] L. L. Everett, I. -W. Kim, P. Ouyang and K. M. Zurek, Phys. Rev. Lett. **101**, 101803 (2008).

TeV scale (deflected) mirage mediation

- gaugino masses at the GUT scale

$$M_a(M_{\text{GUT}}) = \frac{F^T}{T + \bar{T}} + \frac{g_0^2}{16\pi^2} b_a \frac{F^C}{C}$$
$$\simeq m_0 \left(1 + \frac{1}{4} g_0^2 b_a \alpha_m \right) \quad \text{where } (b_1, b_2, b_3) = \left(\frac{33}{5}, 1, -3 \right)$$

if $\alpha_m \sim 2$ \longrightarrow $M_2/M_3 \sim 5$ at the GUT scale

TeV scale mirage mediation [2]

- threshold corrections

$$\Delta M_a(M_{\text{mess}}) = -N_{\text{mess}} \frac{g_a^2(M_{\text{mess}})}{16\pi^2} \left(\frac{F^C}{C} + \frac{F^X}{X} \right)$$

\longrightarrow various patterns of gaugino masses

moduli stabilization

m_0, α_m, α_g are determined by moduli stabilization scenarios

➤ the KKLT-type moduli stabilization

- the original KKLT model predicts $\alpha_m = 1$ [3]
- similar setups can lead various $O(1)$ values [4]
- value of α_m could depend on only discrete parameters
winding numbers, # of fluxes, e.t.c.

➔ we can take $\alpha_m \sim 2$

➤ stabilization of X [5]

e.g.)
$$W_1(X) = \frac{X^n}{\Lambda^{n-3}}, \quad (n < 0 \text{ or } n > 3) \quad \longrightarrow \quad \alpha_g = -\frac{2}{n-1}$$

➔ gauge mediation can also be comparable

[3] S. Kachru, R. Kallosh, A. D. Linde and S. P. Trivedi, Phys. Rev. D **68**, 046005 (2003).

[4] M. Abe, T. Higaki and T. Kobayashi, Phys. Rev. D **73**, 046005 (2006), Nucl. Phys. B **742**, 187 (2006).

[5] L. L. Everett, I. -W. Kim, P. Ouyang and K. M. Zurek, JHEP **0808**, 102 (2008). A. Pomarol and R. Rattazzi, JHEP **9905**, 013 (1999).

input parameters

eight input parameters

$$(m_0, \alpha_m, \alpha_g, n_Q, n_H, N_{\text{mess}}, M_{\text{mess}}, \tan \beta)$$

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$$(\underline{m_0, \alpha_m, \alpha_g}, n_Q, n_H, N_{\text{mess}}, M_{\text{mess}}, \tan \beta)$$

size of mediation

input parameters

eight input parameters

$$(\underline{m_0}, \alpha_m, \alpha_g, n_Q, n_H, N_{\text{mess}}, M_{\text{mess}}, \tan \beta)$$

size of mediation

➤ modular weights n_i

$$a^{ijk}|_{\text{modulus}} = m_0 \sum_{l=i,j,k} (1 - n_l), \quad m^2_i{}^j|_{\text{modulus}} = m_0^2 (1 - n_i) \delta_i^j$$

we assume universal values for quarks/leptons and Higgses, respectively

$$n_Q \qquad n_H$$

input parameters

eight input parameters

$$(m_0, \alpha_m, \alpha_g, n_Q, n_H, N_{\text{mess}}, M_{\text{mess}}, \tan \beta)$$

size of mediation



modular weight

input parameters

eight input parameters

$$(m_0, \alpha_m, \alpha_g, n_Q, n_H, N_{\text{mess}}, M_{\text{mess}}, \tan \beta)$$

size of mediation



messenger sector

modular weight

input parameters

eight input parameters

$$\underbrace{(m_0, \alpha_m, \alpha_g, n_Q, n_H)}_{\text{size of mediation}} \underbrace{(N_{\text{mess}}, M_{\text{mess}})}_{\text{messenger sector}} \underbrace{\tan \beta}_{v_u/v_d}$$

modular weight

input parameters

eight input parameters

$$(m_0, \alpha_m, \alpha_g, n_Q, n_H, N_{\text{mess}}, M_{\text{mess}}, \tan \beta)$$

