

BSM Primary Effects

Rick Sandeepan Gupta (IFAE, Barcelona)

in collaboration with A. Pomarol and F. Riva ([arxiv: 1405.0181](#))

& J. Elias-Miro, C. Grojean, D Marzocca ([arxiv: 1312.2928](#))

+ Plan of Talk



- I. BSM Primary effects with A. Pomarol and F. Riva ([arxiv: 1405.0181](#))

- II. RG-induced constraints on BSM Primaries
with J. Elias-Miro, C. Grojean, D Marzocca ([arxiv: 1312.2928](#))

+ SM as an EFT

- The absence at the LHC of new states beyond the SM (BSM) suggests that the new-physics scale must be heavier than the electroweak (EW) scale and we can write:

$$\mathcal{L}_{\text{eff}} = \frac{\Lambda^4}{g_*^2} \mathcal{L} \left(\frac{D_\mu}{\Lambda}, \frac{g_* H}{\Lambda}, \frac{g_* f_{L,R}}{\Lambda^{3/2}}, \frac{g F_{\mu\nu}}{\Lambda^2} \right) \simeq \mathcal{L}_4 + \mathcal{L}_6 + \dots$$

- **No of measurements > No of couplings/parameters at a fixed order => Predictions relating different measurements.** Predictions from \mathcal{L}_4 (the SM lagrangian) well known and tested:

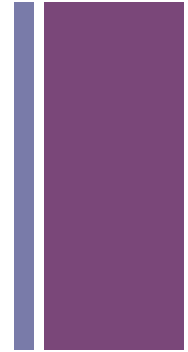
+ SM as an EFT

- The absence at the LHC of new states beyond the SM (BSM) suggests that the new-physics scale must be heavier than the electroweak (EW) scale and we can write:

$$\mathcal{L}_{\text{eff}} = \frac{\Lambda^4}{g_*^2} \mathcal{L} \left(\frac{D_\mu}{\Lambda}, \frac{g_* H}{\Lambda}, \frac{g_* f_{L,R}}{\Lambda^{3/2}}, \frac{g F_{\mu\nu}}{\Lambda^2} \right) \simeq \mathcal{L}_4 + \mathcal{L}_6 + \dots$$

- **No of measurements > No of couplings/parameters at a fixed order => Predictions relating different measurements.** Predictions from \mathcal{L}_4 (the SM lagrangian) well known and tested:

$$m_W = m_Z c\theta_W \quad Y_f = \sqrt{2}m_f/v \quad \text{etc}$$



Main Goal: What are the **predictions from \mathcal{L}_6** ? For eg. **which (non SM) Higgs interactions** are already **constrained** by **EWPT and TGC** data and **which** are still **independent** ?

+ BSM Primaries

- 18 quantities **best** constrain the important deformations in \mathcal{L}_6
- We call these **BSM Primaries**. (see also Pomarol & Riva, 2013, Elias-Miro, Espinosa, Masso & Pomarol, 2013)

Higgs (8)
Physics

$$h \rightarrow \gamma\gamma, h \rightarrow \gamma Z, h \rightarrow gg$$

$$h \rightarrow VV, h \rightarrow ff, pp \rightarrow h^* \rightarrow hh$$

$$hA_{\mu\nu}A^{\mu\nu}, hA_{\mu\nu}Z^{\mu\nu}, hG_{\mu\nu}G^{\mu\nu}$$

$$hW^{+\mu}W_{\mu}^{-}, h\bar{f}f, h^3$$

EWPT (7)
Data

$$Z \rightarrow ff$$

(2 can be traded for S, T)

$$Z_{\mu}f_{L,R}\bar{\gamma}^{\mu}f_{L,R}$$

TGC (3)
Data

$$ee \rightarrow WW$$

$$g_1^Z c_{\theta_W} Z^{\mu} \left(W^{+\nu} \hat{W}_{\mu\nu}^{-} - W^{-\nu} \hat{W}_{\mu\nu}^{+} \right)$$

$$\kappa_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} W_{\mu}^{+} W_{\nu}^{-}$$

$$\lambda_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} \hat{W}_{\mu}^{-\rho} \hat{W}_{\rho\nu}^{+}$$

(18 is not considering four fermions and MFV suppressed and CPV deformations)

+ BSM Primaries

- 18 quantities **best** constrain the important deformations in \mathcal{L}_6
- We call these **BSM Primaries**. (see also Pomarol & Riva, 2013, Elias-Miro, Espinosa, Masso & Pomarol, 2013)

Higgs (8)
Physics

$$h \rightarrow \gamma\gamma, h \rightarrow \gamma Z, h \rightarrow gg$$

$$h \rightarrow VV, h \rightarrow ff, pp \rightarrow h^* \rightarrow hh$$

$$hA_{\mu\nu}A^{\mu\nu}, hA_{\mu\nu}Z^{\mu\nu}, hG_{\mu\nu}G^{\mu\nu}$$

$$hW^{+\mu}W_{\mu}^-, h\bar{f}f, h^3$$

EWPT (7)
Data

$$Z \rightarrow ff$$

(2 can be traded for S, T)

$$Z_{\mu}f_{L,R}\bar{\gamma}^{\mu}f_{L,R}$$

TGC (3)
Data

$$ee \rightarrow WW$$

$$g_1^Z c_{\theta_W} Z^{\mu} \left(W^{+\nu} \hat{W}_{\mu\nu}^- - W^{-\nu} \hat{W}_{\mu\nu}^+ \right)$$

$$\kappa_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} W_{\mu}^+ W_{\nu}^-$$

$$\lambda_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} \hat{W}_{\mu}^{-\rho} \hat{W}_{\rho\nu}^+$$

(18 is not considering four fermions and MFV suppressed and CPV deformations)

+ Correlated deformations

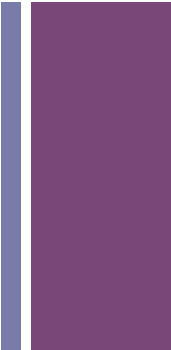
$$hA_{\mu\nu}A^{\mu\nu}, hA_{\mu\nu}Z^{\mu\nu}, hG_{\mu\nu}G^{\mu\nu}$$
$$hW^{+\mu}W_{\mu}^{-}, h\bar{f}f, h^3$$

$$Z_{\mu}\bar{f}_{L,R}\gamma^{\mu}f_{L,R}$$

$$g_1^Z c_{\theta_W} Z^{\mu} \left(W^{+\nu} \hat{W}_{\mu\nu}^{-} - W^{-\nu} \hat{W}_{\mu\nu}^{+} \right)$$
$$\kappa_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} W_{\mu}^{+} W_{\nu}^{-}$$
$$\lambda_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} \hat{W}_{\mu}^{-\rho} \hat{W}_{\rho\nu}^{+}$$

**Primary
Deformations**

+ Correlated deformations



$$hA_{\mu\nu}A^{\mu\nu}, hA_{\mu\nu}Z^{\mu\nu}, hG_{\mu\nu}G^{\mu\nu}$$

$$hW^{+\mu}W_{\mu}^{-}, h\bar{f}f, h^3$$

$$Z_{\mu}\bar{f}_{L,R}\gamma^{\mu}f_{L,R}$$

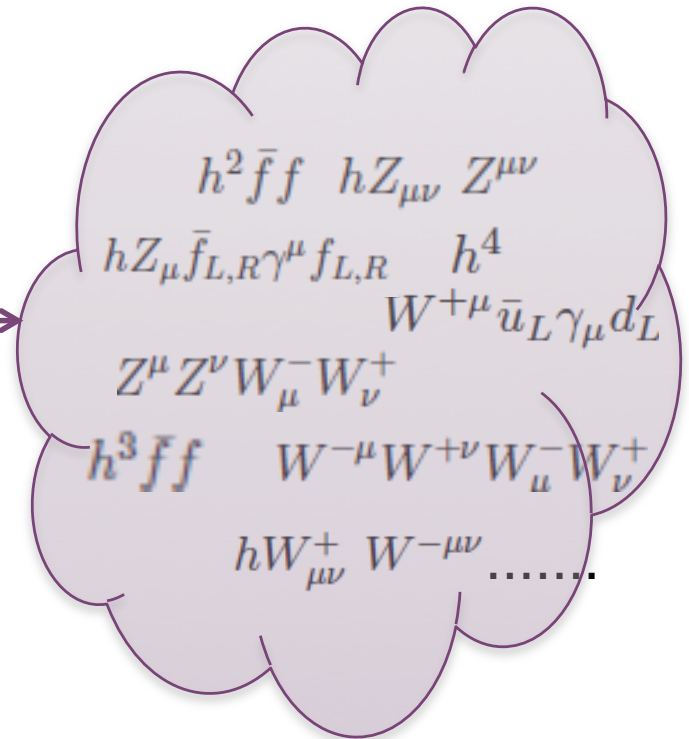
Deformations **correlated**
at dim-6 level

$$g_1^Z c_{\theta_W} Z^{\mu} \left(W^{+\nu} \hat{W}_{\mu\nu}^{-} - W^{-\nu} \hat{W}_{\mu\nu}^{+} \right)$$

$$\kappa_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} W_{\mu}^{+} W_{\nu}^{-}$$

$$\lambda_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} \hat{W}_{\mu}^{-\rho} \hat{W}_{\rho\nu}^{+}$$

**Primary
Deformations**



**Correlated
Deformations**

+ BSM Primary directions

- Cannot generate only BSM primary deformation and no other deformation :

BSM Primary direction

$$\Delta\mathcal{L}_{Z\gamma}^h = 4\kappa_{Z\gamma} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[t_{\theta_W} A_{\mu\nu} Z^{\mu\nu} + \frac{e_{2\theta_W}}{2c_{\theta_W}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^+ W^{-\mu\nu} \right]$$

BSM Primary

Correlated deformations

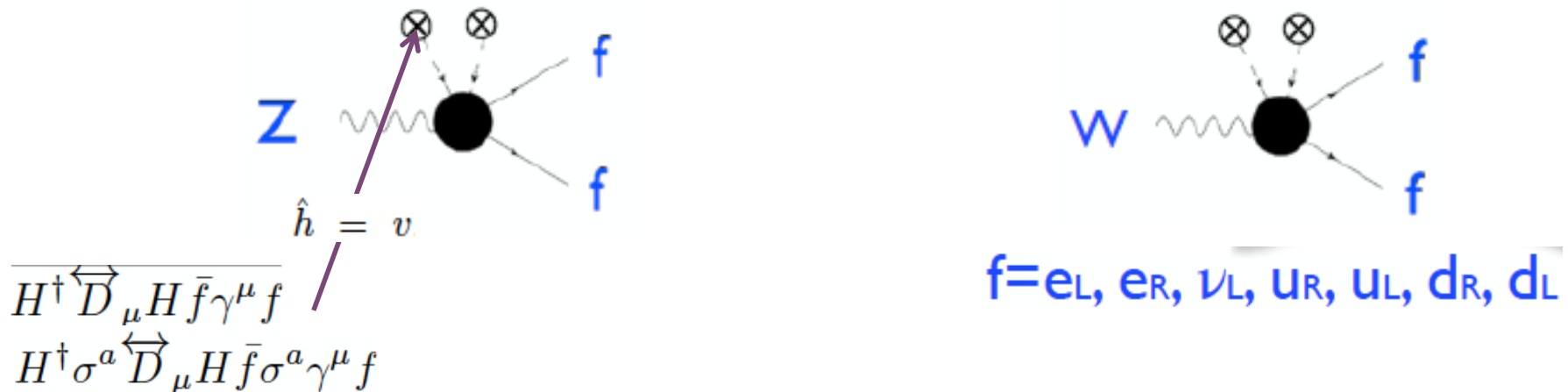
- In other words just the **BSM primary itself is not a dim-6 operator.** There will be other terms from the operator.
- BSM primary directions must be **mutually orthogonal**. For eg. the above terms must not contribute to other 17 primaries like $h \rightarrow \gamma\gamma$.



- We will take a **bottom up approach** to construct the dim-6 Lagrangian by building up these **BSM Primary directions (operators in disguise)** corresponding to each BSM Primary.
- The **dim-6 Lagrangian would be the sum of all these Primary directions.**

+ Z-pole Primaries

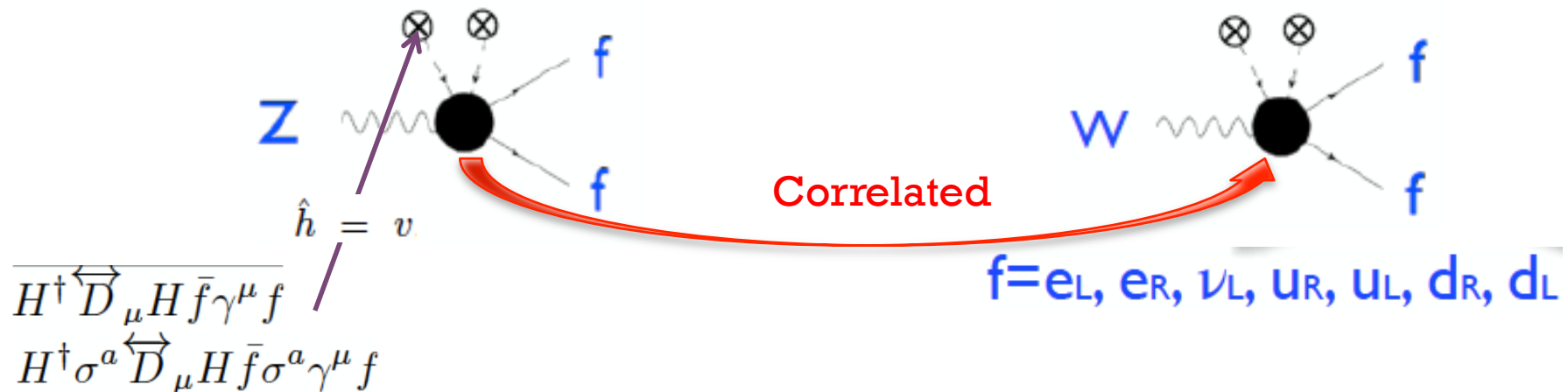
- Zff coupling deformations more general (beyond universal theories) and physical than S and T parameters.



- 3 dim-6 structures for 4 possible W/Z -couplings to leptons **W -couplings related to Z -couplings at dim-6 level**

+ Z-pole Primaries

- Zff coupling deformations more general (beyond universal theories) and physical than S and T parameters.



- 3 dim-6 structures for 4 possible W/Z -couplings to leptons **W -couplings related to Z -couplings at dim-6 level**

+

Deviation from gauge coupling universality

- W/Z couplings can be *individually* altered by operators having product of currents:

$$\begin{aligned}
 \Delta\mathcal{L}_{ee}^V &= \delta g_{eR}^Z \frac{\hat{h}^2}{v^2} Z^\mu \bar{e}_R \gamma_\mu e_R && (15) && \text{Leptons} \\
 &+ \delta g_{eL}^Z \frac{\hat{h}^2}{v^2} \left[Z^\mu \bar{e}_L \gamma_\mu e_L - \frac{c_{\theta_W}}{\sqrt{2}} (W^{+\mu} \bar{\nu}_L \gamma_\mu e_L + \text{h.c.}) \right] \\
 &+ \delta g_{\nu L}^Z \frac{\hat{h}^2}{v^2} \left[Z^\mu \bar{\nu}_L \gamma_\mu \nu_L + \frac{c_{\theta_W}}{\sqrt{2}} (W^{+\mu} \bar{\nu}_L \gamma_\mu e_L + \text{h.c.}) \right]
 \end{aligned}$$

Primary \leftarrow $Z^\mu \bar{e}_L \gamma_\mu e_L$ \rightarrow Correlated interaction

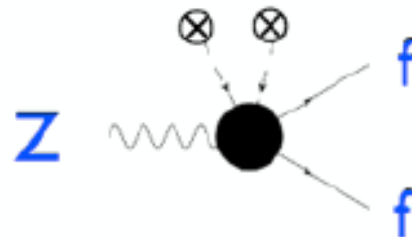
$$\begin{aligned}
 \Delta\mathcal{L}_{qq}^V &= \delta g_{uR}^Z \frac{\hat{h}^2}{v^2} Z^\mu \bar{u}_R \gamma_\mu u_R + \delta g_{dR}^Z \frac{\hat{h}^2}{v^2} Z^\mu \bar{d}_R \gamma_\mu d_R && \text{Quarks} \\
 &+ \delta g_{dL}^Z \frac{\hat{h}^2}{v^2} \left[Z^\mu \bar{d}_L \gamma_\mu d_L - \frac{c_{\theta_W}}{\sqrt{2}} (W^{+\mu} \bar{u}_L \gamma_\mu d_L + \text{h.c.}) \right] \\
 &+ \delta g_{uL}^Z \frac{\hat{h}^2}{v^2} \left[Z^\mu \bar{u}_L \gamma_\mu u_L + \frac{c_{\theta_W}}{\sqrt{2}} (W^{+\mu} \bar{u}_L \gamma_\mu d_L + \text{h.c.}) \right]
 \end{aligned}$$

+ S and T ?

