

### **PRISM and PRIME**



## Measuring $\mu^{-} \rightarrow e^{-}$ conversion at the 10<sup>-18</sup> level and beyond

Roger Barlow Manchester University



#### Contents



- •The Challenge: 10<sup>-18</sup>
- •PRISM: Phase Rotated Intense Slow Muon source
  - Phase Rotation
  - FFAGs
  - Magnets
  - RF
- •PRIME: PRIsm Mu-E conversion
- •The PRISM task force
- •Outlook

#### µ-e conversion



Rate down by factor ~200 from  $\mu \rightarrow e\gamma$  BUT probes different BSM physics AND has very clean signature of monochromatic105 MeV electron, beyond decay spectrum.

To go beyond COMET we will need: 1) More muons For statistics 2) No pion contamination Pions can decay to electrons with energies up to 140 MeV 3) Extra beam extinction As discussed 4) No muons above 76 MeV/c As these can decay in flight to 105 MeV electrons 5) Very small muon energy/momentum spread

This can achieve an experiment essentially free of any background



Want very narrow energy spread so muons all stop in same foil thickness (just 1 foil). Reduces electron energy straggling, and improves spectrometer accuracy as z known.

#### <u>But</u>

Muons are produced with wide spread in energy/momentum. Just selecting a narrow momentum bite is too inefficient

**Solution** (preserving Liouville's theorem)

Generate muons in very short pulses. Trade certainty in time for certainty in energy





#### **Phase Rotation**



All muons in a bunch start together

Muons with higher energy travel faster and arrive earlier at the RF

Arrange RF phase so that they are decelerated

Muons with lower energy travel slower and arrive later at the RF

Arrange RF phase so that they are accelerated

Reduce momentum spread from 20% to 2% in a few turns (~1.5 µs)



TAU2010 Manchester







Proton Accelerator 2-8 GeV ~1 MW with bunch compression

Solid target, pions captured by 4-10T solenoid and allowed to decay

Storage ring:Comprising•Magnets•RF (~40 gaps as acceleration must be fast)•Injection and extraction systemsHas to operate at ~1000 HzAcceptance 3.8 π cm rad horizontally, 0.57 π cm rad verticallyMomentum 68 MeV/c (below 75 MeV/c)Momentum spread 20% → 2%

#### Stopping foil

Electron spectrometer.



#### The system





Manchester

Roger Barlow: PRISM and PRIME

Slide 7/23



## Why an FFAG?



#### What is an FFAG?

Cyclic accelerator in which particle moves from low-field to high-field region



#### What are the advantages?

Acceleration not limited by ramping magnets. So can be rapid.

Very large acceptance

Simplicity, robustness...

Changes in dipole field provide quadrupole field for strong focussing

TAU2010 Manchester



### FFAG prototypes





#### Nested proton FFAGs at KURRI



A prototype electron nsFFAG has been built in Daresbury and is now being tested.

EMMA stands for...

Roger Barlow: PRISM and PRIME



### The Magnets



Wide aperture, complicated field shape

Prototype built, and magnetic fields measured at KEK to be as predicted/required



Roger Barlow: PRISM and PRIME



## The RF



Require high gradient – 200 kV/m – to give 2-3 MV/turn

At low frequency – 4 MHz

In a magnetic field

With triangular shape: add harmonics (Asymmetric sawtooth is even better)

Low duty factor (0.1%) so cooling not a problem

Prototype using (expensive) magnetic alloy with very large cores (~1.7 m X 1.0 m) successfully constructed

Experiments with beam have been performed to simulate the bunch rotation.

Development of a new (cheaper!) material, FT3L, is underway TAU2010 Manchester







Prototype tested with alpha particles



Roger Barlow: PRISM and PRIME

Slide 12/23



## **PRIME: the detector**

PRISM

Magnetic spectrometer – curved solenoid

Resolution ~ 150 keV

Rate is very high. Off energy particles must be lost. They never make it to the tracking chambers.