UV model setups $\longrightarrow n_Q, n_H, N_{\text{mess}}$

moduli stabilization $\longrightarrow m_0, \alpha_m, \alpha_g, M_{\text{mess}}$

μ, b -term are chosen to realize m_Z and $\tan \beta = 15$

input parameters

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μ , b-term are chosen to realize m_Z and $\tan \beta = 15$

➤ EWSB vacuum

$$m_Z^2 = -2\mu(M_{\text{EW}})^2 + \underbrace{\left[\frac{|m_{H_d}^2(M_{\text{EW}}) - m_{H_u}^2(M_{\text{EW}})|}{\sqrt{1 - \sin 2\beta}} - m_{H_u}^2(M_{\text{EW}}) - m_{h_d}^2(M_{\text{EW}}) \right]}_{\text{moduli stabilization}}$$

➔ we focus on the tuning of the μ parameter

How explain the non-universal gaugino masses ?

→ TeV scale (deflected) mirage mediation

modular weights

- mirage unification scenario $(n_Q, n_H) = (1/2, 1)$

If $\sum_{l=i,j,k} (1 - n_l) = 1$ for sizable Yukawa couplings y^{ijk}

➔ all soft terms are unified at the “mirage unification scale”

➔ small $m_{H_u}^2(M_{EW})$ can be obtained easily

- large A-term scenario $(n_Q, n_H) = (0, 0)$

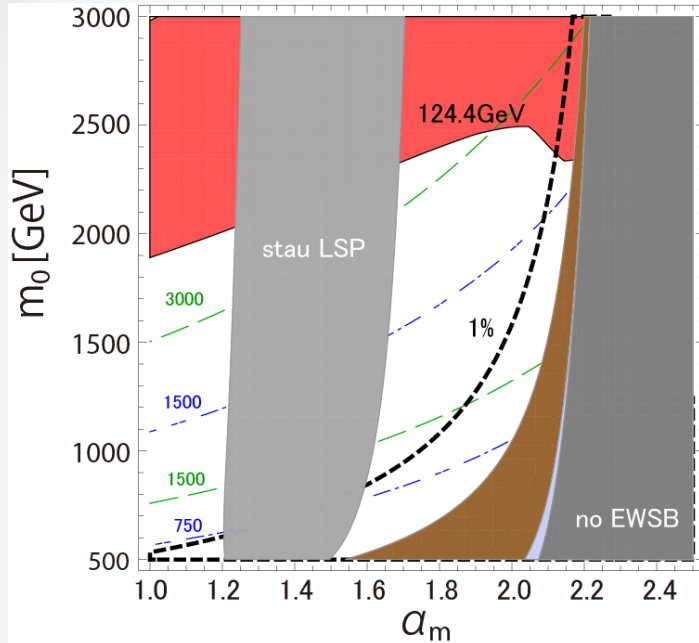
$$a_{\text{moduli}}^{ijk} = m_0 \sum_{l=i,j,k} (1 - n_l), \quad m_{i\text{moduli}}^2{}^j = m_0^2 (1 - n_i) \delta_i^j$$