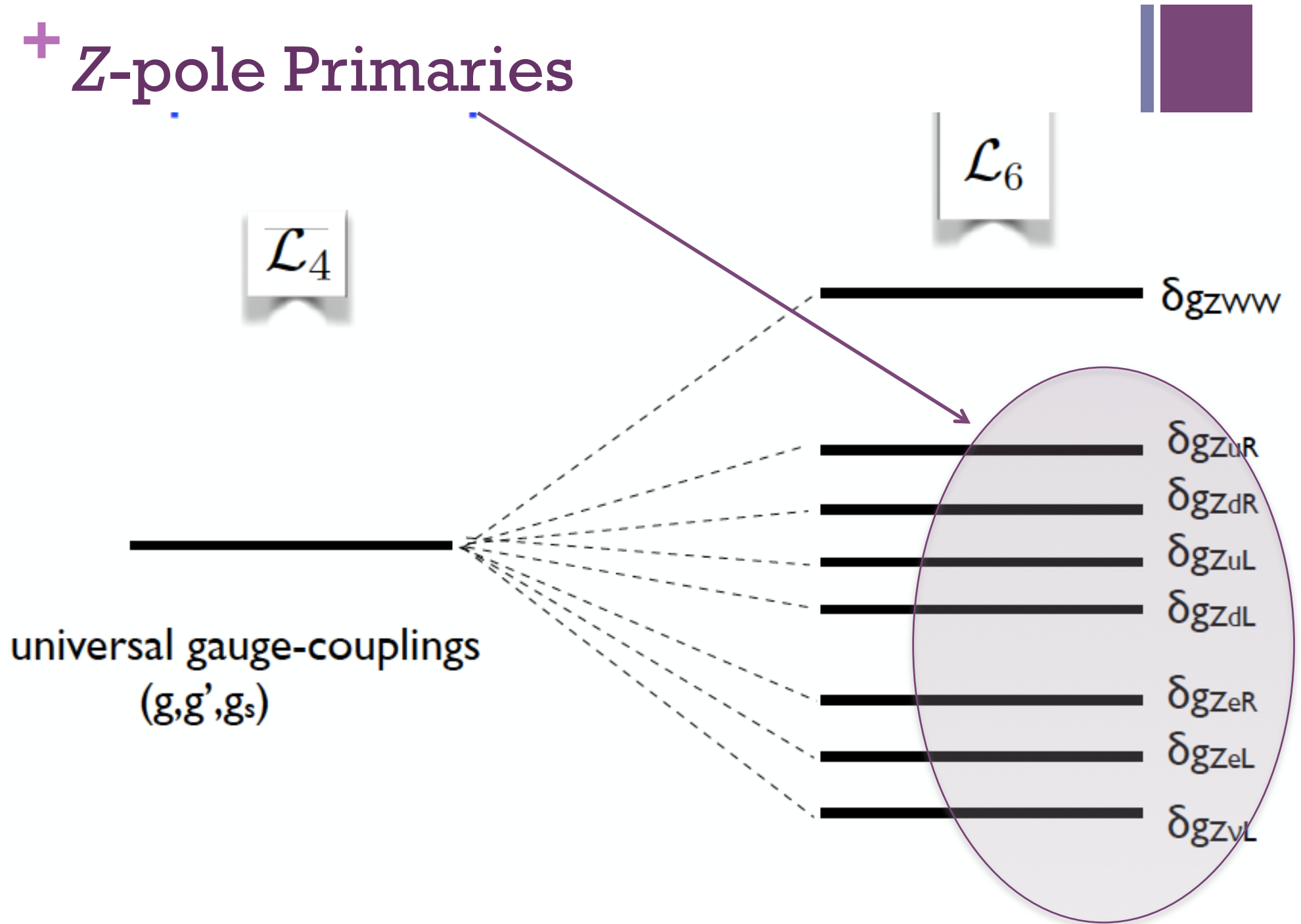
- S and T are linear combinations of above parameters:

$$\Delta\mathcal{L}_{\hat{S}} = \hat{S} \frac{gs_{\theta_w}^2 \hat{h}^2}{c_{\theta_w}^3 v^2} Z_\mu \left[J_Z^\mu - c_{\theta_w}^2 J_{em}^\mu + \frac{g}{c_{\theta_w}} \frac{\hat{h}^2}{4} Z^\mu \right]$$
$$\Delta\mathcal{L}_T = -\frac{\hat{T} \hat{h}^4}{2 v^4} m_Z^2 Z^\mu Z_\mu .$$

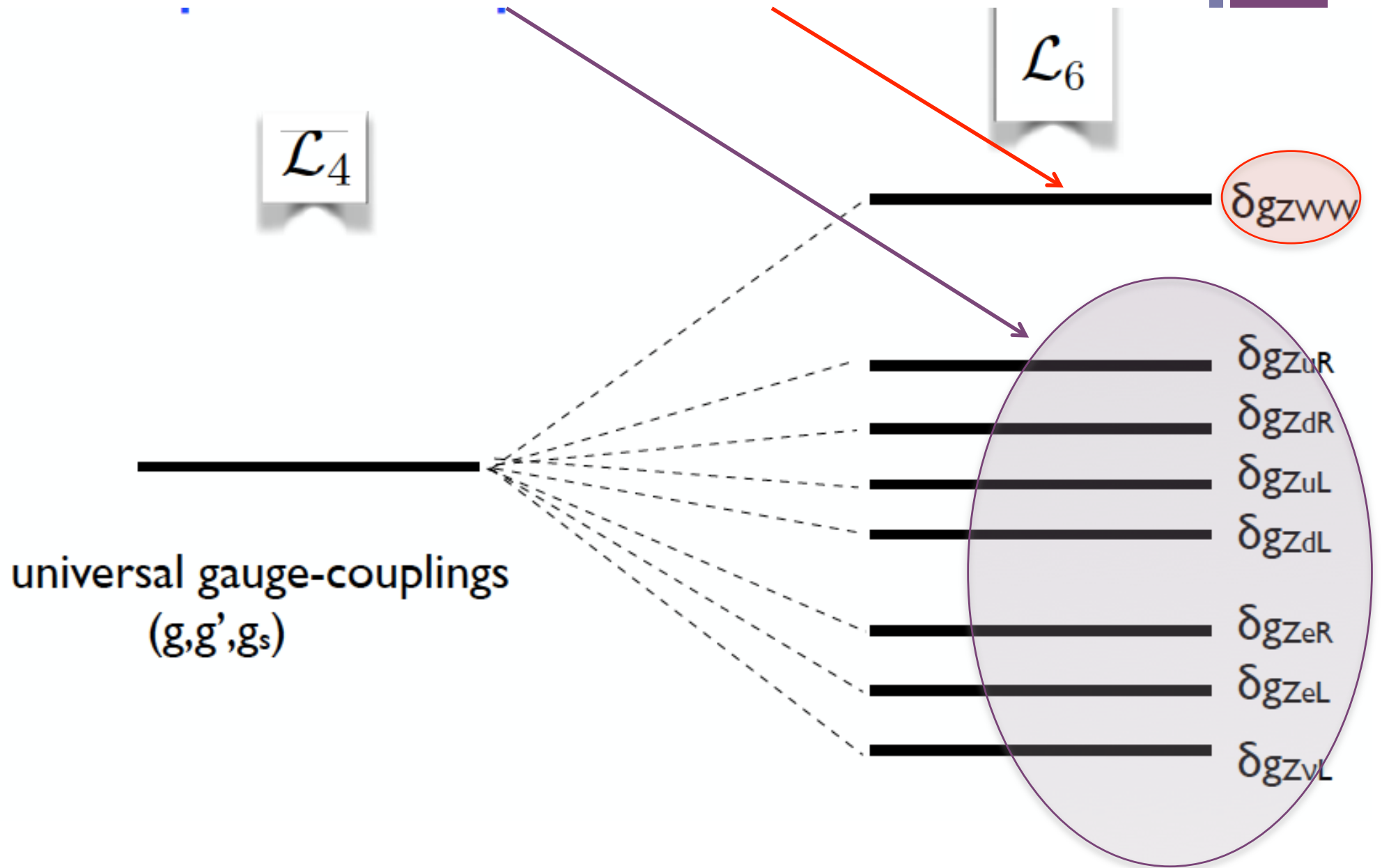
- In our parametrization we have eliminated all corrections to propagators (using EoM) so there is a **one to one correspondence between our δg_f^Z and the Z partial widths.**



+ Z-pole Primaries



+ Z-pole Primaries + TGC



+ Deviation from gauge coupling universality

- A **shift** $s_{\theta_w}^2 \rightarrow s_{\theta_w}^2 (1 + 2\delta g_1^Z c_{\theta_w}^2 \hat{h}^2/v^2)$ keeping e constant in the **fermion –Higgs sector of SM lagrangian** gives:

$$\Delta\mathcal{L}_{g^Z} = \delta g_1^Z c_{\theta_w}^2 \frac{\hat{h}^2}{v^2} \left[\frac{e^2 \hat{h}^2}{4c_{\theta_w}^4} Z^\mu Z_\mu - g(W_\mu^- J_Z^\mu + \text{h.c.}) - \frac{gc_{2\theta_w}}{c_{\theta_w}^3} Z_\mu J_Z^\mu - 2et_{\theta_w} Z_\mu J_{em}^\mu \right]$$

- For $\hat{h} = v$ this shift is just a redefinition of $s_{\theta_w}^2$.
- **Opposite shift** $s_{\theta_w}^2 \rightarrow s_{\theta_w}^2 (1 - 2\delta g_1^Z c_{\theta_w}^2)$ in only pure gauge sector gives **TGC and QGCs**:

$$ig\delta g_1^Z c_{\theta_w} Z^\mu (W^{-\nu} W_{\mu\nu}^+ - W^{+\nu} W_{\mu\nu}^-) + \text{QGCs}$$

+ Deviation from gauge coupling universality

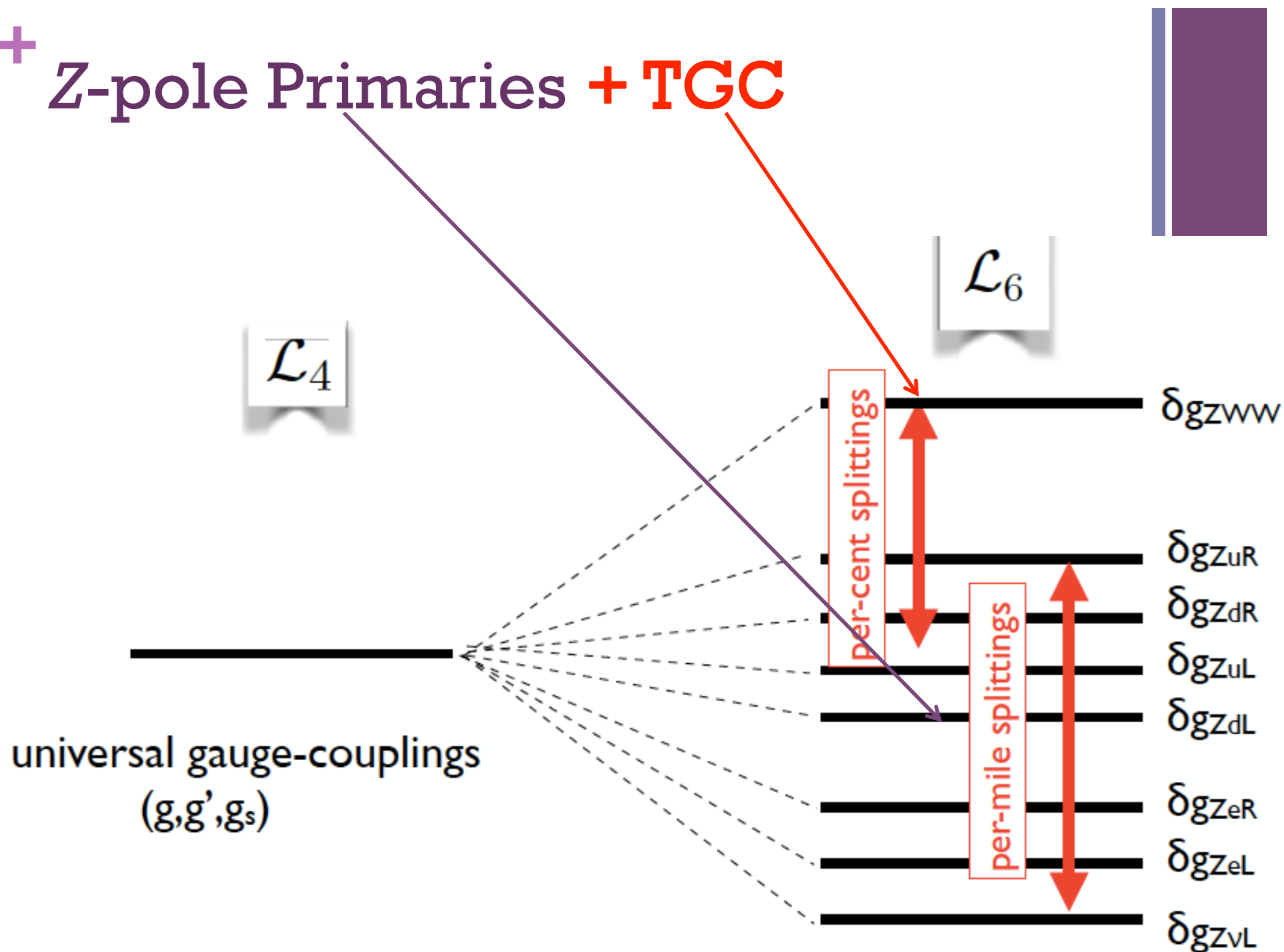
- A **shift** $s_{\theta_w}^2 \rightarrow s_{\theta_w}^2 (1 + 2\delta g_1^Z c_{\theta_w}^2 \hat{h}^2/v^2)$ keeping e constant in the **fermion –Higgs sector of SM lagrangian** gives:

$$\Delta\mathcal{L}_{g^Z} = \delta g_1^Z c_{\theta_w}^2 \frac{\hat{h}^2}{v^2} \left[\frac{e^2 \hat{h}^2}{4c_{\theta_w}^4} Z^\mu Z_\mu - g(W_\mu^- J_Z^\mu + \text{h.c.}) - \frac{gc_{2\theta_w}}{c_{\theta_w}^3} Z_\mu J_Z^\mu - 2et_{\theta_w} Z_\mu J_{em}^\mu \right]$$

- For $\hat{h} = v$ this shift is just a redefinition of $s_{\theta_w}^2$.
- **Opposite shift** $s_{\theta_w}^2 \rightarrow s_{\theta_w}^2 (1 - 2\delta g_1^Z c_{\theta_w}^2)$ in only pure gauge sector gives **TGC and QGCs**:

$$\delta g_1^Z = \frac{\delta g^{ZWW}}{g_{SM}^{ZWW}} = \frac{\delta g^{WWWW}}{2c_{\theta_w}^2 g_{SM}^{WWWW}} = \frac{\delta g^{ZZWW}}{2g_{SM}^{ZZWW}} = \frac{\delta g^{\gamma ZWW}}{g_{SM}^{\gamma ZWW}}$$

+ Z-pole Primaries + TGC



+ Other TGC primary directions

- Notice that the deformation below contains the $\delta\kappa_\gamma$ TGC:

$$\hat{h}^2 \eta^a W_{\mu\nu}^a B^{\mu\nu} = \hat{h}^2 \left[\hat{W}_{\mu\nu}^3 B^{\mu\nu} + 2igc_{\theta_w} W_\mu^- W_\nu^+ (A^{\mu\nu} - t_{\theta_w} Z^{\mu\nu}) \right]$$

$$-\frac{g'\hat{h}^2}{2gv^2} W_{\mu\nu}^3 B^{\mu\nu} = \frac{\Delta\mathcal{L}_{\hat{S}}}{\hat{S}} - \frac{\Delta\mathcal{L}_{\gamma\gamma}^h}{4\kappa_{\gamma\gamma}} - \frac{c_{2\theta_w} \Delta\mathcal{L}_{Z\gamma}^h}{4\kappa_{Z\gamma}} + \frac{\Delta\mathcal{L}_{\kappa_\gamma}}{\delta\kappa_\gamma}$$

- We find a combination that does not contribute to other primaries but only to $\delta\kappa_\gamma$:

$$\Delta\mathcal{L}_{\kappa_\gamma} = \frac{\delta\kappa_\gamma}{v^2} \left[ie\hat{h}^2 (A_{\mu\nu} - t_{\theta_w} Z_{\mu\nu}) W^{+\mu} W^{-\nu} \right. \\ \left. + Z_\nu \partial_\mu \hat{h}^2 (t_{\theta_w} A^{\mu\nu} - t_{\theta_w}^2 Z^{\mu\nu}) + \frac{(\hat{h}^2 - v^2)}{2} \right. \\ \left. \times \left(t_{\theta_w} Z_{\mu\nu} A^{\mu\nu} + \frac{c_{2\theta_w}}{2c_{\theta_w}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^+ W^{-\mu\nu} \right) \right]$$

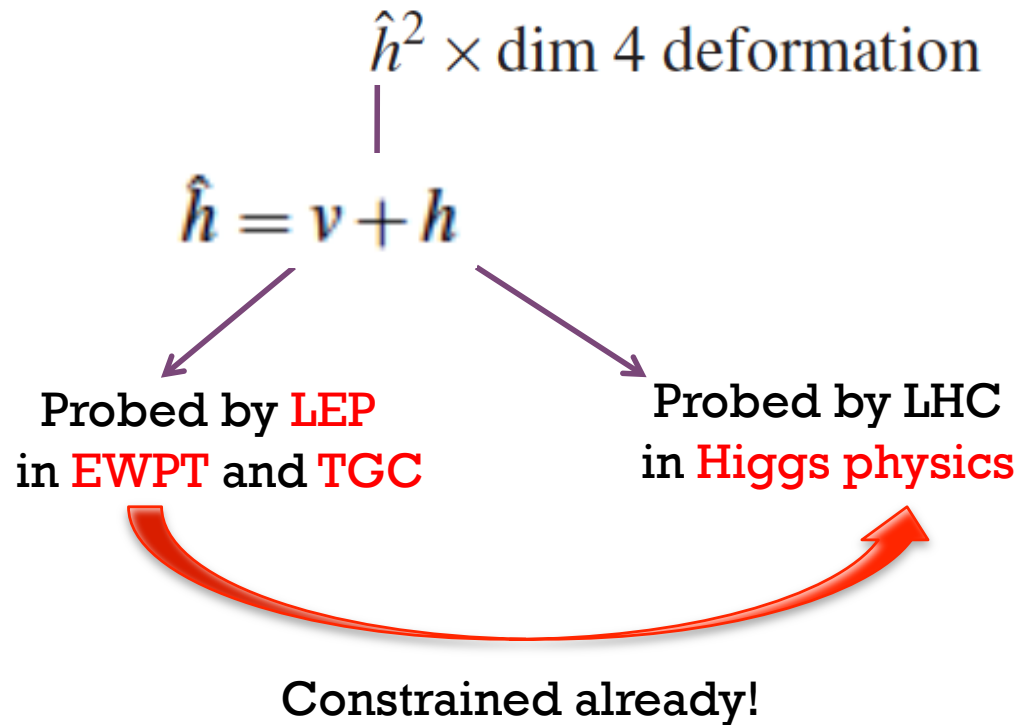
+ Other TGC primary directions

- Finally we also have:

$$\Delta\mathcal{L}_{\lambda_\gamma} = \frac{i\lambda_\gamma}{m_W^2} [(eA^{\mu\nu} + gc\theta_W Z^{\mu\nu})W_\nu^{-\rho}W_{\rho\mu}^+]$$

+ Relation to Higgs physics

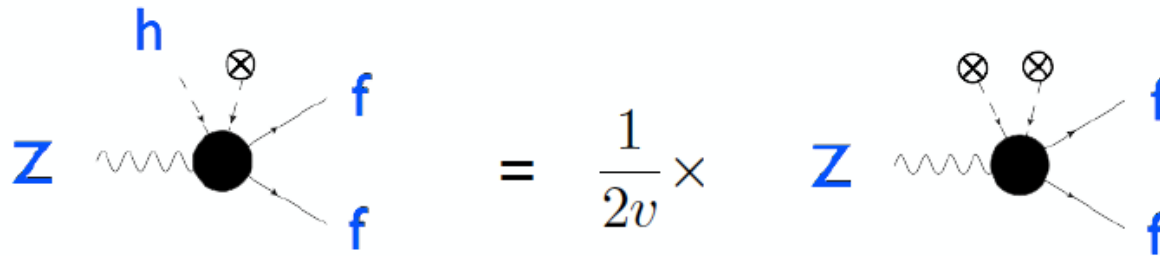
All Primary deformations considered so far have been of the form



+ Relation to Higgs physics

All Primary deformations considered so far have been of the form

Eg.: $\hat{h}^2 \times Z_\mu \bar{f}_{L,R} \gamma^\mu f_{L,R}$



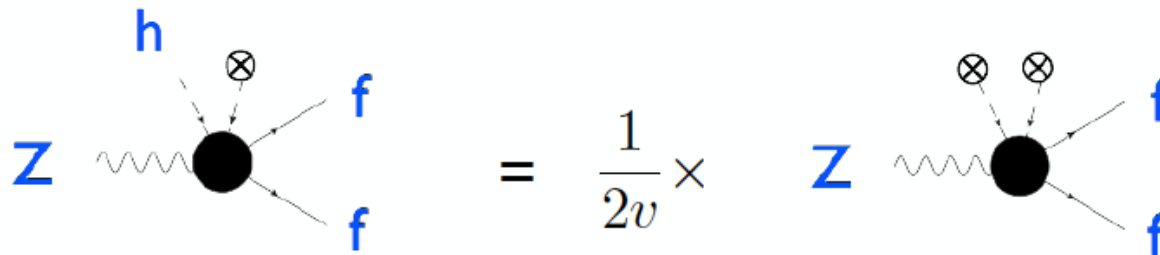
$$H^\dagger D_\mu H f \gamma^\mu f$$

Modifications in $h \rightarrow Zff$ related to $Z \rightarrow ff$

+ Relation to Higgs physics

All Primary deformations considered so far have been of the form

Eg.: $\hat{h}^2 \times Z_\mu \bar{f}_{L,R} \gamma^\mu f_{L,R}$



$$H^\dagger D_\mu H f \gamma^\mu f$$

Modifications in $h \rightarrow Zff$ related to $Z \rightarrow ff$

Contributions to Higgs physics already constrained by TGC, EWPT!