With 10<sup>9</sup> muons/pulse, about 10 electrons/pulse will reach the spectrometer

Calorimeter is for triggering/calibration/checking







#### Peformance



 $10^9$  muons per pulse  $10^3$  pulses per second  $10^{12}$  muons per second ~20x more than COMET

10<sup>7</sup> seconds per year0.4 detector (PRIME) efficiency

~ 5x more than COMET due to better beam quality

Expect ~0.06 background events

Branching ratios of 10<sup>-18</sup> and below will be observable with a few years' running

## The PRISM Task Force



Aim: to address the technological challenges in realising an FFAG based muon-to-electron conversion experiment, and strengthen R&D for muon accelerators in the context of the Neutrino Factory and future muon physics experiments, especially

- the physics of muon to electron conversion,
- proton source,
- pion capture,
- muon beam transport,
- injection & extraction for PRISM
- FFAG ring design including the search for a new improved version,

- FFAG hardware R&D for RF and injection extraction kicker and septum magnets.

**TAU2010** Manchester Roger Barlow: PRISM a

J. Pasternak, Imperial College + RAL, UK L. J. Jenner, A. Kurup, Imperial College, UK/FNAL, USA Y. Üchida, M. Aslaninejad, **Imperial College, UK** B. Muratori, S. L. Smith, D. Kelliher, S. Machida, C. Prior, ASTeC, (RAL and DL) UK K. M. Hock, **Cockcroft Institute/Liverpool, UK R. J. Barlow**, **Cockcroft Institute/ManchesterUK** C. Ohmori, **KEK/JAEA**, Ibaraki-ken, Japan H. Witte, T. Yokoi, JAI, Oxford University, UK J-B. Lagrange, Y. Mori, Kyoto University, KURRI, Japan Y. Kuno, A. Sato, Osaka University, Osaka, Japan M. Lancaster,

UCL, London, UK



## PRISM Task Force Design Strategy



Option 1: Adopt current design and work out injection/extraction, and hardware

Option 2: Find a new design

They should be evaluated in parallel and confronted with the number of muons delivered to target/cost

#### Requirements for a new design:

- •High transverse acceptance (at least 38h/5.7v [Pi mm]).
- High momentum acceptance (at least ± 20% or more).
- Small orbit excursion.
- Compact ring size
- Relax or at least conserve the level of technical difficulties. for hardware (kickers, RF) with respect to the current design.

TAU2010 Manchester

Roger Barlow: PRISM and PRIME

# Task force activity: PRISM kicker studies



- •length 1.6 m
- •B 0.02 T
- •Aperture: 0.95 m x 0.5 m
- Flat top 40 /210 ns (injection / extraction)
  rise time 80 ns (for extraction)
- •fall time ~200 ns (for injection)
- •W<sub>mag</sub>=186 J
- •L = 3 uH (preliminary)
- •I<sub>max</sub>=16 kA

TAU2010 Manchester



H. Witte, M. Aslaninejad, J. Pasternak

Roger Barlow: PRISM and PRIME

Slide 17/23





#### New solutions



Scaling FFAG

Options:,

- periodic with extended cell (for example 5 magnets per cell),
- superperiodic (proposed by S. Machida)
- advanced (proposed by Y. Mori and collaborators, see J-P.Lagrange's lattice)

Non-Scaling FFAG

Main motivation for Non-Scaling design is a possibility to obtain "infinite" acceptance

- •Problems to be addressed:
- currently no insertion scheme,
- very difficult injection
- TOF varies with amplitude



## Alternative system





Roger Barlow: PRISM and PRIME

## Non-scaling FFAG design





- Simple linear and rectangular magnet design.
- Tunes vary with momentum, but large acceptance is expacted for ~6 turns.
- TOF variation with amplitude affects the final energy spread, but seems to be acceptable.
- Injection/extraction and matching seem difficult, but not imposible.
- More cells (20) could reduce the orbit excursion (magnet size), but assuming constant drift length in the symmetric ring, more RF would be required!
- This difficulty may be overcome by the design with insertions for the injection/extraction.



## PRISM and beyond?





#### Synergies include:

- high power proton driver
- short proton bunch length
- pion production and capture
- muon beam transport
- need for high acceptances
- requirements for kickers and RF (for proton driver)
- techiniques in machine design and beam dynamics studies

• etc ...

#### •Muon Collider

TAU2010 Manchester

Roger Barlow: PRISM and PRIME

Slide 22/23



## Outlook



#### PRISM task force

Aim: a CDR in 2011 showing how to measure  $\mu \rightarrow e$  conversion at BR<10<sup>-18</sup>

PRISM could be built at JPARC, or as part of Project X, or...

Lots of work to do – help welcome - contact Jarloslaw Pasternak (J.Pasternak@imperial.ac.uk) to get involved