➔ smaller modular weights will lead the larger stop-mixing

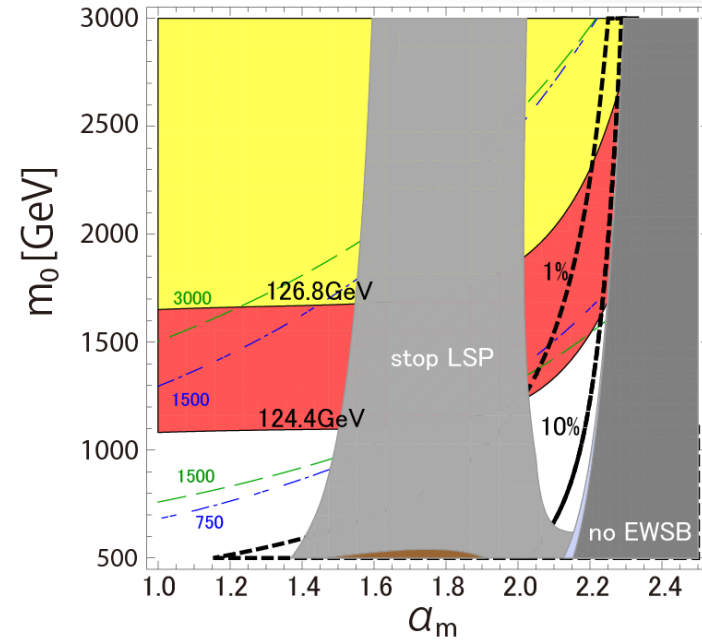
➔ increase the Higgs boson mass

mirage mediation $N_{\text{mess}} = 0$

mirage unification



large A-term



m_0 : SUSY breaking scale

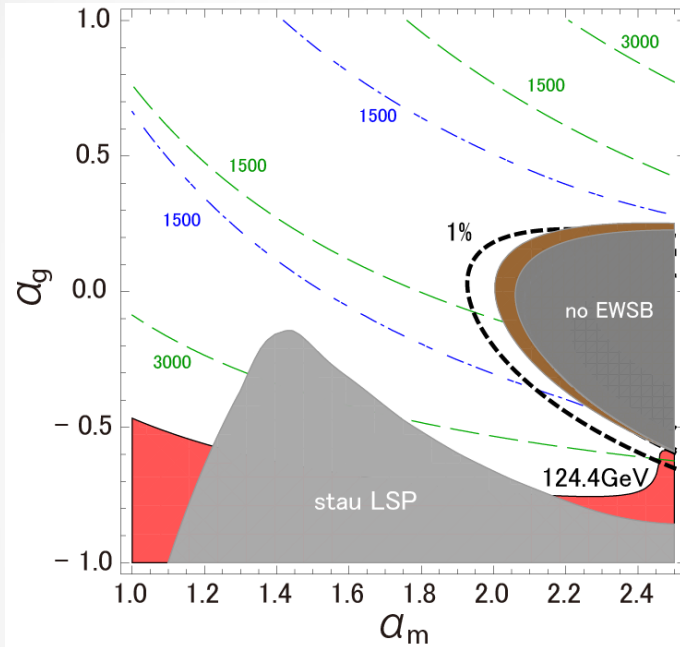
α_m : anomaly/gravity ratio

- large A-term increases the Higgs boson mass
- the tuning is relaxed at $\alpha_m \sim 2$ in **both** cases