+ Any unconstrained BSM effect in Higgs physics?

- Yes! Deformations of the form:

$$|H|^2/\Lambda^2 \text{ (}\mathcal{O}_4\text{)} \longrightarrow \text{SM operator}$$

For $\hat{h} = v$ just a redefinition of SM parameters.

Eg.:

$$\overline{|H|^2 f_L H f_R + h.c.} \longrightarrow Y_f(\hat{h}) = Y_f + \delta Y_f \hat{h}^2/v^2 + \dots$$

$$\Delta \mathcal{L}_{ff}^h = \delta g_{ff}^h (h \bar{f}_L f_R + h.c.) \left(1 + \frac{3h}{2v} + \frac{h^2}{2v^2} \right)$$

+ Higgs Primary directions

$$e(\hat{h}), s_{\theta_W}(\hat{h}), g_s(\hat{h}), Y_f(\hat{h}), \lambda_h(\hat{h}), Z_h(\hat{h})$$

give

$$\begin{aligned} \Delta \mathcal{L}_{\gamma\gamma}^h &= \kappa_{\gamma\gamma} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[A_{\mu\nu} A^{\mu\nu} + Z_{\mu\nu} Z^{\mu\nu} + 2W_{\mu\nu}^+ W^{-\mu\nu} \right] \\ \Delta \mathcal{L}_{Z\gamma}^h &= \kappa_{Z\gamma} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[t_{\theta_W} A_{\mu\nu} Z^{\mu\nu} + \frac{c_{2\theta_W}}{2c_{\theta_W}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^+ W^{-\mu\nu} \right] \\ \Delta \mathcal{L}_{GG}^h &= \kappa_{GG} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) G_{\mu\nu}^A G^{A\mu\nu}, \\ \Delta \mathcal{L}_{ff}^h &= \delta g_{ff}^h \left(h \bar{f}_L f_R + \text{h.c.} \right) \left(1 + \frac{3h}{2v} + \frac{h^2}{2v^2} \right) \\ \Delta \mathcal{L}_{3h} &= \delta g_{3h} h^3 \left(1 + \frac{3h}{2v} + \frac{3h^2}{4v^2} + \frac{h^3}{8v^3} \right) \\ \Delta \mathcal{L}_{VV}^h &= \delta g_{VV}^h \left[h \left(W^{+\mu} W_{\mu}^- + \frac{Z^{\mu} Z_{\mu}}{2c_{\theta_W}^2} \right) + \Delta h \right] \end{aligned}$$

+ Remaining deformations:

- Other deformations include four-fermion deformations
CP- violating deformations and:

$$\Delta\mathcal{L}_R^W = \delta g_R^W \frac{\hat{h}^2}{v^2} W_\mu^+ \bar{u}_R \gamma^\mu d_R + \text{h.c.},$$

$$\Delta\mathcal{L}_{\text{dipole}}^V = \frac{Y_q \hat{h}}{m_W^2} \left[\delta\kappa_q^G \bar{q}_L T^A \sigma^{\mu\nu} q_R G_{\mu\nu}^A \right.$$

$$\left. + \delta\kappa_q^A (T_3 \bar{q}_L \sigma^{\mu\nu} q_R A_{\mu\nu} + \frac{s_{\theta_W}}{\sqrt{2}} \bar{u}_L \sigma^{\mu\nu} d_R W_{\mu\nu}^+) \right.$$

$$\left. + \delta\kappa_q^Z (T_3 \bar{q}_L \sigma^{\mu\nu} q_R Z_{\mu\nu} + \frac{c_{\theta_W}}{\sqrt{2}} \bar{u}_L \sigma^{\mu\nu} d_R W_{\mu\nu}^+) + \text{h.c.} \right]$$

$$\Delta\mathcal{L}_{3G} = \kappa_{3G} \epsilon_{ABC} G_\mu^{A\nu} G_{\nu\rho}^B G^{C\rho\mu}$$

MFV
suppressed

+ BSM Primaries

- 18 quantities **best** constrain all deformations in \mathcal{L}_6 .
- We call these **BSM Primaries**.

**Higgs (8)
Physics**

$$h \rightarrow \gamma\gamma, h \rightarrow \gamma Z, h \rightarrow gg$$

$$h \rightarrow VV, h \rightarrow ff, pp \rightarrow h^* \rightarrow hh$$

$$hA_{\mu\nu}A^{\mu\nu}, hA_{\mu\nu}Z^{\mu\nu}, hG_{\mu\nu}G^{\mu\nu}$$

$$hW^{+\mu}W_{\mu}^{-}, h\bar{f}f, h^3$$

**EWPT (7)
Data**

$$Z \rightarrow ff$$

(2 can be traded for S, T)

$$Z_{\mu}f_{L,R}\bar{\gamma}^{\mu}f_{L,R}$$

**TGC (3)
Data**

$$ee \rightarrow WW$$

$$g_1^Z c_{\theta_W} Z^{\mu} \left(W^{+\nu} \hat{W}_{\mu\nu}^{-} - W^{-\nu} \hat{W}_{\mu\nu}^{+} \right)$$

$$\kappa_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} W_{\mu}^{+} W_{\nu}^{-}$$

$$\lambda_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} \hat{W}_{\mu}^{-\rho} \hat{W}_{\rho\nu}^{+}$$

(18 is not considering four fermions and MFV suppressed and CPV deformations)

+

Higgs Primaries (8)

$$\Delta\mathcal{L}_{GG}^h = \kappa_{GG} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) G_{\mu\nu}^A G^{A\mu\nu}$$

$$\Delta\mathcal{L}_{ff}^h = \delta g_{ff}^h (h \bar{f}_L f_R + \text{h.c.}) \left(1 + \frac{3h}{2v} + \frac{h^2}{2v^2} \right)$$

$$\Delta\mathcal{L}_{3h} = \delta g_{3h} h^3 \left(1 + \frac{3h}{2v} + \frac{3h^2}{4v^2} + \frac{h^3}{8v^3} \right),$$

$$\begin{aligned} \Delta\mathcal{L}_{VV}^h = \delta g_{VV}^h & \left[h \left(W^{+\mu} W_{\mu}^{-} + \frac{Z^{\mu} Z_{\mu}}{2c_{\theta_w}^2} \right) \left(1 + \frac{2h}{v} \right. \right. \\ & \left. \left. + \frac{4h^2}{3v^2} + \frac{h^3}{3v^3} \right) + \frac{m_h^2}{12m_W^2} \left(\frac{h^4}{v} + \frac{3h^5}{4v^2} + \frac{h^6}{8v^3} \right) \right. \\ & \left. + \frac{m_f}{4m_W^2} \left(\frac{h^2}{v} + \frac{h^3}{3v^2} \right) (\bar{f}_L f_R + \text{h.c.}) \right], \end{aligned}$$

$$\begin{aligned} \Delta\mathcal{L}_{\gamma\gamma}^h = 4\kappa_{\gamma\gamma} s_{\theta_w}^2 & \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[A_{\mu\nu} A^{\mu\nu} \right. \\ & \left. + Z_{\mu\nu} Z^{\mu\nu} + 2W_{\mu\nu}^{+} W^{-\mu\nu} \right], \end{aligned}$$

$$\begin{aligned} \Delta\mathcal{L}_{Z\gamma}^h = 4\kappa_{Z\gamma} & \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[t_{\theta_w} A_{\mu\nu} Z^{\mu\nu} \right. \\ & \left. + \frac{c_{2\theta_w}}{2c_{\theta_w}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^{+} W^{-\mu\nu} \right]. \end{aligned}$$

EWPT Primaries(7)

$$\begin{aligned} \Delta\mathcal{L}_{ee}^V = \delta g_{eR}^Z \frac{\hat{h}^2}{v^2} & Z^{\mu} \bar{e}_R \gamma_{\mu} e_R \\ & + \delta g_{eL}^Z \frac{\hat{h}^2}{v^2} \left[Z^{\mu} \bar{e}_L \gamma_{\mu} e_L - \frac{c_{\theta_w}}{\sqrt{2}} (W^{+\mu} \bar{\nu}_L \gamma_{\mu} e_L + \text{h.c.}) \right] \\ & + \delta g_{\nu L}^Z \frac{\hat{h}^2}{v^2} \left[Z^{\mu} \bar{\nu}_L \gamma_{\mu} \nu_L + \frac{c_{\theta_w}}{\sqrt{2}} (W^{+\mu} \bar{\nu}_L \gamma_{\mu} e_L + \text{h.c.}) \right] \end{aligned}$$

$$\begin{aligned} \Delta\mathcal{L}_{qq}^V = \delta g_{uR}^Z \frac{\hat{h}^2}{v^2} & Z^{\mu} \bar{u}_R \gamma_{\mu} u_R + \delta g_{dR}^Z \frac{\hat{h}^2}{v^2} Z^{\mu} \bar{d}_R \gamma_{\mu} d_R \\ & + \delta g_{dL}^Z \frac{\hat{h}^2}{v^2} \left[Z^{\mu} \bar{d}_L \gamma_{\mu} d_L - \frac{c_{\theta_w}}{\sqrt{2}} (W^{+\mu} \bar{u}_L \gamma_{\mu} d_L + \text{h.c.}) \right] \\ & + \delta g_{uL}^Z \frac{\hat{h}^2}{v^2} \left[Z^{\mu} \bar{u}_L \gamma_{\mu} u_L + \frac{c_{\theta_w}}{\sqrt{2}} (W^{+\mu} \bar{u}_L \gamma_{\mu} d_L + \text{h.c.}) \right] \end{aligned}$$

TGC Primaries (3)

$$\begin{aligned} \Delta\mathcal{L}_{g_1^Z} = \delta g_1^Z c_{\theta_w}^2 \frac{\hat{h}^2}{v^2} & \left[\frac{e^2 \hat{h}^2}{4c_{\theta_w}^4} Z^{\mu} Z_{\mu} \right. \\ & \left. - g(W_{\mu}^{-} J_{-}^{\mu} + \text{h.c.}) - \frac{gc_{2\theta_w}}{c_{\theta_w}^3} Z_{\mu} J_Z^{\mu} - 2et_{\theta_w} Z_{\mu} J_{em}^{\mu} \right] \\ \Delta\mathcal{L}_{\kappa\gamma} = \frac{\delta\kappa_{\gamma}}{v^2} & \left[ie\hat{h}^2 (A_{\mu\nu} - t_{\theta_w} Z_{\mu\nu}) W^{+\mu} W^{-\nu} \right. \\ & \left. + Z_{\nu} \partial_{\mu} \hat{h}^2 (t_{\theta_w} A^{\mu\nu} - t_{\theta_w}^2 Z^{\mu\nu}) + \frac{(\hat{h}^2 - v^2)}{2} \right. \\ & \left. \times (t_{\theta_w} Z_{\mu\nu} A^{\mu\nu} + \frac{c_{2\theta_w}}{c_{\theta_w}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^{+} W^{-\mu\nu}) \right] \\ \Delta\mathcal{L}_{\lambda\gamma} = \frac{i\lambda_{\gamma}}{m_W^2} & [(eA^{\mu\nu} + gc_{\theta_w} Z^{\mu\nu}) W_{\nu}^{-\rho} W_{\rho\mu}^{+}] \end{aligned}$$

+ Dimension 6 lagrangian

- So we have finally **constructed the dim-6 lagrangian** in a **bottom up way** (not starting from operators but from measurable deformations):

$$\begin{aligned}\Delta\mathcal{L}_{\text{BSM}} = & \Delta\mathcal{L}_{\gamma\gamma}^h + \Delta\mathcal{L}_{Z\gamma}^h + \Delta\mathcal{L}_{GG}^h + \Delta\mathcal{L}_{ff}^h + \Delta\mathcal{L}_{3h} + \Delta\mathcal{L}_{VV}^h + \Delta\mathcal{L}_{ee}^V + \Delta\mathcal{L}_{qq}^V \\ & + \Delta\mathcal{L}_{g_1^Z} + \Delta\mathcal{L}_{\kappa\gamma} + \Delta\mathcal{L}_{\lambda\gamma} + \Delta\mathcal{L}_{3G} + \Delta\mathcal{L}_{4f} + \Delta\mathcal{L}_{\text{MFV}}^V + \Delta\mathcal{L}_{\text{CPV}}.\end{aligned}$$

- All physical processes, eg. $h \rightarrow Vff$, $pp \rightarrow Vh$, $VV \rightarrow h$ etc can be computed **as a function** of the **BSM primary parameters** using the above Lagrangian.

+ Dimension 6 lagrangian

- So we have finally **constructed the dim-6 lagrangian** in a **bottom up way** (not starting from operators but from measurable deformations):

$$\Delta\mathcal{L}_{\text{BSM}} = \Delta\mathcal{L}_{\gamma\gamma}^h + \Delta\mathcal{L}_{Z\gamma}^h + \Delta\mathcal{L}_{GG}^h + \Delta\mathcal{L}_{ff}^h + \Delta\mathcal{L}_{3h} + \Delta\mathcal{L}_{VV}^h + \Delta\mathcal{L}_{ee}^V + \Delta\mathcal{L}_{qq}^V \\ + \Delta\mathcal{L}_{g_1^Z} + \Delta\mathcal{L}_{\kappa\gamma} + \Delta\mathcal{L}_{\lambda\gamma} + \Delta\mathcal{L}_{3G} + \Delta\mathcal{L}_{4f} + \Delta\mathcal{L}_{\text{MFV}}^V + \Delta\mathcal{L}_{\text{CPV}}.$$

18 Primary Directions

- All physical processes, eg. $h \rightarrow Vff$, $pp \rightarrow Vh$, $VV \rightarrow h$ etc can be computed **as a function** of the **BSM primary parameters** using the above Lagrangian.

+ Dimension 6 lagrangian

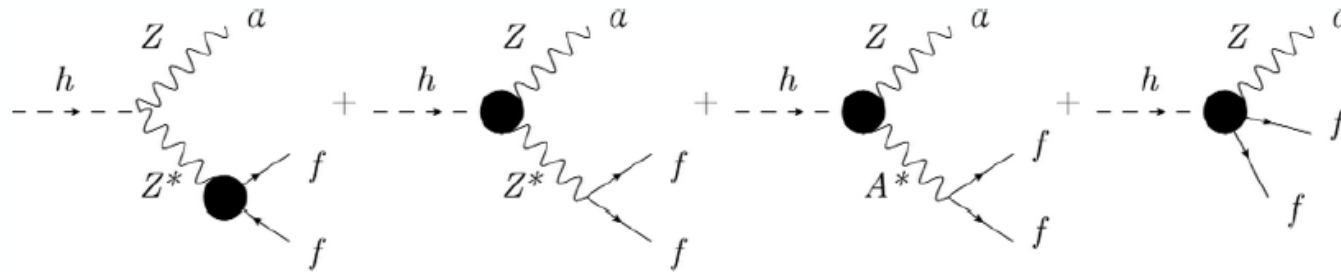
- So we have finally **constructed the dim-6 lagrangian** in a **bottom up way** (not starting from operators but from measurable deformations):

$$\Delta\mathcal{L}_{\text{BSM}} = \Delta\mathcal{L}_{\gamma\gamma}^h + \Delta\mathcal{L}_{Z\gamma}^h + \Delta\mathcal{L}_{GG}^h + \Delta\mathcal{L}_{ff}^h + \Delta\mathcal{L}_{3h} + \Delta\mathcal{L}_{VV}^h + \Delta\mathcal{L}_{ee}^V + \Delta\mathcal{L}_{qq}^V \\ + \Delta\mathcal{L}_{g_1^Z} + \Delta\mathcal{L}_{\kappa\gamma} + \Delta\mathcal{L}_{\lambda\gamma} + \Delta\mathcal{L}_{3G} + \Delta\mathcal{L}_{4f} + \Delta\mathcal{L}_{\text{MFV}}^V + \Delta\mathcal{L}_{\text{CPV}}.$$

Total:59 Primary Directions

- All physical processes, eg. $h \rightarrow Vff$, $pp \rightarrow Vh$, $VV \rightarrow h$ etc can be computed **as a function** of the **BSM primary parameters** using the above Lagrangian.