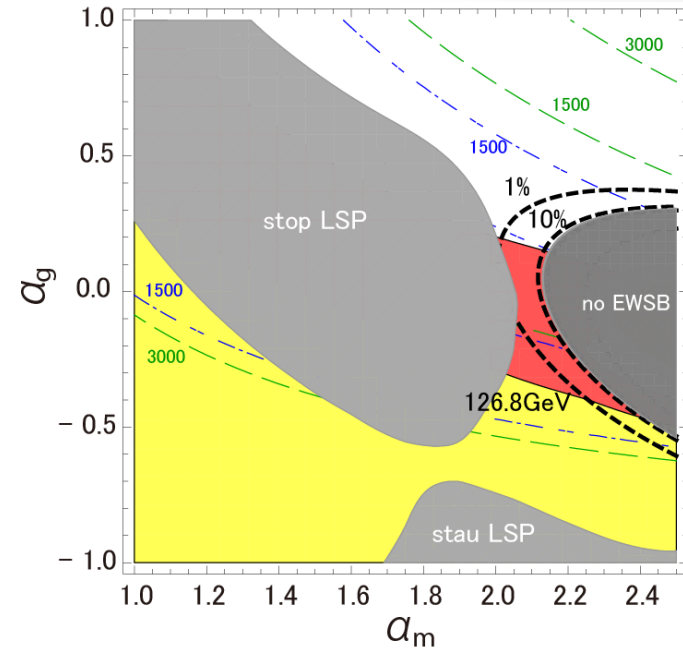
deflected mirage mediation

$$m_0 = 2.0[\text{TeV}], N_{\text{mess}} = 3$$

mirage unification



large A-term



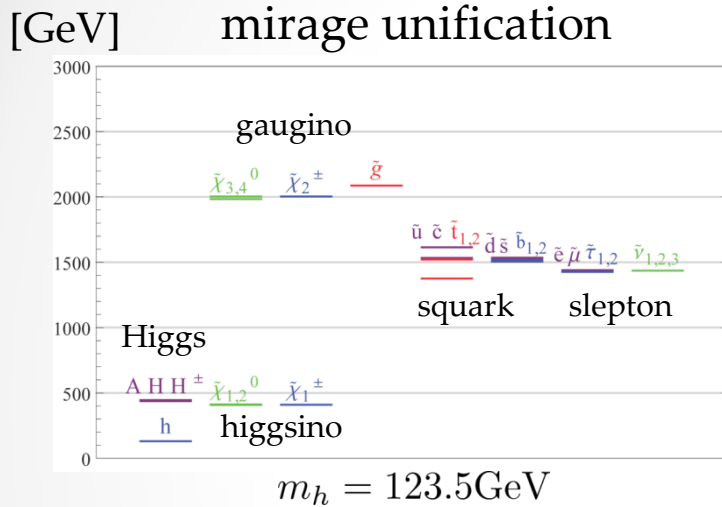
α_m : anomaly/gravity ratio

α_g : gauge/anomaly ratio

- large A-term increases the Higgs boson mass
- the tuning is relaxed at $\alpha_m \sim 2$, $\alpha_g \lesssim 0$ in both cases

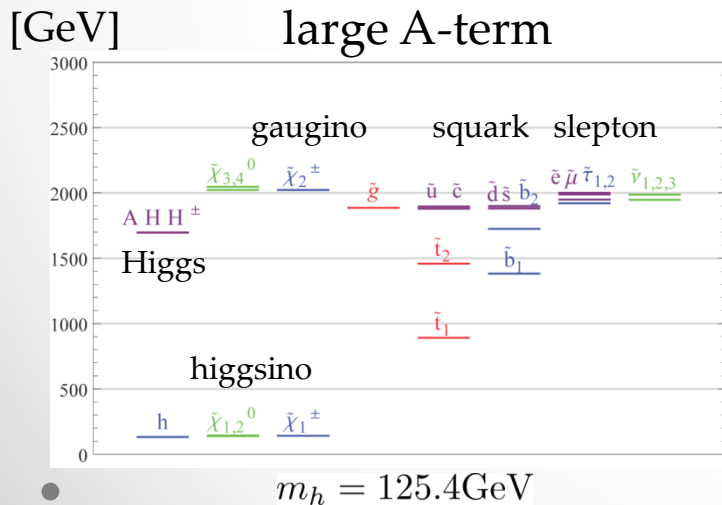
Typical Natural Mass Spectra

naturalness restricts the parameter space of the DMM



$$N_{\text{mess}} = 0, m_0 = 2.0\text{TeV}$$

- gaugino masses are roughly degenerate
- the LSP is higgsino
- spectra are similar in the DMM case



conclusions

➤ 126 GeV Higgs boson and the relaxed tuning

- $M_2 / M_3 \sim 5$ at the GUT scale
- $\alpha_m \sim 2$, $-1 \lesssim \alpha_g \lesssim 0$ in (deflected) mirage mediation
- small modular weights are favored from the 126 GeV Higgs boson

conclusions

➤ 126 GeV Higgs boson and the relaxed tuning

- $M_2 / M_3 \sim 5$ at the GUT scale
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natural SUSY can go !!