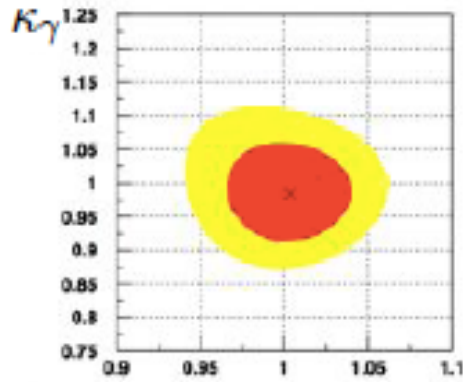
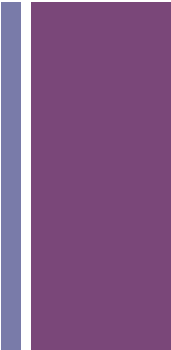
+ Example: $h \rightarrow Zff$



- The relevant primaries (ignoring ones constrained at per-mille level) are:

$$\begin{aligned}
 \Delta\mathcal{L}_{Z\gamma}^h &= 4\kappa_{Z\gamma} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[t_{\theta_W} A_{\mu\nu} Z^{\mu\nu} + \frac{c_{2\theta_W}}{2c_{\theta_W}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^+ W^{-\mu\nu} \right] \\
 \Delta\mathcal{L}_{g_1^Z} &= \delta g_1^Z c_{\theta_W}^2 \frac{\hat{h}^2}{v^2} \left[\frac{e^2 \hat{h}^2}{4c_{\theta_W}^4} Z^\mu Z_\mu - g(W_\mu^- J_-^\mu + \text{h.c.}) - \frac{gc_{2\theta_W}}{c_{\theta_W}^3} Z_\mu J_Z^\mu - 2et_{\theta_W} Z_\mu J_{em}^\mu \right] \\
 \Delta\mathcal{L}_{\kappa\gamma} &= \frac{\delta\kappa_\gamma}{v^2} \left[ie\hat{h}^2 (A_{\mu\nu} - t_{\theta_W} Z_{\mu\nu}) W^{+\mu} W^{-\nu} \right. \\
 &\quad + Z_\nu \partial_\mu \hat{h}^2 (t_{\theta_W} A^{\mu\nu} - t_{\theta_W}^2 Z^{\mu\nu}) + \frac{(\hat{h}^2 - v^4)}{2} \\
 &\quad \left. \times (t_{\theta_W} Z_{\mu\nu} A^{\mu\nu} + \frac{c_{2\theta_W}}{2c_{\theta_W}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^+ W^{-\mu\nu}) \right] \quad (24)
 \end{aligned}$$

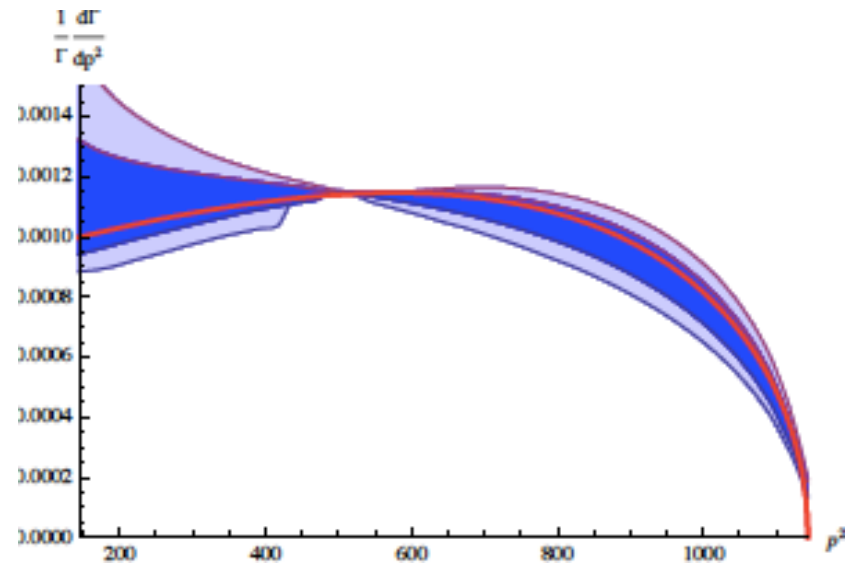
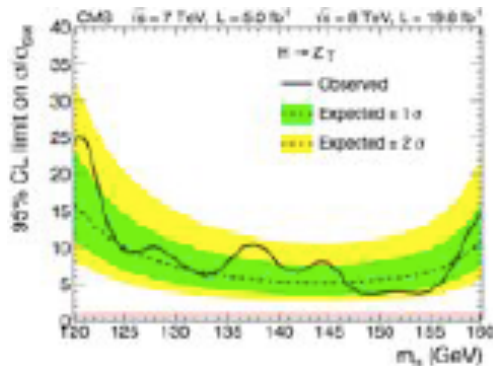
+ Example: $h \rightarrow Zff$



δg_1^Z

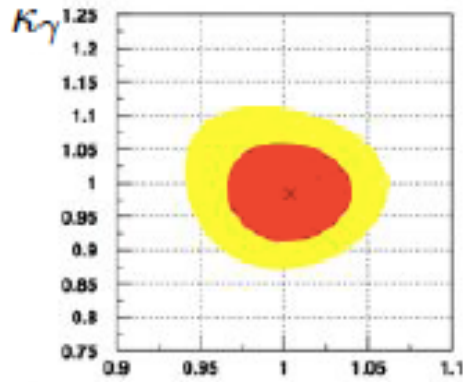
+

$h \rightarrow \gamma Z$



Gupta, Liu, Pomarol, Riva (in preparation)

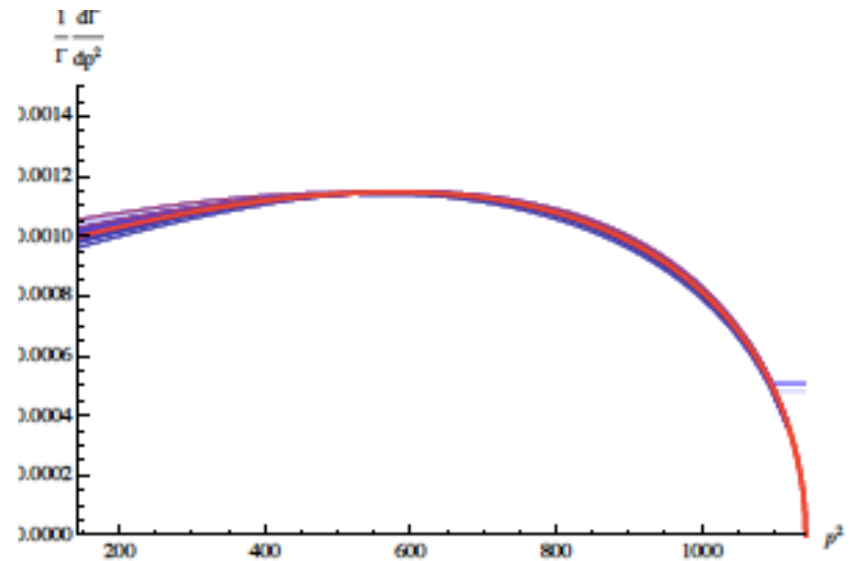
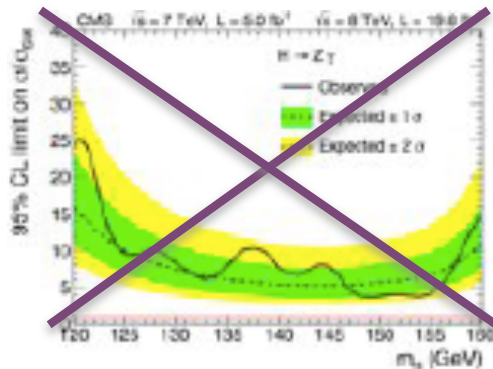
+ Example: $h \rightarrow Zff$



δg_1^Z

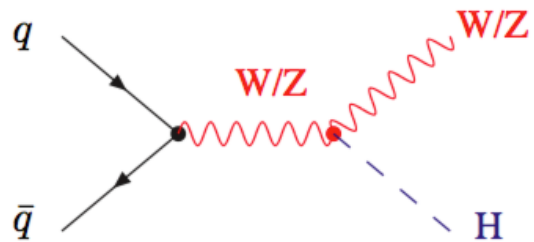
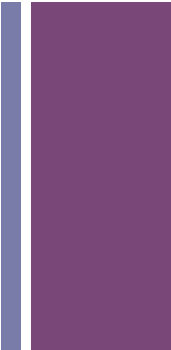
+

$h \rightarrow \gamma Z$



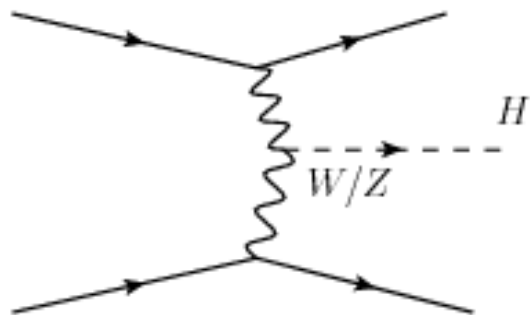
Gupta, Liu, Pomarol, Riva (in preparation)

+ Prediction for any BSM Process in terms of BSM primaries (at dim-6)



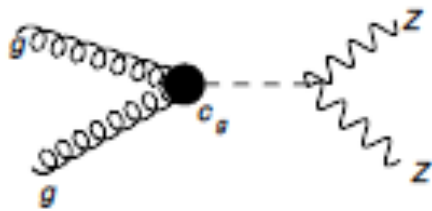
$$hA_{\mu\nu}A^{\mu\nu}, hA_{\mu\nu}Z^{\mu\nu}, hG_{\mu\nu}G^{\mu\nu}$$

$$hW^{+\mu}W_{\mu}^{-}, hf\bar{f}, h^3$$



as a function of:

$$Z_{\mu}\bar{f}_{L,R}\gamma^{\mu}f_{L,R}$$



etc

$$g_1^Z c_{\theta_W} Z^{\mu} \left(W^{+\nu} \hat{W}_{\mu\nu}^{-} - W^{-\nu} \hat{W}_{\mu\nu}^{+} \right)$$

$$\kappa_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} W_{\mu}^{+} W_{\nu}^{-}$$

$$\lambda_{\gamma} s_{\theta_W} \hat{A}^{\mu\nu} \hat{W}_{\mu}^{-\rho} \hat{W}_{\rho\nu}^{+}$$

Gupta, Liu, Pomarol, Riva (in preparation)

+ A Hierarchy of Constraints

$$\left. \delta g_{ff}^h \quad \delta g_{3h} \quad \delta g_{VV}^h \right\} \mathcal{O}(1)$$

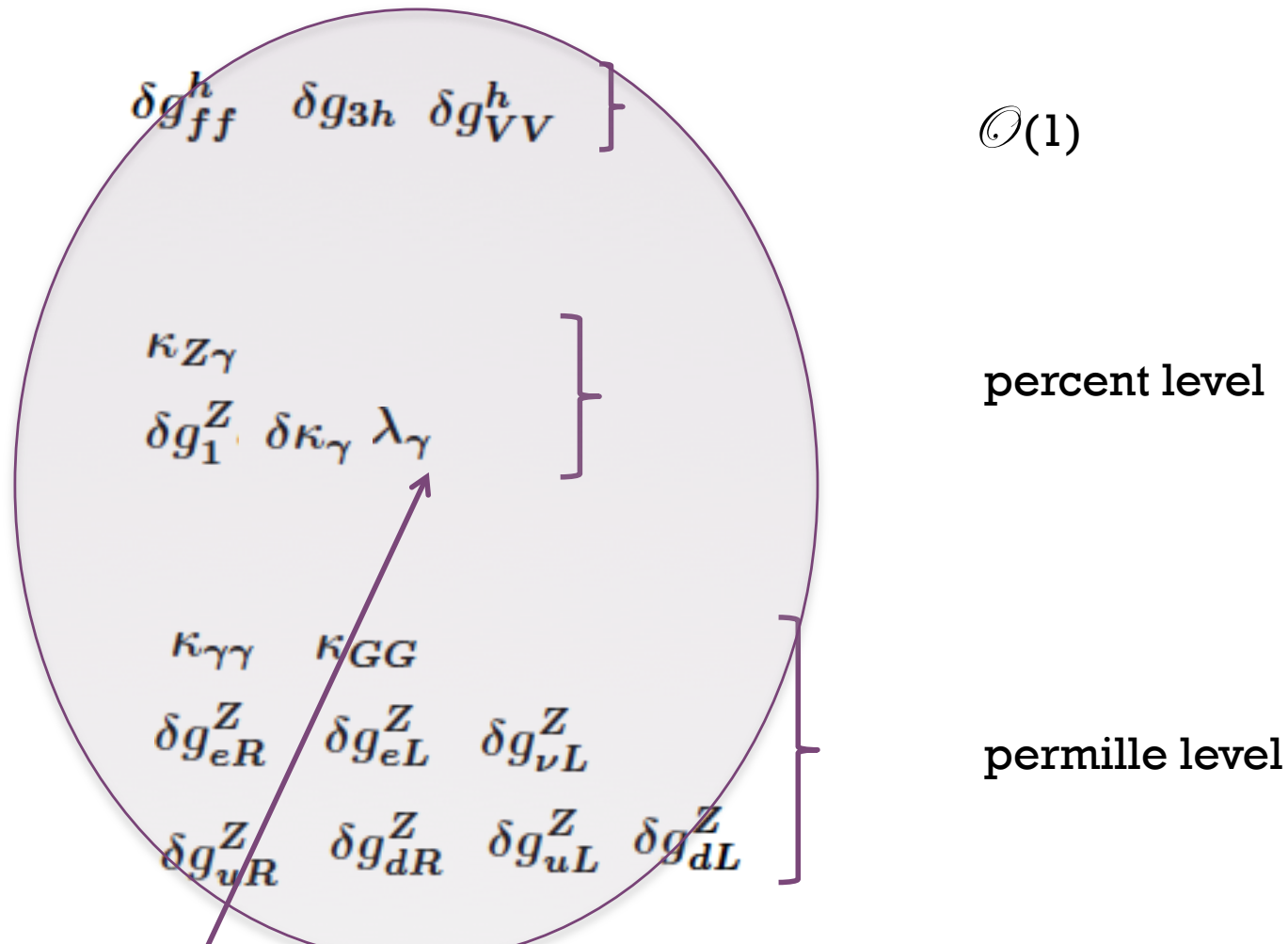
$$\left. \begin{array}{l} \kappa_{Z\gamma} \\ \delta g_1^Z, \delta \kappa_\gamma, \lambda_\gamma \end{array} \right\} \text{percent level}$$

$$\left. \begin{array}{l} \kappa_{\gamma\gamma} \quad \kappa_{GG} \\ \delta g_{eR}^Z \quad \delta g_{eL}^Z \quad \delta g_{\nu L}^Z \\ \delta g_{uR}^Z \quad \delta g_{dR}^Z \quad \delta g_{uL}^Z \quad \delta g_{dL}^Z \end{array} \right\} \text{permille level}$$

These parameters can be identified with the Wilson coefficients of dim-6 operators c_i (mw).

(Pomarol & Riva 2013)

+ A Hierarchy of Constraints



These parameters can be identified with the Wilson coefficients of dim-6 operators c_i (mw).

(Pomarol & Riva 2013)

+ RG-induced Constraints (diphoton example)



BSM matching scale Λ

$c_1(\Lambda), c_2(\Lambda), \dots, c_i(\Lambda)$

Theoretically important;
To constrain these need to
know RG running.

$c_1(m_W), c_2(m_W), \dots, c_i(m_W)$

Directly **constrained by**
experiments

Experimental Observable scale $m_H \sim m_W$

Jenkins, Grojean, Manohar, Trott (2013)
Elias-Miro, Espinosa, Masso, Pomarol (2013)

+ RG-induced Constraints (diphoton example)

BSM matching scale Λ

$c_1(\Lambda), c_2(\Lambda), \dots, c_i(\Lambda)$

Theoretically important;
To constrain these need to
know RG running.

RG running and mixing

for eg. take the diphoton operator:

$$\hat{c}_{\gamma\gamma}(m_h) = \hat{c}_{\gamma\gamma}(\Lambda) - \frac{1}{16\pi^2} \left[\left(\frac{3}{2}g^2 - 2\lambda \right) \hat{c}_{\kappa\gamma} + 3g^2 \hat{c}_{\lambda\gamma} \right] \log \left(\frac{\Lambda}{m_h} \right) < \epsilon_{h\gamma\gamma}$$

$c_1(m_W), c_2(m_W), \dots, c_i(m_W)$

Directly **constrained by**
experiments

Experimental Observable scale $m_H \sim m_W$

Jenkins, Grojean, Manohar, Trott (2013)

Elias-Miro, Espinosa, Masso, Pomarol (2013)



- But aren't these effects one loop suppressed and thus unimportant ?

+ RG-induced Constraints (diphoton example)

$$\hat{c}_{\gamma\gamma}(m_h) = \hat{c}_{\gamma\gamma}(\Lambda) - \frac{1}{16\pi^2} \left[\left(\frac{3}{2}g^2 - 2\lambda \right) \hat{c}_{\kappa\gamma} + 3g^2 \hat{c}_{\lambda\gamma} \right] \log \left(\frac{\Lambda}{m_h} \right) < \epsilon_{h\gamma\gamma}$$

One loop suppression

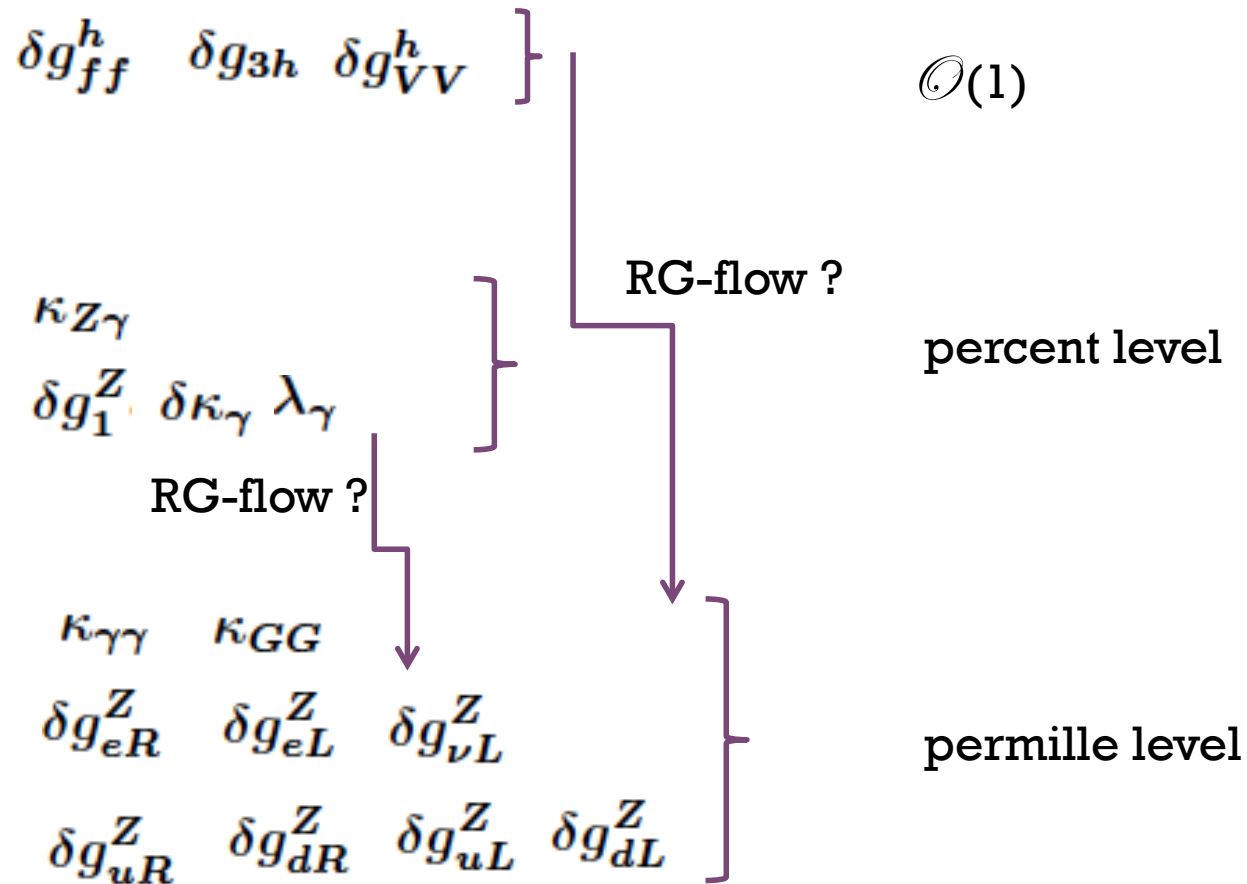
Constrained **per mille** level

Assuming **no tuning/correlation** between the RHS contributions we derive **RG-induced bounds**:

Constrained only at 10 % level thus allowed to be much larger than bound on $h\gamma\gamma$. This and the log enhancement can compensate for the loop factor.

$$|\hat{c}_{\kappa\gamma}| < \Delta_{FT} \frac{16\pi^2}{\log(\Lambda/m_h)} \left| \left(\frac{3}{2}g^2 - 2\lambda \right)^{-1} \right| \epsilon_{h\gamma\gamma}, \quad |\hat{c}_{\lambda\gamma}| < \Delta_{FT} \frac{16\pi^2}{\log(\Lambda/m_h)} \left| \frac{1}{3g^2} \right| \epsilon_{h\gamma\gamma}$$

+ A Hierarchy of Constraints

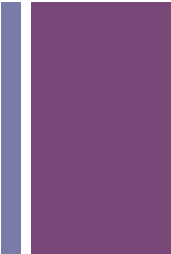


These parameters can be identified with the Wilson coefficients of dim-6 operators c_i (mw).

(Pomarol & Riva 2013)



Anomalous Dimensional Matrix



	\hat{c}_S	\hat{c}_T	\hat{c}_Y	\hat{c}_W	$\hat{c}_{\gamma\gamma}$
$\gamma_{\hat{c}_S}$	$\frac{1}{3}g'^2 + 6y_t^2$	$-\frac{g'^2}{2}$	$\frac{1}{8}g'^2 \left(147 - 106\frac{g'^2}{g^2}\right)$	$\frac{1}{8}(77g^2 + 58g'^2)$	$16e^2$
$\gamma_{\hat{c}_T}$	$-9g'^2 - 24t_{\theta_W}^2 \lambda$	$\frac{9}{2}g^2 + 12y_t^2 + 12\lambda$	$\frac{9}{2}g'^2 + 12t_{\theta_W}^2 (g'^2 + \lambda)$	$\frac{9}{2}g'^2$	0
$\gamma_{\hat{c}_Y}$	$-\frac{2}{3}g'^2$	0	$\frac{94}{3}g'^2$	0	0
$\gamma_{\hat{c}_W}$	0	0	$\frac{53}{12}g'^2 \left(1 - 3t_{\theta_W}^2\right)$	$\frac{331}{12}g^2 + \frac{29}{4}g'^2$	0
$\gamma_{\hat{c}_{\gamma\gamma}}$	0	0	0	0	$-\frac{9}{2}g^2 - \frac{3}{2}g'^2 + 6y_t^2 + 12\lambda$
$\gamma_{\hat{c}_H}$	$18g'^2 - t_{\theta_W}^2 (9g'^2 + 24\lambda)$	$-9g^2 + \frac{9}{2}g'^2 + 12\lambda$	$t_{\theta_W}^2 \left(-\frac{141}{4}g'^2 + 12\lambda\right)$	$\frac{63}{2}g^2 + \frac{51}{4}g'^2 + 72\lambda$	0
$\gamma_{\hat{c}_{\gamma Z}}$	0	0	0	0	0
$\gamma_{\hat{c}_{kZ}}$	0	0	0	0	$-16e^2$
$\gamma_{\hat{c}_{gZ}}$	$-\frac{g'^2}{6c_{\theta_W}^2}$	$\frac{g'^2}{12c_{\theta_W}^2}$	$\frac{g'^2}{8c_{\theta_W}^2} (106t_{\theta_W}^2 - 29)$	$-\frac{1}{8c_{\theta_W}^2} (79g^2 + 58g'^2)$	0
$\gamma_{\hat{c}_{\lambda\gamma}}$	0	0	0	0	0

	\hat{c}_H	$\hat{c}_{\gamma Z}$	$\hat{c}_{\kappa\gamma}$	\hat{c}_{gZ}	$\hat{c}_{\lambda\gamma}$
$\gamma_{\hat{c}_S}$	$-\frac{1}{6}g'^2$	$4(g^2 - g'^2)$	$-\frac{11}{2}g^2 - \frac{1}{6}g'^2 - 4\lambda$	$c_{\theta_W}^2 (9g^2 - \frac{1}{3}g'^2)$	$-2g^2$
$\gamma_{\hat{c}_T}$	$\frac{3}{2}g'^2$	0	$-9g'^2 - 24t_{\theta_W}^2 \lambda$	$24s_{\theta_W}^2 \lambda$	0
$\gamma_{\hat{c}_Y}$	0	0	$-\frac{2}{3}g'^2$	$\frac{2}{3}e^2$	0
$\gamma_{\hat{c}_W}$	0	0	0	$-\frac{2}{3}c_{\theta_W}^2 g^2$	0
$\gamma_{\hat{c}_{\gamma\gamma}}$	0	0	$\frac{3}{2}g^2 - 2\lambda$	0	$3g^2$
$\gamma_{\hat{c}_H}$	$-\frac{9}{2}g^2 - 3g'^2 + 12y_t^2 + 24\lambda$	0	$9g'^2(2 - t_{\theta_W}^2) - 24t_{\theta_W}^2 \lambda$	$9(g'^2 s_{\theta_W}^2 - g^2 c_{\theta_W}^2) - 24\lambda(6c_{\theta_W}^2 - s_{\theta_W}^2)$	0
$\gamma_{\hat{c}_{\gamma Z}}$	0	$-\frac{7}{2}g^2 - \frac{1}{2}g'^2 + 6y_t^2 + 12\lambda$	$c_{\theta_W}^2 (2g^2 - 2\lambda) - s_{\theta_W}^2 (g^2 - 2\lambda)$	0	$\frac{g^2}{2} (11c_{\theta_W}^2 - s_{\theta_W}^2)$
$\gamma_{\hat{c}_{\kappa\gamma}}$	0	$4(g^2 - g'^2)$	$\frac{11}{2}g^2 + \frac{g'^2}{2} + 6y_t^2 + 4\lambda$	0	$2g^2$
$\gamma_{\hat{c}_{gZ}}$	$\frac{g'^2}{12c_{\theta_W}^2}$	0	$\frac{g'^2}{6c_{\theta_W}^2}$	$\frac{17}{2}g^2 - \frac{g'^2}{6} + 6y_t^2$	0
$\gamma_{\hat{c}_{\lambda\gamma}}$	0	0	0	0	$\frac{53}{3}g^2$

+

Anomalous Dimensional Matrix

$$(\hat{c}_S, \hat{c}_T, \hat{c}_Y, \hat{c}_W, \hat{c}_{\gamma\gamma}, \hat{c}_{\gamma Z}, \hat{c}_{\kappa\gamma}, \hat{c}_{gz}, \hat{c}_{\lambda\gamma}, \hat{c}_H)^t (m_t) \simeq \quad (4)$$

$$\begin{array}{c}
 \mathbb{R} \\
 \left(\begin{array}{ccccccccccc}
 0.9 & 0.003 & -0.03 & -0.08 & -0.02 & -0.02 & -0.04 & 0.05 & -0.01 & 0.001 \\
 0.03 & 0.8 & -0.02 & -0.009 & 0 & 0 & -0.03 & 0.01 & 0 & -0.003 \\
 0.001 & 0 & 0.9 & 0 & 0 & 0 & -0.001 & 0.001 & 0 & 0 \\
 0 & 0 & -0.001 & 0.8 & 0 & 0 & 0 & -0.003 & 0 & 0 \\
 0 & 0 & 0 & 0 & 0.9 & 0 & 0.006 & 0 & 0.02 & 0 \\
 0 & 0 & 0 & 0 & 0 & 0.9 & 0.007 & 0 & 0.03 & 0 \\
 0 & 0 & 0 & 0 & -0.02 & -0.02 & 0.9 & 0 & -0.01 & 0 \\
 0.0004 & -0.0007 & -0.0004 & 0.1 & 0 & 0 & -0.0004 & 0.9 & 0 & -0.0007 \\
 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0.9 & 0 \\
 -0.02 & 0.03 & 0.01 & -0.4 & 0 & 0 & 0.02 & -0.3 & 0 & 0.8
 \end{array} \right) \begin{array}{l}
 \hat{c}_S(\Lambda) \\
 \hat{c}_T(\Lambda) \\
 \hat{c}_Y(\Lambda) \\
 \hat{c}_W(\Lambda) \\
 \hat{c}_{\gamma\gamma}(\Lambda) \\
 \hat{c}_{\gamma Z}(\Lambda) \\
 \hat{c}_{\kappa\gamma}(\Lambda) \\
 \hat{c}_{gz}(\Lambda) \\
 \hat{c}_{\lambda\gamma}(\Lambda) \\
 \hat{c}_H(\Lambda)
 \end{array}
 \end{array}$$

- We focus on the part of the matrix, where **weakly bound couplings contribute to strongly bound couplings.**

+ Numerical Results

Coupling	Direct Constraint	RG-induced Constraint
$\hat{c}_S(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_T(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_Y(m_t)$	$[-3, 3] \times 10^{-3}$ [22]	-
$\hat{c}_W(m_t)$	$[-2, 2] \times 10^{-3}$ [22]	-
$\hat{c}_{\gamma\gamma}(m_t)$	$[-1, 2] \times 10^{-3}$ [18]	-
$\hat{c}_{\gamma Z}(m_t)$	$[-0.6, 1] \times 10^{-2}$ [18]	$[-2, 6] \times 10^{-2}$
$\hat{c}_{\kappa\gamma}(m_t)$	$[-10, 7] \times 10^{-2}$ [27]	$[-5, 2] \times 10^{-2}$
$\hat{c}_{gZ}(m_t)$	$[-4, 2] \times 10^{-2}$ [27]	$[-3, 1] \times 10^{-2}$
$\hat{c}_{\lambda\gamma}(m_t)$	$[-6, 2] \times 10^{-2}$ [27]	$[-2, 8] \times 10^{-2}$
$\hat{c}_H(m_t)$	$[-6, 5] \times 10^{-1}$ [32]	$[-2, 0.5] \times 10^{-1}$

- We assume that there is **no tuning/correlation among different contributions** so that **each RG-induced** term in the RGE is **smaller than the bound**. This gives us new **RG-induced constraints**.
- We get bounds on some **TGC** and on **C_H** mainly from their RG-induced contribution to **{S, T, W, Y}** that are **stronger** than the direct bounds.

+ Numerical Results

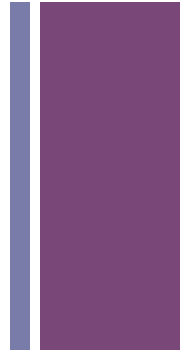
Coupling	Direct Constraint	RG-induced Constraint
$\hat{c}_S(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_T(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_Y(m_t)$	$[-3, 3] \times 10^{-3}$ [22]	-
$\hat{c}_W(m_t)$	$[-2, 2] \times 10^{-3}$ [22]	-
$\hat{c}_{\gamma\gamma}(m_t)$	$[-1, 2] \times 10^{-3}$ [18]	-
$\hat{c}_{\gamma Z}(m_t)$	$[-0.6, 1] \times 10^{-2}$ [18]	$[-2, 6] \times 10^{-2}$
$\hat{c}_{\kappa\gamma}(m_t)$	$[-10, 7] \times 10^{-2}$ [27]	$[-5, 2] \times 10^{-2}$
$\hat{c}_{gZ}(m_t)$	$[-4, 2] \times 10^{-2}$ [27]	$[-3, 1] \times 10^{-2}$
$\hat{c}_{\lambda\gamma}(m_t)$	$[-6, 2] \times 10^{-2}$ [27]	$[-2, 8] \times 10^{-2}$
$\hat{c}_H(m_t)$	$[-6, 5] \times 10^{-1}$ [32]	$[-2, 0.5] \times 10^{-1}$

- We assume that there is **no tuning** so that **each RG-induced** term in the RGE is **smaller than the bound**. This gives us new **RG-induced constraints**.
- We get bounds on some **TGC** and on **C_H** mainly from their RG-induced contribution to **{S, T, W, Y}** that are **stronger** than the direct bounds.

+ Summary

- We present an **efficient choice of independent primary BSM deformations**. All other deformations are generated in a correlated way and we derive these correlations.
- Barring 4-fermions and CPV and MFV deformations, there are 18 BSM Primary effects: **8 Higgs primaries, 7 EWPT primaries and 3 TGC Primaries**.
- **Predictions: W coupling deviations not independent of Z coupling deviations. All $hVff$ deformations constrained by EWPT and TGC; all QGC constrained by the g_{1Z} TGC.**
- We find that RG-induced constraints on the **hVV and TGC primaries due to mixing with the $H\gamma\gamma$ and S -parameter primary directions can be stronger to (or of the same order as) tree level constraints.**

+ Back Up slides



+ Part II: RG-induced constraints

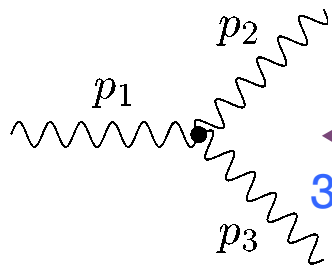
- We calculate the **one-loop anomalous dimension matrix** for 13 bosonic dimension-6 operators relevant for **electroweak (including TGC) and Higgs physics**.
- **New RG-induced bounds**, stronger than the direct constraints, on a **universal shift of the Higgs couplings and the anomalous triple gauge couplings**.

+

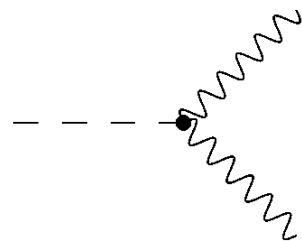
Operators to Observables

- The 10 EW & Higgs operators can be related to 10 observables:

4 precision observables



3 triple gauge coupling observables



3 Higgs observables

$$\sum c_i \mathcal{O}_i$$

$$\hat{T} = \hat{c}_T(m_W) = \frac{v^2}{\Lambda^2} c_T(m_W)$$

$$\hat{S} = \hat{c}_S(m_W) = \frac{m_W^2}{\Lambda^2} [c_W(m_W) + c_B(m_W) + 4c_{WB}(m_W)],$$

$$Y = \hat{c}_Y(m_W) = \frac{m_W^2}{\Lambda^2} c_{2B}(m_W)$$

$$W = \hat{c}_W(m_W) = \frac{m_W^2}{\Lambda^2} c_{2W}(m_W)$$

$$\delta g_1^Z \equiv \hat{c}_{g^Z}(m_W) = \frac{m_W^2}{\Lambda^2} \frac{1}{c_{\theta_W}^2} c_W(m_W)$$

$$\delta \kappa_\gamma \equiv \hat{c}_{\kappa_\gamma}(m_W) = \frac{m_W^2}{\Lambda^2} (-4c_{WB}(m_W))$$

$$\lambda_Z \equiv \hat{c}_{\lambda_\gamma}(m_W) = \frac{m_W^2}{\Lambda^2} c_{3W}(m_W)$$

$$\hat{c}_H(m_h) = \frac{v^2}{\Lambda^2} c_H(m_h),$$

$$\hat{c}_{\gamma\gamma}(m_h) = \frac{m_W^2}{\Lambda^2} (c_{BB}(m_h) + c_{WW}(m_h) - c_{WB}(m_h))$$

$$\hat{c}_{\gamma Z}(m_h) =$$

$$\frac{m_W^2}{\Lambda^2} (2c_{\theta_W}^2 c_{WW}(m_h) - 2s_{\theta_W}^2 c_{BB}(m_h) - (c_{\theta_W}^2 - s_{\theta_W}^2) c_{WB}(m_h)).$$

(Pomarol & Riva 2013)

+ A Hierarchy of Constraints

$$\left. \delta g_{ff}^h \quad \delta g_{3h} \quad \delta g_{VV}^h \right\} \mathcal{O}(1)$$

$$\left. \begin{array}{l} \kappa_{Z\gamma} \\ \delta g_1^Z, \delta \kappa_\gamma, \lambda_\gamma \end{array} \right\} \text{percent level}$$

$$\left. \begin{array}{l} \kappa_{\gamma\gamma} \quad \kappa_{GG} \\ \delta g_{eR}^Z \quad \delta g_{eL}^Z \quad \delta g_{\nu L}^Z \\ \delta g_{uR}^Z \quad \delta g_{dR}^Z \quad \delta g_{uL}^Z \quad \delta g_{dL}^Z \end{array} \right\} \text{permille level}$$

+ Anomalous Dimensional Matrix

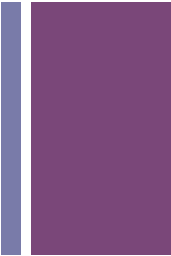
$$\hat{c}_i(m_h) = \hat{c}_i(\Lambda) - \frac{1}{16\pi^2} \hat{\gamma}_{ij} \hat{c}_j(\Lambda) \log \left(\frac{\Lambda}{m_h} \right)$$

Anomalous dimension matrix: connects Wilson coefficients at BSM scale to those at experimental scale.