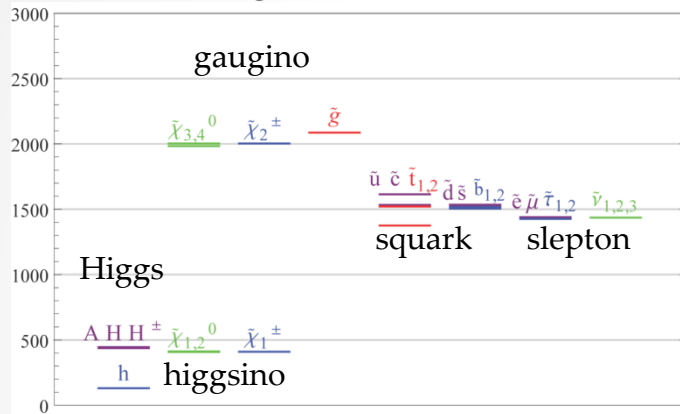
Thank you for your attention

back up

Typical Natural Mass Spectrum

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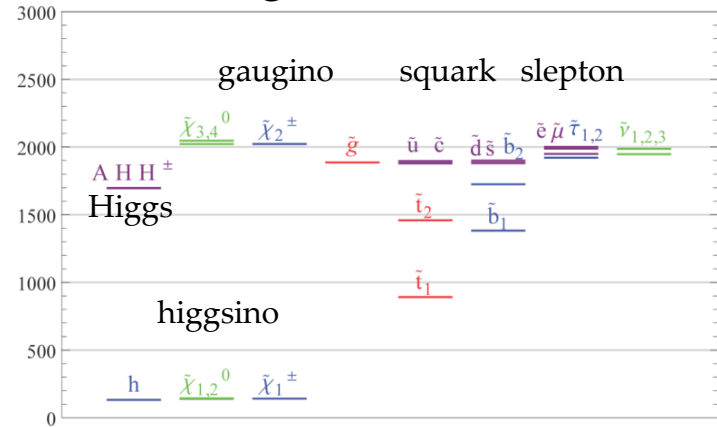
[GeV] mirage unification $\alpha_m = 2.14$



$$m_h = 123.5 \text{ GeV}$$

$$\Delta_\mu^{-1} \times 100 = 2.31\%$$

[GeV] large A-term $\alpha_m = 2.26$



$$m_h = 125.4 \text{ GeV}$$

$$\Delta_\mu^{-1} \times 100 = 55.6\%$$

- stop can be lighter than 1 TeV in the large A-term case
- heavy Higgs bosons tend to be light in the mirage unification case due to the mirage unification \rightarrow enhance $BR(b \rightarrow s \gamma)$

implications of the mass spectra

➤ typical mass spectra

- light colored particles, especially light stop
- uncolored sparticle masses are almost same as colored sparticle masses
- higgsino LSP

- can be tested at the LHC

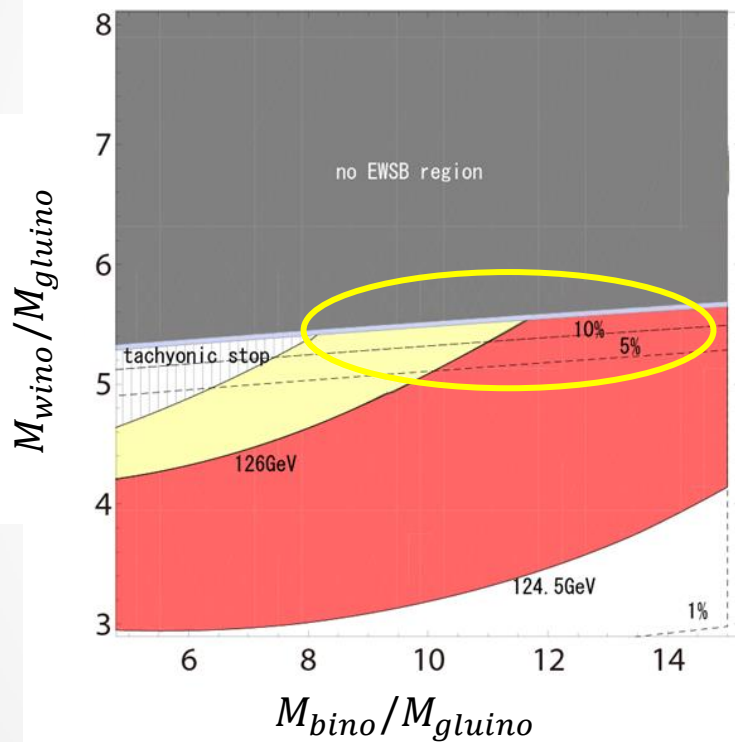


- will be excluded when some signatures for light uncolored particles are detected
- it's challenging to explain the observations for dark matters

axino LSP, non-minimal cosmological scenario, ...

non-universal gaugino masses

[1] H. Abe, J. K. and H. Otsuka, PTEP 2013, 013B02 (2013).