- The full anomalous dimension matrix is shown in the next slide:



Anomalous Dimensional Matrix

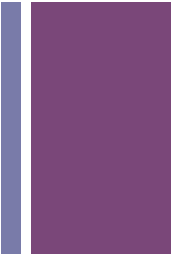


	\hat{c}_S	\hat{c}_T	\hat{c}_Y	\hat{c}_W	$\hat{c}_{\gamma\gamma}$
$\gamma_{\hat{c}_S}$	$\frac{1}{3}g'^2 + 6y_t^2$	$-\frac{g^2}{2}$	$\frac{1}{8}g'^2 \left(147 - 106\frac{g'^2}{g^2}\right)$	$\frac{1}{8}(77g^2 + 58g'^2)$	$16e^2$
$\gamma_{\hat{c}_T}$	$-9g'^2 - 24t_{\theta_W}^2 \lambda$	$\frac{9}{2}g^2 + 12y_t^2 + 12\lambda$	$\frac{9}{2}g'^2 + 12t_{\theta_W}^2 (g'^2 + \lambda)$	$\frac{9}{2}g'^2$	0
$\gamma_{\hat{c}_Y}$	$-\frac{2}{3}g'^2$	0	$\frac{94}{3}g'^2$	0	0
$\gamma_{\hat{c}_W}$	0	0	$\frac{53}{12}g'^2 \left(1 - 3t_{\theta_W}^2\right)$	$\frac{331}{12}g^2 + \frac{29}{4}g'^2$	0
$\gamma_{\hat{c}_{\gamma\gamma}}$	0	0	0	0	$-\frac{9}{2}g^2 - \frac{3}{2}g'^2 + 6y_t^2 + 12\lambda$
$\gamma_{\hat{c}_H}$	$18g'^2 - t_{\theta_W}^2 (9g'^2 + 24\lambda)$	$-9g^2 + \frac{9}{2}g'^2 + 12\lambda$	$t_{\theta_W}^2 \left(-\frac{141}{4}g'^2 + 12\lambda\right)$	$\frac{63}{2}g^2 + \frac{51}{4}g'^2 + 72\lambda$	0
$\gamma_{\hat{c}_{\gamma Z}}$	0	0	0	0	0
$\gamma_{\hat{c}_{kZ}}$	0	0	0	0	$-16e^2$
$\gamma_{\hat{c}_{gZ}}$	$-\frac{g^2}{6c_{\theta_W}^2}$	$\frac{g^2}{12c_{\theta_W}^2}$	$\frac{g^2}{8c_{\theta_W}^2} (106t_{\theta_W}^2 - 29)$	$-\frac{1}{8c_{\theta_W}^2} (79g^2 + 58g'^2)$	0
$\gamma_{\hat{c}_{\lambda\gamma}}$	0	0	0	0	0

	\hat{c}_H	$\hat{c}_{\gamma Z}$	$\hat{c}_{\kappa\gamma}$	\hat{c}_{gZ}	$\hat{c}_{\lambda\gamma}$
$\gamma_{\hat{c}_S}$	$-\frac{1}{6}g'^2$	$4(g^2 - g'^2)$	$-\frac{11}{2}g^2 - \frac{1}{6}g'^2 - 4\lambda$	$c_{\theta_W}^2 \left(9g^2 - \frac{1}{3}g'^2\right)$	$-2g^2$
$\gamma_{\hat{c}_T}$	$\frac{3}{2}g'^2$	0	$-9g'^2 - 24t_{\theta_W}^2 \lambda$	$24s_{\theta_W}^2 \lambda$	0
$\gamma_{\hat{c}_Y}$	0	0	$-\frac{2}{3}g'^2$	$\frac{2}{3}e^2$	0
$\gamma_{\hat{c}_W}$	0	0	0	$-\frac{2}{3}c_{\theta_W}^2 g^2$	0
$\gamma_{\hat{c}_{\gamma\gamma}}$	0	0	$\frac{3}{2}g^2 - 2\lambda$	0	$3g^2$
$\gamma_{\hat{c}_H}$	$-\frac{9}{2}g^2 - 3g'^2 + 12y_t^2 + 24\lambda$	0	$9g'^2(2 - t_{\theta_W}^2) - 24t_{\theta_W}^2 \lambda$	$9(g^2 s_{\theta_W}^2 - g^2 c_{\theta_W}^2) - 24\lambda(6c_{\theta_W}^2 - s_{\theta_W}^2)$	0
$\gamma_{\hat{c}_{\gamma Z}}$	0	$-\frac{7}{2}g^2 - \frac{1}{2}g'^2 + 6y_t^2 + 12\lambda$	$c_{\theta_W}^2 (2g^2 - 2\lambda) - s_{\theta_W}^2 (g^2 - 2\lambda)$	0	$\frac{g^2}{2} (11c_{\theta_W}^2 - s_{\theta_W}^2)$
$\gamma_{\hat{c}_{\kappa\gamma}}$	0	$4(g^2 - g'^2)$	$\frac{11}{2}g^2 + \frac{g'^2}{2} + 6y_t^2 + 4\lambda$	0	$2g^2$
$\gamma_{\hat{c}_{gZ}}$	$\frac{g^2}{12c_{\theta_W}^2}$	0	$\frac{g^2}{6c_{\theta_W}^2}$	$\frac{17}{2}g^2 - \frac{g'^2}{6} + 6y_t^2$	0
$\gamma_{\hat{c}_{\lambda\gamma}}$	0	0	0	0	$\frac{53}{3}g^2$



Anomalous Dimensional Matrix



	\hat{c}_S	\hat{c}_T	\hat{c}_Y	\hat{c}_W	$\hat{c}_{\gamma\gamma}$
$\gamma_{\hat{c}_S}$	$\frac{1}{3}g'^2 + 6y_t^2$	$-\frac{g'^2}{2}$	$\frac{1}{8}g'^2 \left(147 - 106\frac{g'^2}{g^2}\right)$	$\frac{1}{8}(77g^2 + 58g'^2)$	$16e^2$
$\gamma_{\hat{c}_T}$	$-9g'^2 - 24t_{\theta_W}^2 \lambda$	$\frac{9}{2}g^2 + 12y_t^2 + 12\lambda$	$\frac{9}{2}g'^2 + 12t_{\theta_W}^2 (g'^2 + \lambda)$	$\frac{9}{2}g'^2$	0
$\gamma_{\hat{c}_Y}$	$-\frac{2}{3}g'^2$	0	$\frac{94}{3}g'^2$	0	0
$\gamma_{\hat{c}_W}$	0	0	$\frac{53}{12}g'^2 \left(1 - 3t_{\theta_W}^2\right)$	$\frac{331}{12}g^2 + \frac{29}{4}g'^2$	0
$\gamma_{\hat{c}_{\gamma\gamma}}$	0	0	0	0	$-\frac{9}{2}g^2 - \frac{3}{2}g'^2 + 6y_t^2 + 12\lambda$
$\gamma_{\hat{c}_H}$	$18g'^2 - t_{\theta_W}^2 (9g'^2 + 24\lambda)$	$-9g^2 + \frac{9}{2}g'^2 + 12\lambda$	$t_{\theta_W}^2 \left(-\frac{141}{4}g'^2 + 12\lambda\right)$	$\frac{63}{2}g^2 + \frac{51}{4}g'^2 + 72\lambda$	0
$\gamma_{\hat{c}_{\gamma Z}}$	0	0	0	0	0
$\gamma_{\hat{c}_{kZ}}$	0	0	0	0	$-16e^2$
$\gamma_{\hat{c}_{gZ}}$	$-\frac{g'^2}{6c_{\theta_W}^2}$	$\frac{g'^2}{12c_{\theta_W}^2}$	$\frac{g'^2}{8c_{\theta_W}^2} (106t_{\theta_W}^2 - 29)$	$-\frac{1}{8c_{\theta_W}^2} (79g^2 + 58g'^2)$	0
$\gamma_{\hat{c}_{\lambda\gamma}}$	0	0	0	0	0

	\hat{c}_H	$\hat{c}_{\gamma Z}$	$\hat{c}_{\kappa\gamma}$	\hat{c}_{gZ}	$\hat{c}_{\lambda\gamma}$
$\gamma_{\hat{c}_S}$	$-\frac{1}{6}g'^2$	$4(g^2 - g'^2)$	$-\frac{11}{2}g^2 - \frac{1}{6}g'^2 - 4\lambda$	$c_{\theta_W}^2 (9g^2 - \frac{1}{3}g'^2)$	$-2g^2$
$\gamma_{\hat{c}_T}$	$\frac{3}{2}g'^2$	0	$-9g'^2 - 24t_{\theta_W}^2 \lambda$	$24s_{\theta_W}^2 \lambda$	0
$\gamma_{\hat{c}_Y}$	0	0	$-\frac{2}{3}g'^2$	$\frac{2}{3}e^2$	0
$\gamma_{\hat{c}_W}$	0	0	0	$-\frac{2}{3}c_{\theta_W}^2 g^2$	0
$\gamma_{\hat{c}_{\gamma\gamma}}$	0	0	$\frac{3}{2}g^2 - 2\lambda$	0	$3g^2$
$\gamma_{\hat{c}_H}$	$-\frac{9}{2}g^2 - 3g'^2 + 12y_t^2 + 24\lambda$	0	$9g'^2(2 - t_{\theta_W}^2) - 24t_{\theta_W}^2 \lambda$	$9(g'^2 s_{\theta_W}^2 - g^2 c_{\theta_W}^2) - 24\lambda(6c_{\theta_W}^2 - s_{\theta_W}^2)$	0
$\gamma_{\hat{c}_{\gamma Z}}$	0	$-\frac{7}{2}g^2 - \frac{1}{2}g'^2 + 6y_t^2 + 12\lambda$	$c_{\theta_W}^2 (2g^2 - 2\lambda) - s_{\theta_W}^2 (g^2 - 2\lambda)$	0	$\frac{g^2}{2} (11c_{\theta_W}^2 - s_{\theta_W}^2)$
$\gamma_{\hat{c}_{\kappa\gamma}}$	0	$4(g^2 - g'^2)$	$\frac{11}{2}g^2 + \frac{g'^2}{2} + 6y_t^2 + 4\lambda$	0	$2g^2$
$\gamma_{\hat{c}_{gZ}}$	$\frac{g'^2}{12c_{\theta_W}^2}$	0	$\frac{g'^2}{6c_{\theta_W}^2}$	$\frac{17}{2}g^2 - \frac{g'^2}{6} + 6y_t^2$	0
$\gamma_{\hat{c}_{\lambda\gamma}}$	0	0	0	0	$\frac{53}{3}g^2$

+

Anomalous Dimensional Matrix

$$(\hat{c}_S, \hat{c}_T, \hat{c}_Y, \hat{c}_W, \hat{c}_{\gamma\gamma}, \hat{c}_{\gamma Z}, \hat{c}_{\kappa\gamma}, \hat{c}_{gz}, \hat{c}_{\lambda\gamma}, \hat{c}_H)^t (m_t) \simeq \quad (4)$$

$$\begin{array}{c}
 \mathbb{R} \\
 \left(\begin{array}{ccccccccccc}
 0.9 & 0.003 & -0.03 & -0.08 & -0.02 & -0.02 & -0.04 & 0.05 & -0.01 & 0.001 \\
 0.03 & 0.8 & -0.02 & -0.009 & 0 & 0 & -0.03 & 0.01 & 0 & -0.003 \\
 0.001 & 0 & 0.9 & 0 & 0 & 0 & -0.001 & 0.001 & 0 & 0 \\
 0 & 0 & -0.001 & 0.8 & 0 & 0 & 0 & -0.003 & 0 & 0 \\
 0 & 0 & 0 & 0 & 0.9 & 0 & 0.006 & 0 & 0.02 & 0 \\
 0 & 0 & 0 & 0 & 0 & 0.9 & 0.007 & 0 & 0.03 & 0 \\
 0 & 0 & 0 & 0 & -0.02 & -0.02 & 0.9 & 0 & -0.01 & 0 \\
 0.0004 & -0.0007 & -0.0004 & 0.1 & 0 & 0 & -0.0004 & 0.9 & 0 & -0.0007 \\
 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0.9 & 0 \\
 -0.02 & 0.03 & 0.01 & -0.4 & 0 & 0 & 0.02 & -0.3 & 0 & 0.8
 \end{array} \right) \begin{array}{l}
 \hat{c}_S(\Lambda) \\
 \hat{c}_T(\Lambda) \\
 \hat{c}_Y(\Lambda) \\
 \hat{c}_W(\Lambda) \\
 \hat{c}_{\gamma\gamma}(\Lambda) \\
 \hat{c}_{\gamma Z}(\Lambda) \\
 \hat{c}_{\kappa\gamma}(\Lambda) \\
 \hat{c}_{gz}(\Lambda) \\
 \hat{c}_{\lambda\gamma}(\Lambda) \\
 \hat{c}_H(\Lambda)
 \end{array}
 \end{array}$$

- We focus on the part of the matrix, where **weakly bound couplings contribute to strongly bound couplings.**

+ Numerical Results

Coupling	Direct Constraint	RG-induced Constraint
$\hat{c}_S(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_T(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_Y(m_t)$	$[-3, 3] \times 10^{-3}$ [22]	-
$\hat{c}_W(m_t)$	$[-2, 2] \times 10^{-3}$ [22]	-
$\hat{c}_{\gamma\gamma}(m_t)$	$[-1, 2] \times 10^{-3}$ [18]	-
$\hat{c}_{\gamma Z}(m_t)$	$[-0.6, 1] \times 10^{-2}$ [18]	$[-2, 6] \times 10^{-2}$
$\hat{c}_{\kappa\gamma}(m_t)$	$[-10, 7] \times 10^{-2}$ [27]	$[-5, 2] \times 10^{-2}$
$\hat{c}_{gZ}(m_t)$	$[-4, 2] \times 10^{-2}$ [27]	$[-3, 1] \times 10^{-2}$
$\hat{c}_{\lambda\gamma}(m_t)$	$[-6, 2] \times 10^{-2}$ [27]	$[-2, 8] \times 10^{-2}$
$\hat{c}_H(m_t)$	$[-6, 5] \times 10^{-1}$ [32]	$[-2, 0.5] \times 10^{-1}$

- We assume that there is **no tuning/correlation among different contributions** so that **each RG-induced** term in the RGE is **smaller than the bound**. This gives us new **RG-induced constraints**.
- We get bounds on some **TGC** and on **C_H** mainly from their RG-induced contribution to **{S, T, W, Y}** that are **stronger** than the direct bounds.

+ Numerical Results

Coupling	Direct Constraint	RG-induced Constraint
$\hat{c}_S(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_T(m_t)$	$[-1, 2] \times 10^{-3}$ [31]	-
$\hat{c}_Y(m_t)$	$[-3, 3] \times 10^{-3}$ [22]	-
$\hat{c}_W(m_t)$	$[-2, 2] \times 10^{-3}$ [22]	-
$\hat{c}_{\gamma\gamma}(m_t)$	$[-1, 2] \times 10^{-3}$ [18]	-
$\hat{c}_{\gamma Z}(m_t)$	$[-0.6, 1] \times 10^{-2}$ [18]	$[-2, 6] \times 10^{-2}$
$\hat{c}_{\kappa\gamma}(m_t)$	$[-10, 7] \times 10^{-2}$ [27]	$[-5, 2] \times 10^{-2}$
$\hat{c}_{gZ}(m_t)$	$[-4, 2] \times 10^{-2}$ [27]	$[-3, 1] \times 10^{-2}$
$\hat{c}_{\lambda\gamma}(m_t)$	$[-6, 2] \times 10^{-2}$ [27]	$[-2, 8] \times 10^{-2}$
$\hat{c}_H(m_t)$	$[-6, 5] \times 10^{-1}$ [32]	$[-2, 0.5] \times 10^{-1}$

- We assume that there is **no tuning** so that **each RG-induced** term in the RGE is **smaller than the bound**. This gives us new **RG-induced constraints**.
- We get bounds on some **TGC** and on **C_H** mainly from their RG-induced contribution to **{S, T, W, Y}** that are **stronger** than the direct bounds.

+ A Hierarchy of Constraints

$$\left. \delta g_{ff}^h \quad \delta g_{3h} \quad \delta g_{VV}^h \right\} \mathcal{O}(1)$$

$$\left. \begin{array}{l} \kappa_{Z\gamma} \\ \delta g_1^Z, \delta \kappa_\gamma, \lambda_\gamma \end{array} \right\} \text{percent level}$$

$$\left. \begin{array}{l} \kappa_{\gamma\gamma} \quad \kappa_{GG} \\ \delta g_{eR}^Z \quad \delta g_{eL}^Z \quad \delta g_{\nu L}^Z \\ \delta g_{uR}^Z \quad \delta g_{dR}^Z \quad \delta g_{uL}^Z \quad \delta g_{dL}^Z \end{array} \right\} \text{permille level}$$

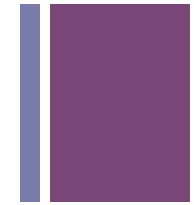
+ A Hierarchy of Constraints

$$\left. \begin{matrix} \delta g_{ff}^h & \delta g_{3h} \end{matrix} \right\} \mathcal{O}(1)$$

$$\left. \begin{matrix} \kappa_{Z\gamma} & \delta g_{VV}^h \\ \delta g_1^Z & \delta \kappa_\gamma & \lambda_\gamma \end{matrix} \right\} \begin{matrix} \text{percent level} \\ \text{Strong RG-induced} \\ \text{constraint from } S, T \end{matrix}$$

$$\left. \begin{matrix} \underline{\kappa_{\gamma\gamma}} & \underline{\kappa_{GG}} \\ \delta g_{eR}^Z & \delta g_{eL}^Z & \delta g_{\nu L}^Z \\ \delta g_{uR}^Z & \delta g_{dR}^Z & \delta g_{uL}^Z & \delta g_{dL}^Z \end{matrix} \right\} \text{permille level}$$

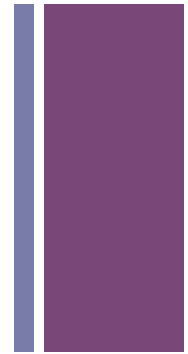
+ RG-effects measurable in future



Direct Measurement	Future Precision	$ \hat{c}_{\kappa\gamma} $	$ \hat{c}_{\gamma Z} $	$ \hat{c}_{\lambda\gamma} $	$ \hat{c}_H $
$\hat{c}_{\gamma\gamma}$	4×10^{-5} [36]	6×10^{-3}	-	2×10^{-3}	-
$\hat{c}_{\gamma Z}$	3×10^{-4} [36]	4×10^{-2}	-	1×10^{-2}	-
$\hat{c}_{\kappa\gamma}$	2×10^{-4} [37]	-	1×10^{-2}	1×10^{-2}	-
\hat{c}_{gZ}	2×10^{-4} [37]	0.4	-	-	0.25

- **Minimum value** of the couplings to which **direct measurements** of the observables in the first column would be **sensitive via the one loop RG-mixing in the long term**.
- A **deviation in one BSM primary parameter** would induce deviations in **other ones in a correlated way**.
- If these RG effects are not seen it would mean would mean that some **tuning** is needed, or it would indicate some **UV correlation** among Wilson coefficients pointing towards a particular structure of the new physics

+ Part III: Explicit Models



- We consider expectations for BSM primary effects in two models:

(1) Composite Models

[Giudice, Grojean, Pomarol and Rattazzi \(2007\)](#)

(2) Integrating out Higgses in SUSY Models

[Gupta, Montull, Riva \(2012\)](#)

+ Composite Models

- **Strongly Interacting Light Higgs (SILH)** Lagrangian:

$$\begin{aligned}
 \mathcal{L}_{\text{SILH}} = & \frac{c_H}{2f^2} \partial^\mu (H^\dagger H) \partial_\mu (H^\dagger H) + \frac{c_T}{2f^2} \left(H^\dagger \overleftrightarrow{D}^\mu H \right) \left(H^\dagger \overleftrightarrow{D}_\mu H \right) \\
 & - \frac{c_6 \lambda}{f^2} (H^\dagger H)^3 + \left(\frac{c_y y_f}{f^2} H^\dagger H \bar{f}_L H f_R + \text{h.c.} \right) \\
 & + \frac{ic_W g}{2m_\rho^2} \left(H^\dagger \sigma^i \overleftrightarrow{D}^\mu H \right) (D^\nu W_{\mu\nu})^i + \frac{ic_B g'}{2m_\rho^2} \left(H^\dagger \overleftrightarrow{D}^\mu H \right) (\partial^\nu B_{\mu\nu}) \\
 & + \frac{ic_{HW} g}{16\pi^2 f^2} (D^\mu H)^\dagger \sigma^i (D^\nu H) W_{\mu\nu}^i + \frac{ic_{HB} g'}{16\pi^2 f^2} (D^\mu H)^\dagger (D^\nu H) B_{\mu\nu} \\
 & + \frac{c_\gamma g'^2}{16\pi^2 f^2} \frac{g^2}{g_\rho^2} H^\dagger H B_{\mu\nu} B^{\mu\nu} + \frac{c_g g_S^2}{16\pi^2 f^2} \frac{y_t^2}{g_\rho^2} H^\dagger H G_{\mu\nu}^a G^{a\mu\nu}.
 \end{aligned}$$

(assumes **Higgs is a pseudo Nambu Goldstone Boson** of a strong sector)

Giudice, Grojean, Pomarol and Rattazzi
(2007)

+ Composite Models

$$\left. \begin{matrix} \delta g_{ff}^h & \delta g_{3h} \end{matrix} \right\}$$

$\mathcal{O}(1)$

$$\frac{g_*^2}{\Lambda^2}$$

$$\left. \begin{matrix} \kappa_{Z\gamma} & \delta g_{VV}^h \\ \delta g_1^Z & \delta \kappa_\gamma & \lambda_\gamma \end{matrix} \right\}$$

percent level

$$\frac{1}{\Lambda^2}$$

Strong RG-induced constraint from S, T

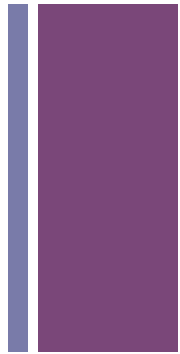
$$\left. \begin{matrix} \kappa_{\gamma\gamma} & \kappa_{GG} \\ \delta g_{eR}^Z & \delta g_{eL}^Z & \delta g_{\nu L}^Z \\ \delta g_{uR}^Z & \delta g_{dR}^Z & \delta g_{uL}^Z & \delta g_{dL}^Z \end{matrix} \right\}$$

$$\frac{g_*^2}{16\pi^2 \Lambda^2}$$

permille level

$$\frac{g_{SM}^2}{16\pi^2 \Lambda^2}$$

$$\hat{S} \quad \hat{T}$$



+ Composite Models

$$\left. \begin{array}{c} \delta g_{ff}^h \quad \delta g_{3h} \end{array} \right\}$$

$\mathcal{O}(1)$

$$\frac{g_*^2}{\Lambda^2}$$

$$\left. \begin{array}{c} \kappa_{Z\gamma} \quad \delta g_{VV}^h \\ \delta g_1^Z \quad \delta \kappa_\gamma \quad \lambda_\gamma \end{array} \right\}$$

percent level

$$\frac{1}{\Lambda^2}$$

Strong RG-induced constraint from S, T

$$\left. \begin{array}{c} \underline{\kappa_{\gamma\gamma}} \quad \underline{\kappa_{GG}} \\ \delta g_{eR}^Z \quad \delta g_{eL}^Z \quad \delta g_{\nu L}^Z \\ \delta g_{uR}^Z \quad \delta g_{dR}^Z \quad \delta g_{uL}^Z \quad \delta g_{dL}^Z \end{array} \right\}$$

\hat{S} \hat{T}

permille level

$$\frac{g_*^2}{16\pi^2 \Lambda^2}$$

$$\frac{g_{SM}^2}{16\pi^2 \Lambda^2}$$

+ Composite Models

$$\left. \begin{array}{c} \delta g_{ff}^h \quad \delta g_{3h} \end{array} \right\}$$

$\mathcal{O}(1)$

$$\frac{g_*^2}{\Lambda^2}$$

$$\left. \begin{array}{c} \kappa_{Z\gamma} \quad \delta g_{VV}^h \\ \delta g_1^Z \quad \delta \kappa_\gamma \quad \lambda_\gamma \end{array} \right\}$$

percent level

$$\frac{1}{\Lambda^2}$$

Strong RG-induced constraint from S, T

$$\left. \begin{array}{c} \kappa_{\gamma\gamma} \quad \kappa_{GG} \\ \delta g_{eR}^Z \quad \delta g_{eL}^Z \quad \delta g_{\nu L}^Z \\ \delta g_{uR}^Z \quad \delta g_{dR}^Z \quad \delta g_{uL}^Z \quad \delta g_{dL}^Z \end{array} \right\}$$

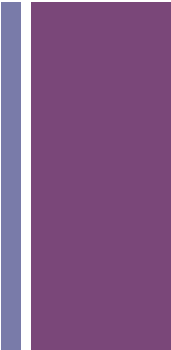
\hat{S} \hat{T}

permille level

$$\frac{g_*^2}{16\pi^2 \Lambda^2}$$

$$\frac{g_{SM}^2}{16\pi^2 \Lambda^2}$$

+ Composite Models



$$\boxed{\delta g_{ff}^h \quad \delta g_{3h}}$$

$\mathcal{O}(1)$

$$\boxed{\frac{g_*^2}{\Lambda^2}}$$

Left-right symmetry

$$\left. \begin{array}{l} \cancel{\kappa_{Z\gamma}} \quad \delta g_{VV}^h \\ \boxed{\delta g_1^Z} \quad \delta \kappa_\gamma \quad \lambda_\gamma \end{array} \right\}$$

percent level

$$\boxed{\frac{1}{\Lambda^2}}$$

Strong RG-induced constraint from S, T

$$\left. \begin{array}{l} \kappa_{\gamma\gamma} \quad \kappa_{GG} \\ \delta g_{eR}^Z \quad \delta g_{eL}^Z \quad \delta g_{\nu L}^Z \\ \delta g_{uR}^Z \quad \delta g_{dR}^Z \quad \delta g_{uL}^Z \quad \delta g_{dL}^Z \end{array} \right\}$$

permille level

$$\frac{g_*^2}{16\pi^2 \Lambda^2}$$

$$\frac{g_{SM}^2}{16\pi^2 \Lambda^2}$$

$$\boxed{\hat{S}}$$

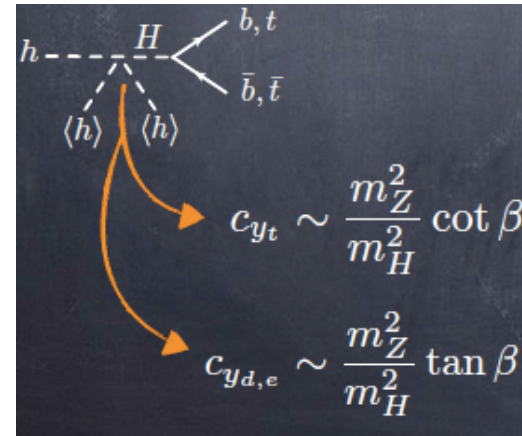
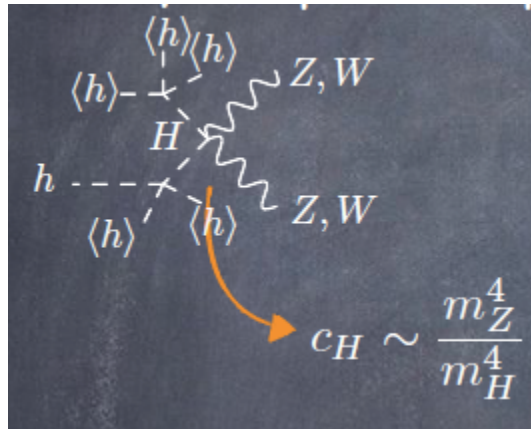
$$\cancel{\boxed{\hat{T}}}$$

Custodial symmetry



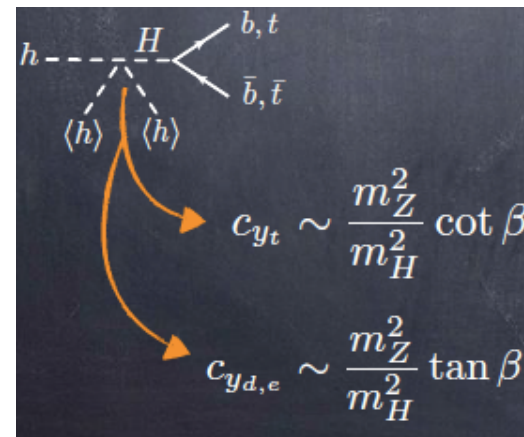
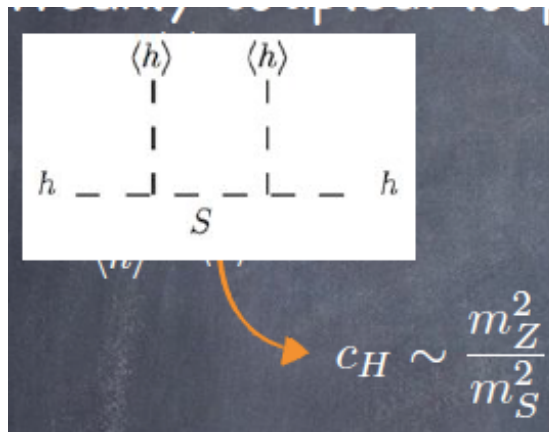
Integrating out heavy Higgses in SUSY

- Supersymmetric models (2HDMS)



- NMSSM

(\mathcal{O}_6 also generated)



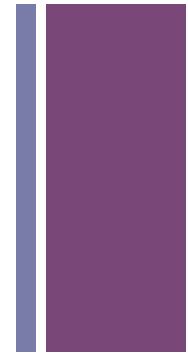
+ Integrating out heavy Higgses in SUSY

2HDM: $\left. \begin{matrix} \delta g_{ff}^h & \delta g_{3h} \end{matrix} \right\} \mathcal{O}(1)$

$\left. \begin{matrix} \kappa_{Z\gamma} & \delta g_{VV}^h \\ \delta g_1^Z & \delta \kappa_\gamma & \lambda_\gamma \end{matrix} \right\}$
percent level
Strong RG-induced
constraint from S, T

$\left. \begin{matrix} \underline{\kappa_{\gamma\gamma}} & \underline{\kappa_{GG}} \\ \delta g_{eR}^Z & \delta g_{eL}^Z & \delta g_{\nu L}^Z \\ \delta g_{uR}^Z & \delta g_{dR}^Z & \delta g_{uL}^Z & \delta g_{dL}^Z \end{matrix} \right\}$
permille level

+ Integrating out heavy Higgses in SUSY



NMSSM: $\left. \begin{matrix} \delta g_{ff}^h & \delta g_{3h} \end{matrix} \right\}$

$\mathcal{O}(1)$

$\left. \begin{matrix} \kappa_{Z\gamma} & \delta g_{VV}^h \\ \delta g_1^Z & \delta \kappa_\gamma & \lambda_\gamma \end{matrix} \right\}$

percent level
Strong RG-induced
constraint from S, T

$\left. \begin{matrix} \underline{\kappa_{\gamma\gamma}} & \underline{\kappa_{GG}} \\ \delta g_{eR}^Z & \delta g_{eL}^Z & \delta g_{\nu L}^Z \\ \delta g_{uR}^Z & \delta g_{dR}^Z & \delta g_{uL}^Z & \delta g_{dL}^Z \end{matrix} \right\}$

permille level

+ Understanding SUSY Higgs coupling deviations

- Write potential in terms of h and H , where:

$$\begin{aligned}h_1^0 &= \cos \beta h + \sin \beta H \\h_2^0 &= \sin \beta h - \cos \beta H\end{aligned}$$

gets full VEV

- H and h almost mass eigenstates if $\delta_i v^2 / m_H^2 \ll 1$
- h has exactly SM couplings as it gives mass to all the particles.

quartics



SUSY modifications to **raise the Higgs mass** would necessarily **change Higgs couplings** in a **correlated** way!

+ Understanding SUSY Higgs coupling deviations

- As quartics are turned on the lightest mass eigenstate is no longer h and the misalignment causes deviations from SM couplings:

$$\Delta V(H_1, H_2) = +\delta_\lambda h^4 + \delta h^3 H + \delta_2 h^2 H^2 + \delta_3 h H^3 + \delta_4 H^4$$

raises Higgs mass

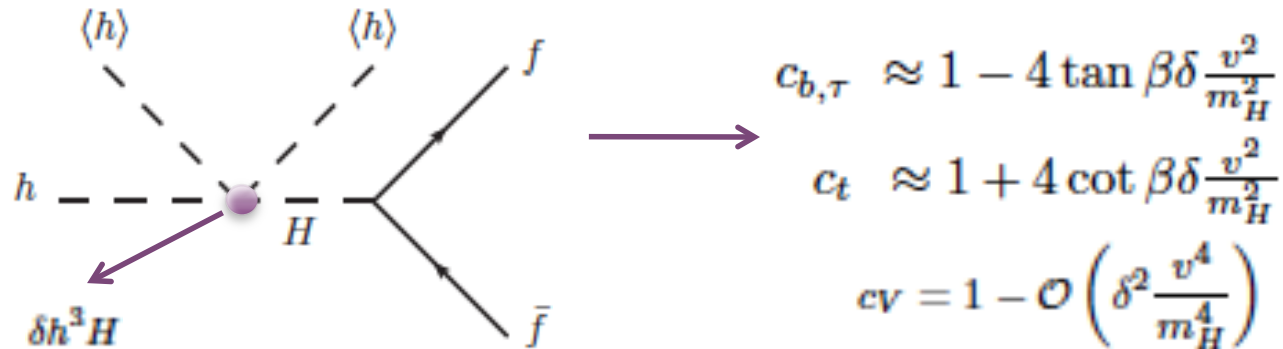
changes Higgs couplings wrt SM by inducing mixing term: hH and causing misalignment of $\{h, H\}$ with the mass eigenstate basis

$$\Delta m_h^2 = 16\delta_\lambda v^2$$

$$\Delta m_h^2 = m_h^{\text{obs}2} - m_Z^2 (\cos 2\beta)^2 \gtrsim (86 \text{ GeV})^2$$

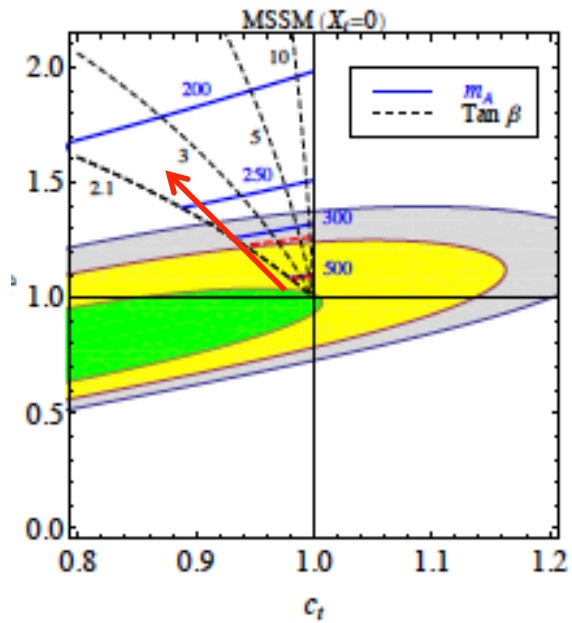
+ Understanding SUSY Higgs coupling deviations