$$\tan \beta = 15$$

$$M_3 = 385 \text{ GeV}$$

$$(m_0)_3 = 200 \text{ GeV}$$

$$(m_0)_{1,2} = 1.5 \text{ TeV}$$

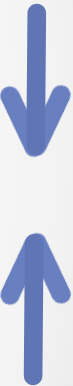
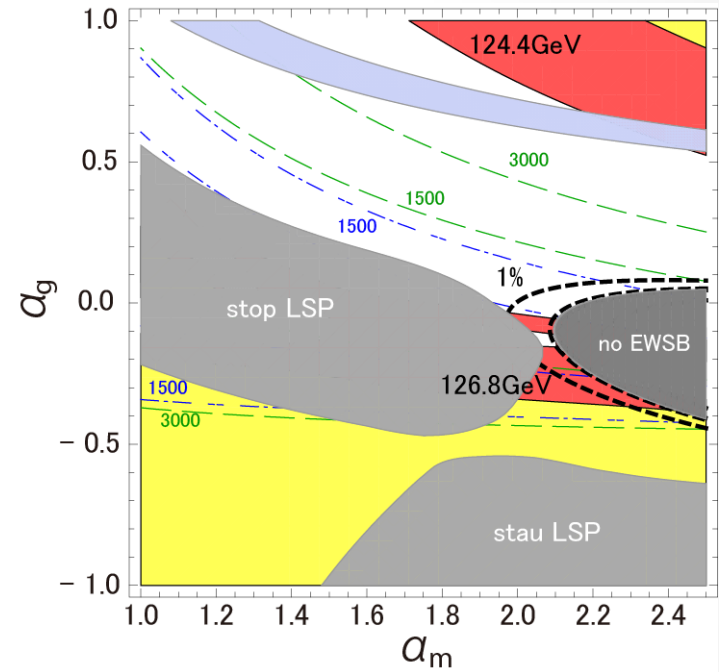
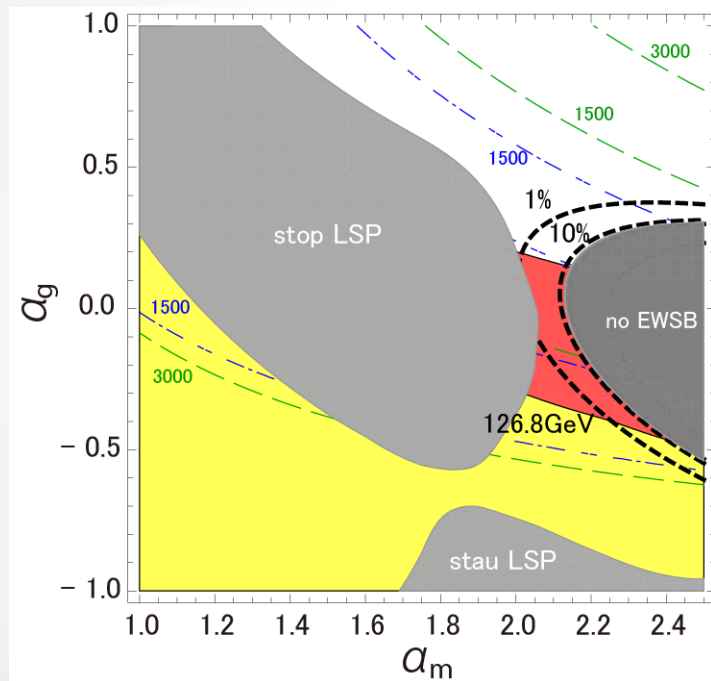
$$A_0 = -400 \text{ GeV}$$