- Integrate out H to obtain:

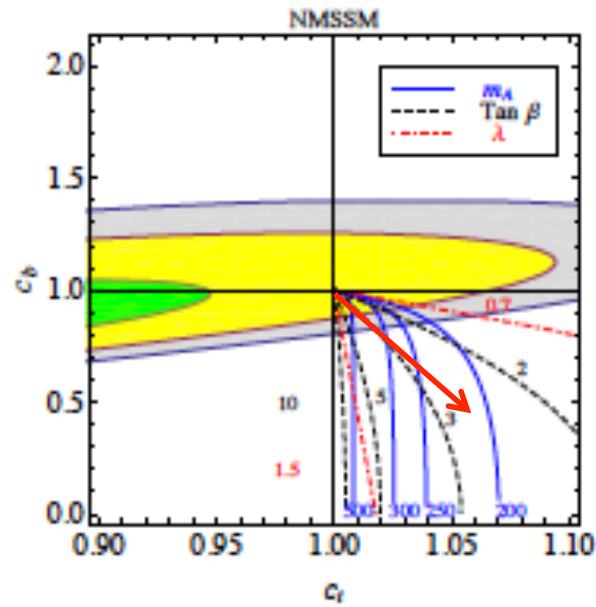


	ΔV	δ_λ	δ
MSSM	$\frac{g^2+g'^2}{8} (H_1^0 ^2 - H_2^0 ^2)^2$	$\frac{m_{\tilde{t}}^2}{16v^2} (c_\beta^2 - s_\beta^2)^2$	$\frac{m_{\tilde{t}}^2}{2v^2} s_\beta c_\beta (c_\beta^2 - s_\beta^2)$
Stops (no mixing)	$\frac{\lambda}{2} H_2 ^4 = \frac{3v_t^4}{8\pi^2} \log[m_{\tilde{t}_1} m_{\tilde{t}_2}/M_t^2] H_2 ^4$	$s_\beta^4 \frac{\lambda_2}{8}$	$-4s_\beta^3 c_\beta \frac{\lambda_2}{8}$
D-term extension	$\kappa (H_1^0 ^2 - H_2^0 ^2)^2$	$\frac{m_{\tilde{t}}^2}{16v^2} (c_\beta^2 - s_\beta^2)^2$	$\frac{m_{\tilde{t}}^2}{2v^2} s_\beta c_\beta (c_\beta^2 - s_\beta^2)$
NMSSM	$\lambda^2 H_1^0 H_2^0 ^2$	$\frac{\lambda^2}{16} \sin^2 2\beta$	$-\frac{\lambda^2}{8} \sin 4\beta$

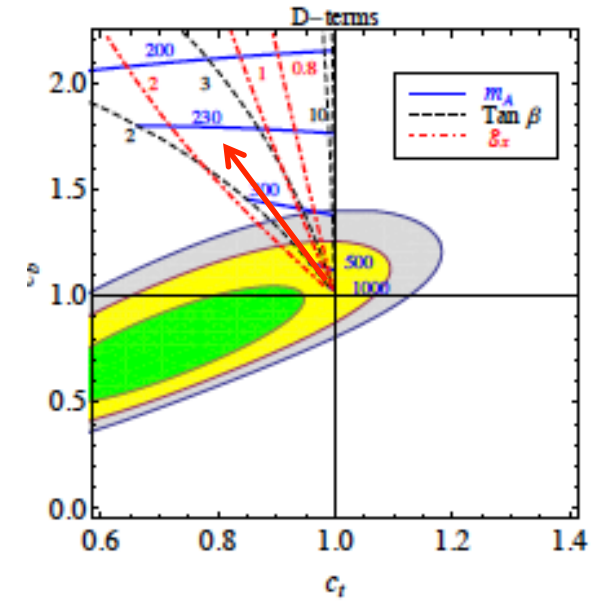
$$m_Z^2/v^2 \rightarrow 4\kappa.$$



MSSM
($\delta < 0$)



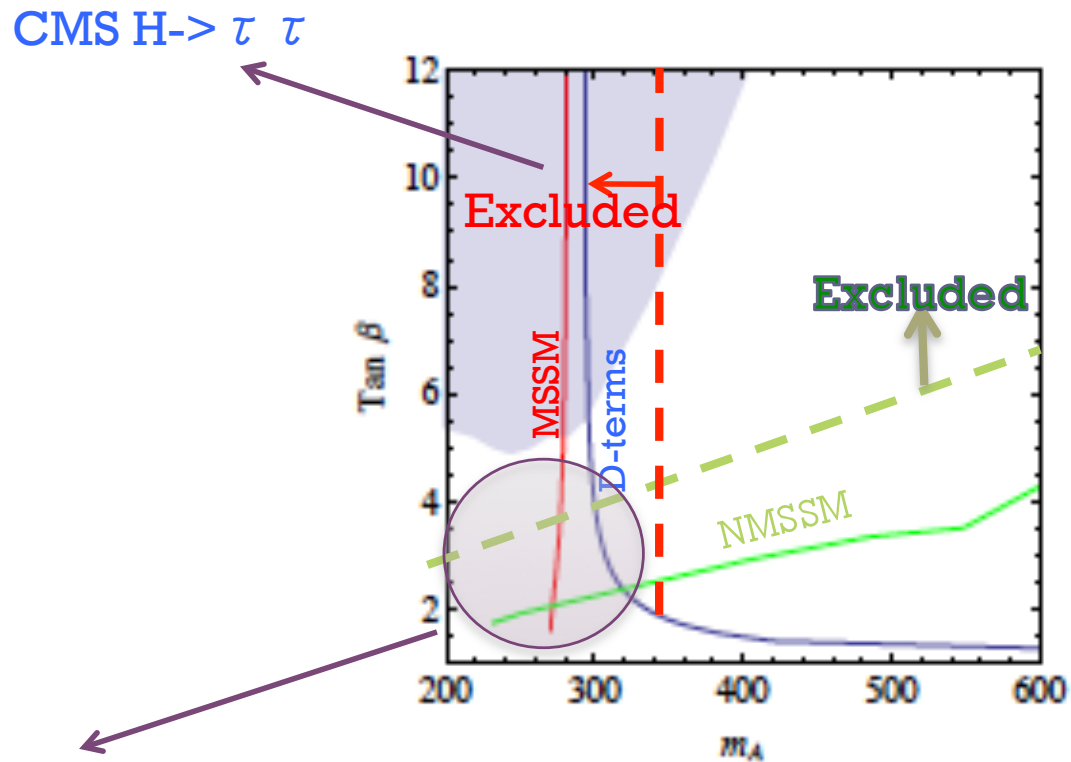
NMSSM
($\delta > 0$)



D-terms
($\delta < 0$)

- All qualitative features of the above plots can be understood using our expansion. Quantitatively it is approximate but works well if $m_A > 350$ GeV.

+ Exclusions



Higgs coupling data
more competitive than
direct searches in
low $\tan \beta$ region

Dashed: Barbieri et al
(2012)
with latest data
Solid lines: our bounds

+ Consistency check: shift invariance

- The anomalous dimension of shift invariant combinations of couplings is a function of shift invariant combination of couplings

$$\gamma_{C_i} = f(C_j)$$

	c_H	c_T	c_T	c_B	c_W
γ_{c_H}	$28\lambda + \frac{3}{2}g^2 - 4\gamma_H$	$\frac{3}{2}(2g^2 + g'^2) - 4\lambda$	$8\lambda - 6g^2 - \frac{3}{2}g'^2$	$\frac{3}{4}g'^2(g'^2 + 4g^2)$	$\frac{3}{4}g^2(3g^2 + 4g'^2) - 6\lambda g^2$
γ_{c_T}	$4\lambda - 3g^2$	$\frac{3}{2}(5g^2 + g'^2) + 20\lambda - 4\gamma_H$	$-4\lambda + 3g^2 - 6g'^2$	$\frac{3}{4}g'^2(2g'^2 - g^2) + 6\lambda g'^2$	$\frac{3}{2}g^2(6g^2 - g'^2) + 30\lambda g^2$
γ_{c_T}	$\frac{3}{2}g'^2$	$-\frac{3}{2}g'^2$	$12\lambda + \frac{27}{2}g^2 + 3g'^2 - 4\gamma_H$	$-\frac{3}{4}g'^2g^2 - 6\lambda g'^2$	$-\frac{3}{4}g'^2g^2$
γ_{c_B}	$-\frac{1}{3}$	$\frac{1}{3}$	$-\frac{5}{3}$	$\frac{g'^2}{6} + 6g^2$	$\frac{g^2}{3}$
γ_{c_W}	$-\frac{1}{3}$	$\frac{1}{3}$	$-\frac{1}{3}$	$\frac{g'^2}{6}$	$\frac{11}{2}g^2 + 6g^2$
$\gamma_{c_{2B}}$	0	0	0	$-\frac{2}{3}g'^2$	0
$\gamma_{c_{2W}}$	0	0	0	0	$-\frac{2}{3}g^2$

- Thus the contribution of c_B and c_T to γ_{c_T} must be connected in such a way that they can combine to form C_T

$$\gamma_{C_T} = \gamma_{c_T} - \frac{1}{4}g'^2(2\gamma_{c_B}) \dots \equiv \#c_T + \#c_B + \#c_{2B} = \frac{1}{6}(72\lambda + 5g'^2 + 27g^2) C_T$$

This will connect for eg. the contribution $c_B \rightarrow C_{T,B,2B}$ $c_T \rightarrow C_{T,B,2B}$ which have been computed completely independently.

+ Consistency check: shift invariance

- The anomalous dimension of shift invariant combinations of couplings is a function of shift invariant combination of couplings

$$\gamma_{C_i} = f(C_j)$$

	c_H	c_T	c_T	c_B	c_W
γ_{c_H}	$28\lambda + \frac{3}{2}g^2 - 4\gamma_H$	$\frac{3}{2}(2g^2 + g'^2) - 4\lambda$	$6\lambda - 6g^2 - \frac{3}{2}g'^2$	$\frac{3}{2}g'^2(g'^2 + 4g^2)$	$\frac{3}{4}g^2(3g^2 + 4g'^2) - 6\lambda g^2$
γ_{c_T}	$4\lambda - 3g^2$	$\frac{3}{2}(5g^2 + g'^2) + 20\lambda - 4\gamma_H$	$-4\lambda + 3g^2 - 6g'^2$	$\frac{3}{2}g'^2(2g'^2 - g^2) + 6\lambda g'^2$	$\frac{3}{2}g^2(6g^2 - g'^2) + 30\lambda g^2$
γ_{c_B}	$\frac{3}{2}g'^2$	$-\frac{3}{2}g'^2$	$12\lambda + \frac{27}{2}g^2 + 3g'^2 - 4\gamma_H$	$-\frac{3}{4}g'^2 g^2 - 6\lambda g'^2$	$-\frac{3}{4}g'^2 g^2$
γ_{c_W}	$-\frac{1}{3}$	$\frac{1}{3}$	$-\frac{5}{3}$	$\frac{g'^2}{6} + 6v_t^2$	$\frac{g^2}{3}$
$\gamma_{c_{2B}}$	$-\frac{1}{3}$	$\frac{1}{3}$	$-\frac{1}{3}$	$\frac{g'^2}{6}$	$\frac{11}{2}g^2 + 6v_t^2$
$\gamma_{c_{2W}}$	0	0	0	$-\frac{2}{3}g'^2$	0
$\gamma_{c_{2W}}$	0	0	0	0	$-\frac{2}{3}g^2$

- Thus the contribution of c_B and c_T to γ_{c_T} must be connected in such a way that they can combine to form C_T

$$\gamma_{C_T} = \gamma_{c_T} - \frac{1}{4}g'^2(2\gamma_{c_B}) \dots \equiv \#c_T + \#c_B + \#c_{2B} = \frac{1}{6}(72\lambda + 5g'^2 + 27g^2) C_T$$

This will connect for eg. the contribu $c_B \rightarrow C_{T,B,2B}$ $c_T \rightarrow C_{T,B,2B}$ which have been computed completely independently.

+ Higgs Primary directions

- For the effect from $e(\hat{h})$, $s_{\theta_W}(\hat{h})$ we write SM Lagrangian with non canonical kinetic terms (we can get back to SM by $A_\mu \rightarrow A_\mu - s_{\theta_W}^2(v)Z_\mu$)

$$\begin{aligned} \mathcal{L}_{EW} = & -\frac{1}{4e^2(\hat{h})} \left(A_{\mu\nu} + s_{\theta_W}^2(\hat{h}) Z_{\mu\nu} \right)^2 - \frac{c_{\theta_W}^2(\hat{h})}{4g^2(\hat{h})} Z_{\mu\nu}^2 \\ & - \frac{1}{2g^2(\hat{h})} W_{\mu\nu}^+ W^{-\mu\nu} + \frac{\hat{h}^2}{2} \left[W_\mu^+ W^{-\mu} + \frac{1}{2} Z^\mu Z_\mu \right] \\ & + A_\mu J_{em}^\mu + Z_\mu J_3^\mu + W_\mu^+ J_+^\mu + W_\mu^- J_-^\mu, \end{aligned}$$

$$e(\hat{h}) = e \left(1 + \kappa_{\gamma\gamma} \frac{\hat{h}^2}{v^2} \right), \quad s_{\theta_W}(\hat{h}) = s_{\theta_W}$$

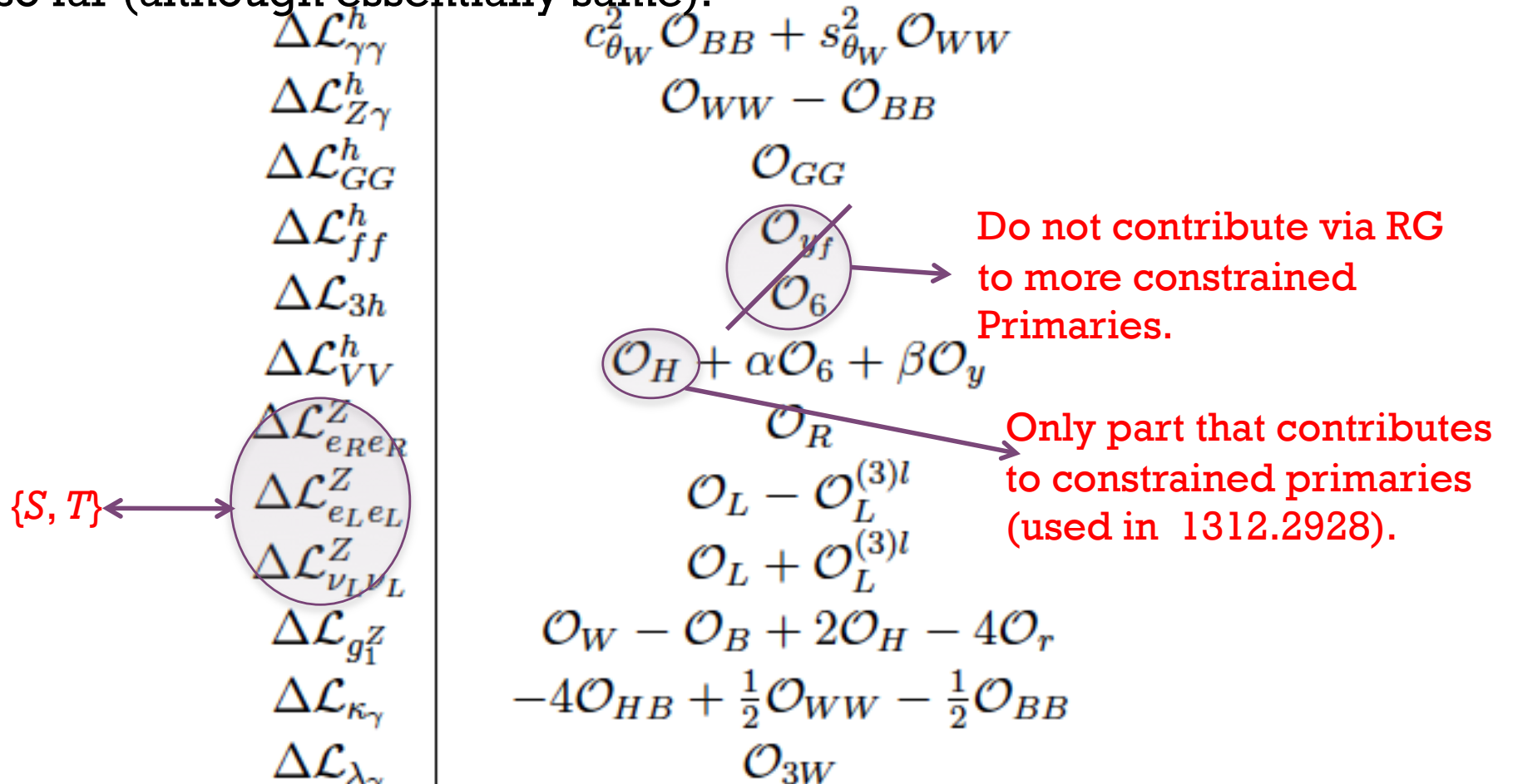
$$\begin{aligned} \Delta \mathcal{L}_{\gamma\gamma}^h = & \kappa_{\gamma\gamma} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[A_{\mu\nu} A^{\mu\nu} \right. \\ & \left. + Z_{\mu\nu} Z^{\mu\nu} + 2W_{\mu\nu}^+ W^{-\mu\nu} \right], \end{aligned}$$

$$e(\hat{h}) = e, \quad s_{\theta_W}^2(\hat{h}) = s_{\theta_W}^2 \left(1 - \kappa_{Z\gamma} \frac{\hat{h}^2}{v^2} \right)$$

$$\begin{aligned} \Delta \mathcal{L}_{Z\gamma}^h = & \kappa_{Z\gamma} \left(\frac{h}{v} + \frac{h^2}{2v^2} \right) \left[t_{\theta_W} A_{\mu\nu} Z^{\mu\nu} \right. \\ & \left. + \frac{c_{2\theta_W}}{2c_{\theta_W}^2} Z_{\mu\nu} Z^{\mu\nu} + W_{\mu\nu}^+ W^{-\mu\nu} \right]. \end{aligned}$$

+ Difference in Parametrization

- Parametrization in this part slightly different from one used so far (although essentially same):



Now we also considered W and Y that have been traded for four-fermions so far.