- relaxed fine-tuning and large stop-mixing

messenger sector dependence

$$m_0 = 2.0\text{TeV}, (n_Q, n_H) = (0, 0)$$

$$N_{\text{mess}} = 3 \rightarrow N_{\text{mess}} = 6$$

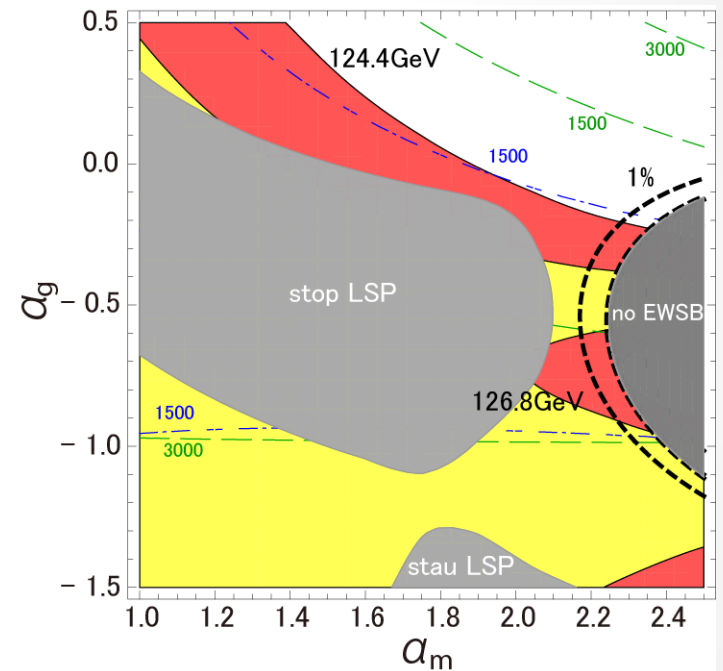
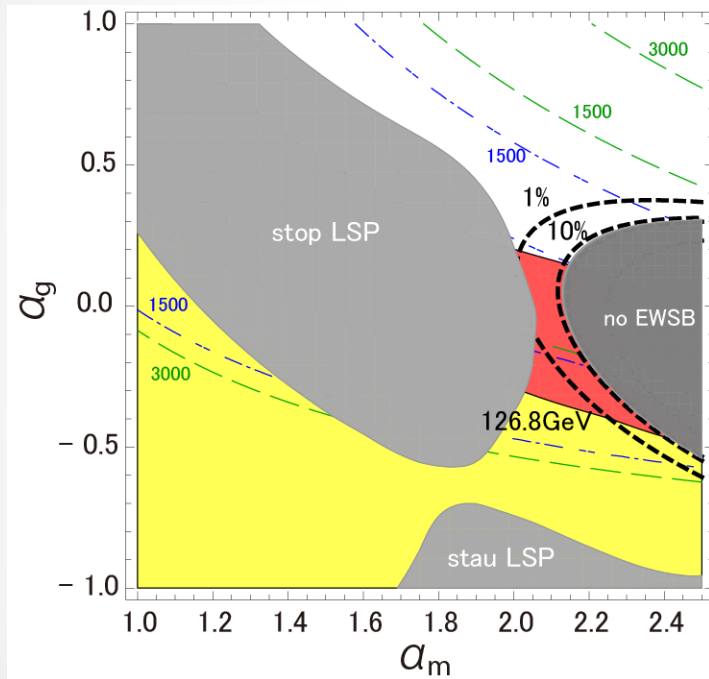


- aimed region is **compressed** along α_g direction

messenger sector dependence

$$m_0 = 2.0\text{TeV}, (n_Q, n_H) = (0, 0)$$

$$M_{\text{mess}} = 10^{12}\text{GeV} \rightarrow M_{\text{mess}} = 10^6\text{GeV}$$



- the aimed region is **shifted** along the α_g direction

model setup

➤ effective SUGRA action with single modulus field

$$\mathcal{L} = - \int d^4\theta \ 3|C|^2 e^{-K/3} + \left[\int d^2\theta f_a \mathcal{W}^a \mathcal{W}^a + \int d^2\theta C^3 W + \text{h.c.} \right]$$

where

$$f_a = T$$

$$K = -3 \ln(T + \bar{T}) + \frac{X\bar{X}}{(T + \bar{T})^{n_X}} + \frac{\Phi_i \bar{\Phi}_i}{(T + \bar{T})^{n_i}}$$

$$W = \underbrace{W_0(T) + W_1(X)}_{\text{stabilize } T, X} + \underbrace{\lambda X \Psi \bar{\Psi}}_{\text{messenger}} + W_{\text{MSSM}}$$

T : moduli

C : compensator

X : SM gauge singlet

Φ_i : MSSM matter

Ψ : $\mathbf{5}$ of $SU(5)$

n_i : modular weights

soft parameters in the DMM

➤ soft parameters at the GUT scale

$$M_a(M_{\text{GUT}}) = m_0 \left[1 + \frac{g_0^2}{16\pi^2} b'_a \alpha_m \ln \frac{M_p}{m_{3/2}} \right]$$
$$a^{ijk}(M_{\text{GUT}}) = m_0 \left[(3 - n_i - n_j - n_k) - \frac{1}{16\pi^2} [y^{ljk} \gamma_l^i + \text{cyclic}] \alpha_m \ln \frac{M_p}{m_{3/2}} \right]$$
$$m^2_i{}^j(M_{\text{GUT}}) = m_0^2 \left[(1 - n_i) \delta_i^j - \frac{2\theta_i^j}{16\pi^2} \alpha_m \ln \frac{M_p}{m_{3/2}} - \frac{\dot{\gamma}_i^j}{(16\pi^2)^2} \left(\alpha_m \ln \frac{M_p}{m_{3/2}} \right)^2 \right]$$

➤ threshold corrections at the messenger scale

$$\Delta M_a(M_{\text{mess}}) = -m_0 N \frac{g_a^2(M_{\text{mess}})}{16\pi^2} \alpha_m (1 + \alpha_g) \ln \frac{M_p}{m_{3/2}}$$
$$\Delta m^2_i{}^j(M_{\text{mess}}) = m_0^2 \sum_a 2c_a(\Phi_i) N \frac{g_a^4(M_{\text{mess}})}{(16\pi^2)^2} \left[\alpha_m (1 + \alpha_g) \ln \frac{M_p}{m_{3/2}} \right]^2 \delta_i^j$$

➤ parameterization

$$\frac{F^T}{T + \bar{T}} \equiv m_0, \quad \frac{F^C}{C} = m_0 \left(\alpha_m \ln \frac{M_p}{m_{3/2}} \right), \quad \frac{F^X}{X} = \alpha_g \frac{F^C}{C}$$

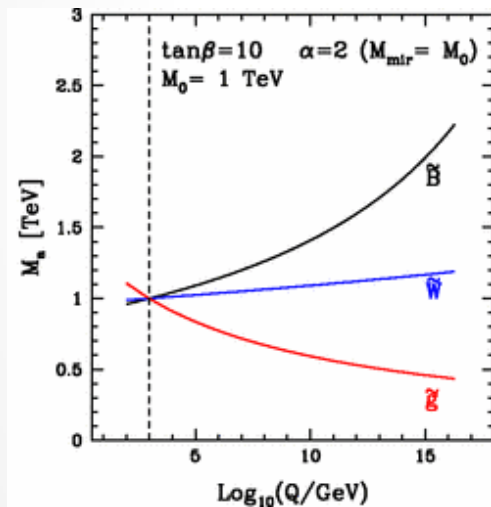
mirage unification

If $\sum_{l=i,j,k} (1 - n_l) = 1$ for sizable Yukawa couplings y^{ijk}

→ all soft terms are unified at the “mirage unification scale”

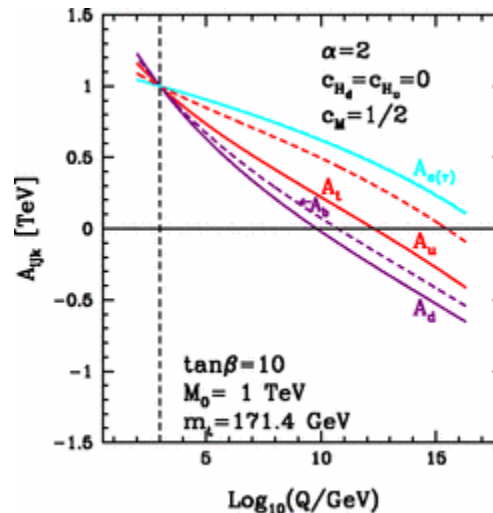
K. Choi, K. S. Jeong, T. Kobayashi and K. -i. Okumura, Phys. Rev. D 75, 095012 (2007).

gaugino masses



↑
always unify

A-terms



↑
depending on modular weights

soft masses

